

A Model-Driven Approach to Estimate Atmospheric Visibility with Ordinary Cameras

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Abstract

Atmospheric visibility is an important input for road and air transportation safety, as well as a good proxy to estimate the air quality. A model-driven approach is presented to monitor the meteorological visibility distance through use of ordinary outdoor cameras. Unlike in previous data-driven approaches, a physics-based model is proposed which describes the mapping function between the contrast in the image and the atmospheric visibility. The model is non-linear, which allows encompassing a large spectrum of applications. The model assumes a continuous distribution of objects with respect to the distance in the scene and is estimated by a novel process. It is more robust to illumination variations by selecting the Lambertian surfaces in the scene. To evaluate the relevance of the approach, a publicly available database is used. When the model is fitted to short range data, the proposed method is shown to be effective and to improve on existing methods. In particular, it allows envisioning an easier deployment of these camera-based techniques on multiple observation sites.

Keywords: camera, visibility, observation, road safety, aviation safety, air quality

1. Introduction

In the presence of fog, haze or air pollution, atmospheric visibility is reduced. This constitutes a common and vexing transportation problem for different public authorities in multiple countries throughout the world.

First, low visibility is obviously a problem of traffic safety. Indeed, the behavior of drivers in fog is often inappropriate (e.g. reduced headways, altered reaction times) but the reasons for these dangerous behaviors are not fully understood (Kang et al., 2008; Caro et al., 2009). A recommendation in order to improve the safety in such situations was to use two rear fog lamps in the vehicles, as far apart as possible (Cavallo et al., 2001). It was also suggested that lowering the height of these lamps could lead to reduced headway estimation (Buchner et al., 2006). Various countermeasures have been tested on the roadside to reduce the impact of critically reduced visibility (Shepard, 1996), among which automated warning systems employing road-side weather stations and visibility meters to provide automated detection (Mac Carley, 2005).

In addition to the road safety problem, reduced visibility is cause of delays and disruption in air, sea and ground transportation for passengers and freight. On highways, massive pile-ups create non-recurrent congestions which sometimes force the operator to momentarily close the road. Fog-related road closures are not an uncommon subject for news headlines. Another example is the Heathrow airport which was blocked for three days during Christmas 2006. Such events have important economic impacts (Pejovic et al., 2009). According to Perry and Symons (1991), in 1974 fog was estimated to have cost over roughly £120 millions (at 2010 prices) on the roads of Great Britain. This figure includes the cost of medical treatment, damage to vehicles and property, as well as the adminis-

trative costs of police, services and insurance, but they do not include the cost of delays to people not directly involved in the accident.

Impaired visibility is also a symptom of environmental problems because it is evidence of air pollution (Hyslop, 2009); in addition, it has been shown that impaired visibility in urban environment and mortality are correlated (Thach et al., 2010). According to Thach et al. (2010), visibility provides a useful proxy for the assessment of environmental health risks from ambient air pollutants, and a valid approach for the assessment of the public health impacts of air pollution where pollutant monitoring data are scarce.

The ability to accurately monitor visibility helps solving these problems. Critical safety at important transportation facilities such as airports are generally instrumented for monitoring visibility with devices that are expensive and hence, scarce. Cost is precisely the reason why highway meteorological stations are seldom equipped with visibility meters. In this context, using existing and ubiquitous highway cameras is of great interest, as these are low cost sensors already deployed for other purposes such as traffic monitoring (Jacobs et al., 2009). Furthermore, introducing new functionalities into road-side cameras would make them multipurpose and thus more cost-effective, easing their deployment along the roads.

Attempts at estimating the visibility using outdoor cameras or webcams are reported in the literature. However, the visibility range differs from one application to another, so that there is no general approach to tackle the problem by camera. For road safety applications, the range 0-400 m is usually considered. For meteorological observation and airport safety, the range 0-1000 m is usually considered. Visual range is also used for monitoring pollution in urban areas. In this case, higher visual

65 ranges, typically 1-5 km, are usually considered. In the follow-122
 66 ing, we address the whole spectrum of visual ranges, typically123
 67 0-10 km. 124

68 Two families of methods are proposed in the literature. The125
 69 first one estimates the maximum distance at which a selected126
 70 target can be detected. The methods differ depending on the127
 71 nature of the target and how to estimate the distance. For in-128
 72 telligent vehicles as well as for visual monitoring of highway129
 73 traffic, a black target at the horizon is chosen and a flat road130
 74 is assumed. Bush and Debes (1998) use a wavelet transform131
 75 to detect the highest edge in the image with a contrast above132
 76 5%. Based on a highway meteorology standard, Hautière et al.133
 77 (2008) proposed a reference-free roadside camera-based sensor134
 78 which not only estimates the visibility range but also estimates135
 79 whether the visibility reduction is caused by fog. For meteoro-136
 80 logical observations, regions of interest whose distance can be137
 81 obtained on standard geographic maps are selected manually138
 82 (Bäumer et al., 2008). An accurate geometric calibration of the139
 83 camera with respect to the scene is necessary to calibrate and140
 84 operate continuously these methods, which may be understood141
 85 as direct approaches.

86 A second family of methods correlates the contrast in the
 87 scene with the visual range estimated by reference additional142
 88 sensors (Hallowell et al., 2007). No accurate geometric cali-143
 89 bration is needed. Conversely, a learning phase is needed to
 90 estimate the function which maps the contrast in the scene to
 91 the visual range. The method proposed in this paper belongs to
 92 this second family. Usually, a gradient based on the Sobel filter
 93 or a high-pass filter in the frequency domain are used to com-
 94 pute the contrast (Liaw et al., 2010; Hagiwara et al., 2007; Xie
 95 et al., 2008). Luo et al. (2005) have shown that the visual range
 96 obtained with these two approaches are highly correlated. Liaw
 97 et al. (2010) proposed to use a homomorphic filter or a Haar
 98 function in addition to the high-pass filter in order to reduce the
 99 effects of non-uniform illumination. Once the contrast is com-144
 100 puted, a linear regression is performed to estimate the mapping145
 101 function (Hallowell et al., 2007; Xie et al., 2008; Liaw et al.,146
 102 2010). Babari et al. (2010) propose a method which is robust147
 103 to illumination variations in the scene by taking into account148
 104 the physical properties of objects in the scene. Unlike previous149
 105 methods, a non-linear data regression is performed which al-
 106 lows covering a wider spectrum of applications. Due to the step
 107 of data regression, these methods can be seen as data-driven
 108 approaches. Nevertheless, the major problem of data-driven
 109 methods is the need of a learning phase, which makes this kind
 110 of method difficult to deploy massively. Indeed, one must wait
 111 for an episode with impaired visibility, so as to collect learning
 112 data and compute the fitting parameters.

113 The direct approaches are very sensitive to the geometric cali-
 114 bration of the camera but no learning phase is necessary to use
 115 them. The data-driven approaches do not use any accurate geo-
 116 metric calibration. However, they need episodes with impaired
 117 visibility before they are operational. We believe that new tech-
 118 niques can be developed which need neither accurate geometric
 119 calibration nor learning phase. In this aim, one must model how
 120 the contrast in the scene is altered by the presence of reduced
 121 visibility conditions, so as to build an *a priori* mapping func-

tion between the contrast and the atmospheric visibility distance
 in the scene. This constitutes a model-driven approach. Haut-
 122 tière et al. (2010) propose such a probabilistic model-driven ap-
 123 proach which allows computing a physics-based mapping func-
 124 tion. In particular, the model takes into account an *a priori*
 distribution of contrasts in the scene. However, a uniform dis-
 125 tribution of targets is assumed which limits the applicability of
 126 the method on any scene. In this article, the method proposed
 127 in (Hautière et al., 2010) is generalized by adding new targets
 128 distributions, as well as a method to estimate the actual distri-
 129 bution of objects in the scene. A great attention is paid to the
 130 data fitting process, which greatly influences the final results.
 131 To assess the relevance of the approach, the different methods
 132 are compared using the MATILDA database (Hautière et al.,
 133 2010).

This article is organized as follows. In section 2,
 Koschmieder’s model of fog visual effects is recalled. In sec-
 134 tion 3, the model-driven approach is presented, whose experi-
 135 mental evaluation is carried out in section 4. Finally, the results
 136 are discussed and perspectives for future work are given.

2. Vision through the Atmosphere

2.1. Koschmieder’s Theory

The attenuation of luminance through the atmosphere was
 studied by Koschmieder (Middleton, 1952), who derived an
 equation relating the extinction coefficient of the atmosphere β ,
 which is the sum of the scattering coefficient and of the absorp-
 137 tion coefficient, the apparent luminance L of an object located at
 138 distance d , and the luminance L_0 measured close to this object:

$$L = L_0 e^{-\beta d} + L_\infty (1 - e^{-\beta d}) \quad (1)$$

(1) indicates that the luminance of the object seen through fog
 is attenuated by $e^{-\beta d}$ (Beer-Lambert law); it also reveals a lu-
 139 minance reinforcement of the form $L_\infty (1 - e^{-\beta d})$ resulting from
 140 daylight scattered by the slab of fog between the object and the
 141 observer, the so-called airlight. L_∞ is the atmospheric lumi-
 142 nance.

On the basis of this equation, Duntley developed a contrast
 attenuation law (Middleton, 1952), stating that a nearby object
 exhibiting contrast C_0 with the fog in the background will be
 143 perceived at distance d with the following contrast:

$$C = \left[\frac{L_0 - L_\infty}{L_\infty} \right] e^{-\beta d} = C_0 e^{-\beta d} \quad (2)$$

This expression serves to base the definition of a standard di-
 144 mension called meteorological visibility distance V , i.e. the
 145 greatest distance at which a black object ($C_0 = -1$) of a suit-
 146 able dimension can be seen on the horizon, with the threshold
 147 contrast set at 5% (CIE, 1987). It is thus a standard parame-
 148 ter that characterizes the opacity of a fog layer. This definition
 149 yields the following expression:

$$V \approx \frac{3}{\beta} \quad (3)$$

150 More recently, Koschmieder's model has received a lot of at-
 151 tention in the computer vision community, e.g. (Narasimhan
 152 and Nayar, 2003; Hautière et al., 2007; Tan, 2008; He et al.,
 153 2009; Tarel and Hautière, 2009). Indeed, it is possible based on
 154 this model to infer the 3D structure of a scene in presence of
 155 fog, or to dehaze/defog images by reversing the model. How-
 156 ever, it is worth mentioning that in these works a relative esti-
 157 mation of the meteorological visibility is enough to restore the
 158 visibility. In this paper, Koschmieder's model is used to esti-
 159 mate the actual meteorological visibility distance, which makes
 160 the problem quite different (an absolute estimation of the visi-
 161 bility is needed).

162 2.2. Contrast of Lambertian Targets

Assuming a linear response function of the camera, the in-173
 tensity I of a distant point located at distance d in an outdoor
 scene is given by Koschmieder's model (1):

$$174 \quad I = Re^{-\beta d} + A_\infty(1 - e^{-\beta d}) \quad (4) \quad 175$$

where R is the intrinsic intensity of the pixel, i.e. the inten-177
 sity corresponding to the intrinsic luminance value of the cor-178
 responding scene point and A_∞ is the background sky intensity.179
 Two points located at roughly the same distance $d_1 \approx d_2 = d$
 with different intensities $I_1 \neq I_2$ form a distant target whose nor-
 malized contrast is given by:

$$180 \quad C = \frac{I_2 - I_1}{A_\infty} = \left[\frac{R_2 - R_1}{A_\infty} \right] e^{-\beta d} = C_0 e^{-\beta d} \quad (5) \quad 181$$

In this equation, the contrast C of a target located at distance183
 d depends on $V = \frac{3}{\beta}$ and on its intrinsic contrast C_0 . If we184
 now assume that the surface of the target is Lambertian, the185
 luminance L at each point i of the target is given by:186

$$187 \quad L = \rho_i \frac{E}{\pi} \quad (6) \quad 188$$

where E denotes the global illumination and ρ_i denotes the
 albedo at i . Moreover, it is a classical assumption to set $L_\infty = \frac{E}{\pi}$
 so that (5) finally becomes:

$$189 \quad C = (\rho_2 - \rho_1) e^{-\beta d} \approx (\rho_2 - \rho_1) e^{-\frac{3d}{V}} = \Delta\rho e^{-\frac{3d}{V}} \quad (7)$$

163 Consequently, the contrast of a distant Lambertian target only
 164 depends on its physical properties and on its distance to the sen-
 165 sor and on the meteorological visibility distance, and no longer
 166 on the illumination. These surfaces are robust to strong illumina-
 167 tion variations in the computation of the contrast in the scene.

168 3. The Model-Driven Approach

169 3.1. Principle

Let us consider an outdoor scene where targets are distributed190
 continuously at increasing distances from the camera. Let us
 denote ϕ the probability density function of observing a con-
 trast C in the scene:

$$191 \quad \mathbb{P}(C < X \leq C + dC) = \phi(C) dC \quad (8)$$

The expectation of the contrast m in the image is expressed as:

$$m = \mathbb{E}[C] = \int_0^1 C \phi(C) dC \quad (9)$$

Based on (7), C is a random variable which depends of the two
 random variables d and $\Delta\rho$. These two variables are assumed
 to be independent, which allows expressing (9) as:

$$m = \mathbb{E}[\Delta\rho] \mathbb{E}\left[e^{-\frac{3d}{V}}\right] = \overline{\Delta\rho} \int_0^{+\infty} \psi(d) e^{-\frac{3d}{V}} dd \quad (10)$$

where $\overline{\Delta\rho}$ denotes the mean albedo difference between the ob-
 jects in the scene and ψ denotes the p.d.f. of there being an
 object at the distance d in the scene. To compute m , a realistic
 expression for the density of objects ψ in the scene is needed.

174 3.2. Expectation of the Contrast

Choosing a suitable target distribution ψ allows us comput-
 ing the expectation of the contrast (10) with respect to the me-
 teorological visibility distance. In (Hautière et al., 2010), (10)
 was solved assuming a uniform distribution of targets between
 0 and d_{\max} , which leads to the following solution:

$$m_u = \frac{V \overline{\Delta\rho}}{3d_{\max}} \left[1 - \exp\left(-\frac{3d_{\max}}{V}\right) \right] \quad (11)$$

This assumption may be useful when the scene is not known
a priori but may limit the applicability of the method on any
 scene. The problem has received little consideration in the liter-
 ature. Torralba and Oliva (2002) proposed some *a priori* depth
 distributions in natural or man-made scenes which are Gaussian
 distributions. To circumvent this problem, a solution is to esti-
 mate the actual distribution and to solve m for this distribution.
 Let us first examine if mathematical solutions exist for classical
 statistical distributions.

Assuming a Gaussian distribution of parameters μ and σ , the
 density of targets is given by:

$$\psi_G(d) = \frac{1}{\sigma \sqrt{2\pi}} \exp\left[-\frac{1}{2} \left(\frac{d-\mu}{\sigma}\right)^2\right] \quad (12)$$

(10) then has an analytical solution m_g , which is given by:

$$m_g(V) = \frac{\overline{\Delta\rho}}{2} \exp\left(\frac{9\sigma^2}{2V^2} - \frac{3\mu}{V}\right) \operatorname{erfc}\left[\frac{1}{\sigma \sqrt{2}} \left(\frac{3\sigma^2}{V} - \mu\right)\right] \quad (13)$$

where erfc denotes the complementary error function:

$$\operatorname{erfc}(z) = \frac{2}{\sqrt{\pi}} \int_z^\infty \exp(-\zeta^2) d\zeta \quad (14)$$

In the same way, assuming a Rayleigh distribution of param-
 eter σ :

$$\psi_R(d) = \frac{d}{\sigma^2} \exp\left(-\frac{d^2}{2\sigma^2}\right) \quad (15)$$

$$m_R(V) = 1 - \frac{3\sigma \overline{\Delta\rho}}{V} \exp\left(\frac{9\sigma^2}{2V^2}\right) \sqrt{\frac{\pi}{2}} \operatorname{erfc}\left(\frac{3\sigma}{V \sqrt{2}}\right) \quad (16)$$

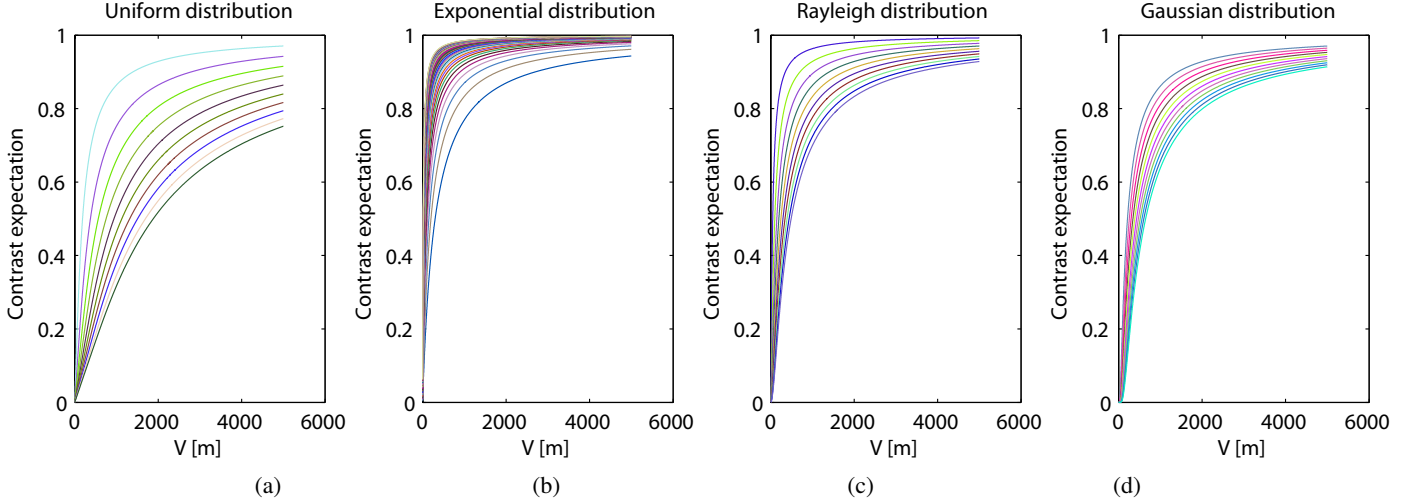


Figure 1: Plots of the different contrast expectation models assuming (a) a uniform distribution ($d_{max} \in [100; 1000]$); (b) an exponential distribution of targets density ($v \in [0.01; 0.1]$); (c) a Rayleigh distribution of targets density ($\sigma \in [10; 100]$); (d) a Gaussian distribution of targets density ($\sigma = 10$ and $\mu \in [50; 150]$).

Finally, assuming an exponential distribution of parameter v :

$$\psi_e(d) = v \exp(-vd) \quad (17)$$

$$m_e(V) = \frac{v \Delta \rho}{v + \frac{3}{V}} \quad (18)$$

Other types of distributions can be tested, such as the log-normal distribution. However, mathematical solutions are not easy to find and then to handle, apart from the uniform and exponential distributions.

3.3. Model Inversion and Error Estimation

The different models are all increasing functions of V and share the same limits towards 0 and ∞ , see Eqs. (11,13,16,18):

$$\lim_{V \rightarrow 0} m = 0 \quad \lim_{V \rightarrow \infty} m = 1 \quad (19)$$

which are obvious physical bounds that data-driven approaches do not always respect. The models of contrast expectation presented in the previous section are plotted as functions of the meteorological visibility distance V in Fig. 1. As one can see, these models have roughly the same shape.

In (Hautière et al., 2010), the solution for the uniform case was found to be invertible:

$$V(m_u) = \frac{3m_u d_{max}}{1 + m_u W\left(\frac{e^{-1/m_u}}{m_u}\right)} \quad (20)$$

where W denotes the Lambert function, which is a transcendental function defined by solutions of the equation $W(x)e^{W(x)} = x$ (Corless et al., 1996). Given the complexity of the equation, it is somehow difficult to compute the partial derivatives of the model and express error bounds of the model. In the case of the Gaussian and Rayleigh distributions, it is also possible to find analytical solutions to invert the models, but these ones are not detailed here for the sake of readability of the article.

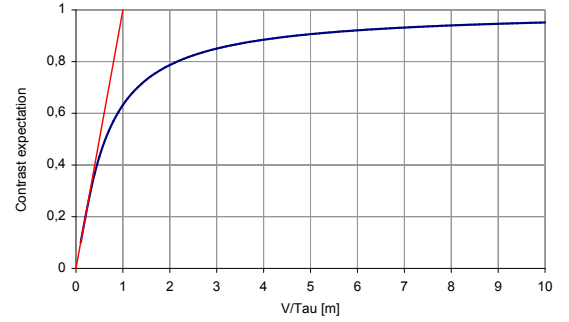


Figure 2: Analogy between the charge/discharge of a capacitor and the shape of the contrast expectation (blue curve) with respect to the meteorological visibility. The red curve denotes the tangent at the origin.

Fortunately in the case of an exponential distribution, a simpler solution is available:

$$V(m_e, v) = \frac{3m_e}{v(1-m_e)} \quad (21)$$

With this model, the partial derivatives of V with respect to m and v (22) can be obtained and an upper bound of the error of the model (23) is derived:

$$dV = \frac{\partial V(m_e, v)}{\partial m_e} dm_e + \frac{\partial V(m_e, v)}{\partial v} dv \quad (22)$$

$$\Delta V \leq \frac{3\Delta m_e}{v(1-m_e)^2} + \frac{3m_e \Delta v}{v^2(1-m_e)} \quad (23)$$

At this stage, we can make a comparison with the charging/discharging of a capacitor. Assuming a uniform distribution, (11) can be expressed as following:

$$m_u = \frac{\Delta \rho}{\tau} V \left[1 - \exp\left(-\frac{\tau}{V}\right) \right] \quad (24)$$

where $\tau = 3d_{max}$. When $V = \tau$, we have $m_u = 1 - e^{-1} \approx 0.63$. This is the same constant as the one used to characterize the

207 charging speed of a capacitor. Fig. 2 shows the curve obtained
 208 when plotting (24) with respect to the ratio $\frac{V}{\tau}$.

209 In the general case, the capacitance of the system is deter-
 210 mined by the distribution of distances in the scene, the texture
 211 of the objects in the scene and the quality (MTF, resolution) of
 212 the camera along with the response of the image processing fil-
 213 ter (e.g. the Sobel filter). The smaller the capacitance of the
 214 system, the faster the curves go to 1. We thus define an indica-
 215 tor τ of the system quality which is the meteorological visibility
 216 distance at which 0.63 of the "capacitance" is reached.

217 3.4. Estimation of the Distribution of Targets

The direct computation of m and V strongly depends on the
 distribution ψ . Thus, an important task is to guess which dis-
 tribution is best suited for a given scene. Following the method
 proposed by Narasimhan and Nayar (2003), the scene struc-
 ture can be approximated from two weather conditions 1 and 2
 thanks to Koschmieder's law (1):

$$(\beta_2 - \beta_1)d = -\log \left[\frac{A_{\infty_2} - I_2}{A_{\infty_1} - I_1} \right] - \log \frac{A_{\infty_1}}{A_{\infty_2}} \quad (25)$$

218 Using this method, it is possible to roughly estimate a depth
 219 for each pixel of the scene image. Starting from Narasimhan
 220 and Nayar (2003), we used landmarks of known depth and we
 221 adjusted the sky intensities A_{∞_1} and A_{∞_2} so as to improve the
 222 accuracy of the global map.

223 Second, due to the noise of the camera sensor, a simple com-
 224 putation of the depth distribution is useless. Soft-voting is often
 225 used to obtain reliable data from multiple uncertain data sources
 226 (Latif-Shabgahi et al., 2004). In computer vision and pattern
 227 recognition, this process is often used to deduce a global infor-
 228 mation from local information, e.g. the Hough transform (Duda
 229 and Hart, 1972), the fast radial symmetry transform (Loy and
 230 Zelinsky, 2003) or the v-disparity transform (Labayrade et al.,
 231 2002). In a similar way, the distribution of Lambertian targets
 232 can be estimated using a Parzen's like approach (Parzen, 1962).²⁴⁸

233 In this aim, a cumulative histogram of depth $h(d)$ is com-
 234 puted for $d \in [0, d_{max}]$ which takes into account a bandwidth²⁴⁹
 235 parameter. This one is related to the confidence u_i on the esti-²⁵⁰
 236 mation of the distance associated to each pixel. For each pixel,²⁵¹
 237 a normal distribution $\mathcal{N}(d|d_i, u_i)$ is cumulated in the histogram²⁵²
 238 with center d_i and standard deviation u_i . In addition to the²⁵³
 239 standard Parzen's approach, we also use a weighting parame-²⁵⁴
 240 ter w_i which accounts for the contribution of each data to the
 241 histogram. This histogram of depth is then expressed by:²⁵⁵

$$h(d) = \sum_{i=1}^P w_i \mathcal{N}(d|d_i, u_i) \quad (26)$$

where P denotes the total number of pixels. The confidence u is²⁵⁹
 obtained by computing the sensitivity of (25) to its parameters:²⁶⁰

$$u \propto \sum \left[\frac{\partial d}{\partial (A_{\infty_{1,2}}, I_{1,2})} d(A_{\infty_{1,2}}, I_{1,2}) \right]^2 \quad (27)$$

Assuming $dA_{\infty_1} \approx dA_{\infty_2} \approx dI_1 \approx dI_2 = dI$, (27) becomes:

$$u \propto \frac{f_1 + f_2}{(\beta_2 - \beta_1)^2} dI^2 \quad (28)$$

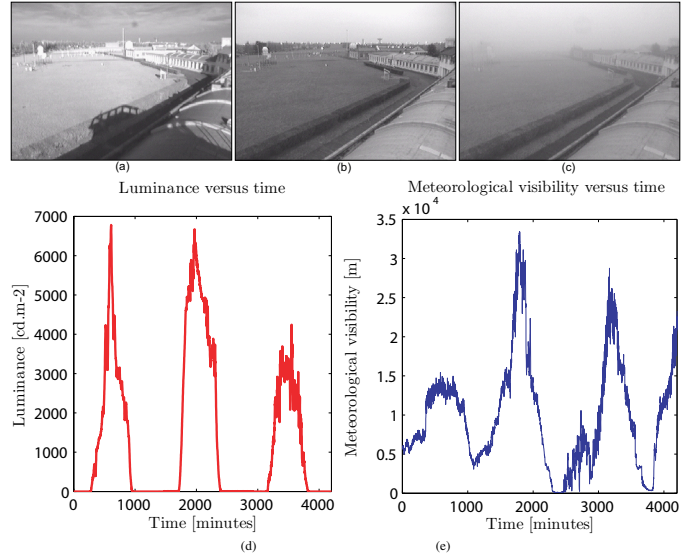


Figure 3: Samples of data collected in winter 2008-2009: (a) images with strong illumination conditions and presence of shadows; (b) cloudy conditions; (c) foggy weather situation; (d) meteorological visibility distance data and (e) background luminance data collected in the field test during two days.

where f_i is given by:

$$f_{i=1,2} = \frac{1}{A_{\infty_i}^2} + 2 \left[\frac{1}{(A_{\infty_i} - I_i)^2} + \frac{1}{A_{\infty_i}(A_{\infty_i} - I_i)} \right] \quad (29)$$

In section 4.4, we apply this method to actual data issued from a test site. In particular, we chose the relevant weight w_i . Having estimated h , the relevant distribution model can be determined empirically or using classical statistical tests. If the distribution has different modes, a probability mixture model can also be used to fit h .

4. Experimental Validation

In this section, an experimental evaluation of the proposed approach for visibility estimation is carried out. In this aim, the publicly available MATILDA database is used. First, the methodology is presented. Second, a method to estimate whether a surface is Lambertian or not is recalled. Third, results are presented and discussed.

4.1. Experimental Data

The observation test field is equipped with a reference transmissometer (Degreane Horizon TI8510). It serves to calibrate different scatterometers (Degreane Horizon DF320) used to monitor the meteorological visibility distance on the French territory, one of which provided our data. They are coupled with a background luminance sensor (Degreane Horizon LU320) which monitors the illumination received by the sensor. A camera grabs images of the field test every ten minutes. This camera is an 8-bit CCD camera (640 × 480 definition, mounting height 8.3 m, pitch angle 9.8°, focal length $f_l = 4$ mm and pixel size $t_{pix} = 9 \mu m$). It is thus a low cost camera which is representative of common video surveillance cameras.



Figure 4: Map of Lambertian surfaces on the field test: The redder the pixel is, the higher the probability that the surface is Lambertian.

Two fog events were collected at the end of February 2009. The fog occurred early in the morning and lasted a few hours after sunrise. During the same days, sunny weather periods also occurred. Fig. 3 shows sample images of (a) sunny weather, (b) cloudy weather and (c) foggy weather. The meteorological visibility distances and luminances are plotted in Figs. 3(d)&(e) versus time on a three day period. As one can see, the meteorological visibility distance ranges from 100 m to 35,000 m and the luminance ranges from 0 to 6,000 $\text{cd}\cdot\text{m}^{-2}$. This database made of 150 images grabbed every ten minutes is available on the LCPC's website <http://www.lcpc.fr/en/produits/matilda/> for research purpose.

4.2. Location of Lambertian surfaces

To estimate m and thus V , the normalized gradient is computed on the Lambertian surfaces of the scene as proposed in section 3. Locating the Lambertian surfaces in the images is thus needed. Following the method proposed in Babari et al. (2010), the Pearson coefficient, denoted $P_{i,j}^L$, is computed between the intensity of pixels in image series where the position of the sun changes and the value of the background luminance estimated by the luminancemeter. The closer $P_{i,j}^L$ is to 1, the stronger the probability that the pixel belongs to a Lambertian surface. This technique provides an efficient way to locate the Lambertian surfaces in the scene. For the MATILDA database, the density map of Lambertian surfaces is shown in Fig.4. The redder the pixel, the higher the probability that the surface is Lambertian.

4.3. Contrast Estimator

Having located the Lambertian surfaces, the gradients in the scene are estimated by means of the module of the Sobel filter. For each pixel, the gradient $\nabla_{i,j}$ is normalized by the intensity of the background A_∞ . Since the camera is equipped with an auto gain control, the background intensity A_∞ is most of the time equal to $2^8 - 1$, so that this step can be skipped. Each gradient is then weighted by $P_{i,j}^L$, the probability that a pixel (i, j) belongs to a Lambertian surface. Consequently, only relevant areas of the image are used for the visibility estimation, and

the scene need not be totally Lambertian. Finally, the estimated contrast in the scene \tilde{m} is given by:

$$\tilde{m} = \frac{1}{N} \sum_{i,j} \Delta\rho_{i,j} \exp\left(-\frac{3d_{i,j}}{V}\right) P_{i,j}^L = \frac{1}{N} \sum_{i,j} \frac{\nabla_{i,j}}{A_\infty} \quad (30)$$

where $\Delta\rho_{i,j}$ is the intrinsic contrast of a pixel (7) and N denotes the number of pixels of the image.

4.4. Selection of the Relevant Distribution

In section 3.4, we have proposed a methodology to estimate the distribution ψ in a scene. In this section, we apply this method to the test site of the MATILDA database. Having the contrast estimator (see previous paragraph), we are now able to derive a relevant weight w_i . Based on (30), the contribution of a data to the histogram is its weighted gradient $\nabla_{i,j} P_{i,j}^L$ computed in good weather conditions, which leads to choose it as weight w_i , see (26). The confidence u_i on the depth of each pixel is given by (28) and it is controlled by the value of dI which is set empirically. The estimated distribution is shown in Fig. 5 using the green plot ($dI = 0.1$), the purple plot ($dI = 0.25$) and the black curve ($dI = 1$). The exponential distribution fits the data quite well and is chosen to model the data of the histogram because it is the most easily revertible and is plotted in red. Based on this curve, we estimate $d_{max} \approx 325$ m. We can thus expect a capacitance τ of approximately 1000 m.

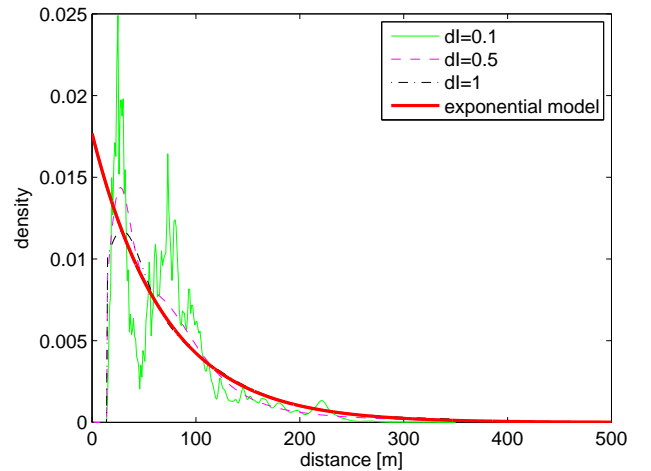


Figure 5: Histogram of weighted contrasts versus depth. The estimated distribution is shown using the green plot ($dI = 0.1$), the purple plot ($dI = 0.5$) and the black curve ($dI = 1$). The fitted exponential distribution is plotted in red.

4.5. Results

As in Babari et al. (2010), \tilde{m} is computed for the collection of 150 images of the MATILDA database using (30). The exponential distribution model (18) has been fitted to all the data using a robust non-linear least squares fitting technique ($R^2 = 0.91$), namely the Levenberg-Marquardt algorithm. We have also fitted upper and lower bound curves which comprise 99% of the data points. The different curves are plotted in Fig. 6(a).

Application	Highway fog	Meteorological fog	Haze	Air quality	
Range [m]	0-400	0-1000	0-5000	0-10000	0-15000
Number of data	13	19	45	70	150
Weighted logarithmic model (Babari et al., 2010)	10.4%	22.5%	23.4%	29.9%	41.9%
Uniform distribution (Hautière et al., 2010)	12.6%	18.1%	29.7%	∞	∞
Exponential distribution	10.0%	16.2%	29%	60%	373%
Exponential distribution + enhanced fitting	9.7%	11.2%	33%	50%	63.5%

Table 1: Mean relative errors of meteorological visibility distance estimation with respect to the envisioned applications.

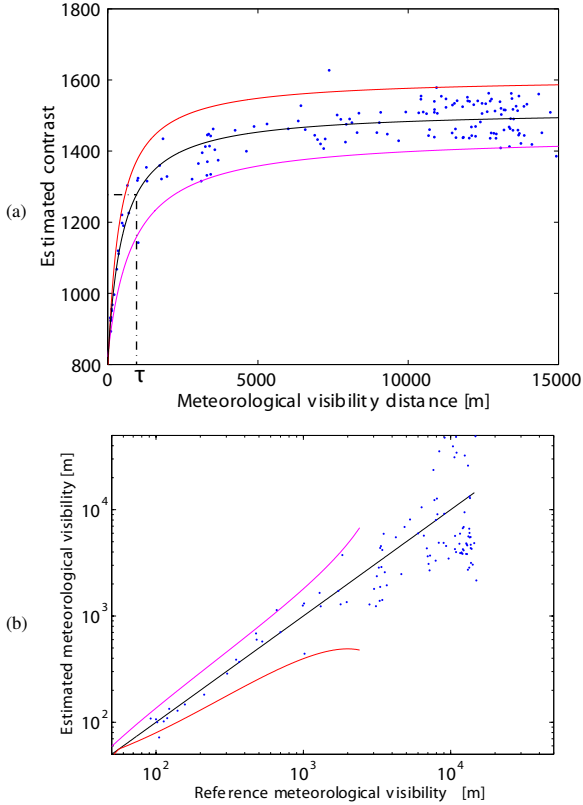


Figure 6: Model fitting: (a) Data fitting with the exponential distribution model in black. The upper bound is plotted in blue and the lower bound in magenta. (b) Plot of estimated visibility distances versus reference visibility distances.

We estimated a capacitance of the scene $\tau \approx 950\text{m} \approx 3d_{max}$ as expected. We invert the fitted model using (21) and estimate the meteorological visibility distance based on the contrast expectation m . Finally, we plot the estimated meteorological visibility distance versus the reference meteorological visibility distance in Fig. 6(b). From the same experimental data, Babari et al. (2010) fit an empirical logarithmic model, whereas Hautière et al. (2010) fit the contrast expectation of a uniform distribution (11).

The mean relative errors are compared in Tab. 1. Since the applications are very different depending on the range of meteorological visibility distances, the relative error for various applications are computed: road safety, meteorological observation and air quality.

Compared to data-driven approaches, one can see that the error remains low with model-driven approaches for critical safety applications, increases for higher visibility ranges, and becomes huge for visibility distances above 7 km. On the test site, using the actual target distribution, i.e. the exponential model, improves the previous results obtained with the uniform distribution (Hautière et al., 2010) and covers a large spectrum of applications with a limited error. Due to the unbalanced data fitting process, the error is slightly higher for low visibility ranges (<1.000 m) using the data-driven approach (Babari et al., 2010) despite the weighting introduced by the authors. It remains limited for higher visibility ranges (>5000 m).

In the previous results, all the data have been used to fit the models. This is the principle underlying the data-driven approach. Conversely, this approach should not be followed for the model-driven approach, since the model may not be valid for the whole ranges of visibility. According to section 4.4, we are sure that the model is valid in the range $0 - \tau$, i.e. 0-1000 m in our case. A new data fitting process is deduced. First, the exponential distribution model (18) has been fitted to the data in the range 0-1000 m using a robust non-linear least squares fitting technique, namely the Levenberg-Marquardt algorithm. The confidence in the fitting is higher ($R^2 = 0.97$). The fitted curve is shown in Fig. 7. Second, the model is extrapolated on the range $\tau - 15000$ m. The mean relative error is then computed between the adjusted model and the ground truth data. The results are given in the last line of Tab. 1. Since the model has been fitted to short visibility data, the results are improved at short ranges. At higher ranges, the errors are reduced as well, which illustrates the benefits of performing a data fitting process only on reliable data.

Finally, according to metrology practices in the field of visibility observations, a measurement device is considered as correct if the error is smaller than 20% in 90% of the cases. The 10% worst cases are thus excluded from the error computation. In this way, we are able to obtain a correct estimate of the meteorological visibility up to 3320 m.

5. Discussion

The data-driven approach requires visibility data for its calibration and implementation. Both model-driven approaches need only to determine the type of targets distribution in the scene. The distributions used in this article, namely uniform and exponential, are parameterized by a single parameter d_{max}

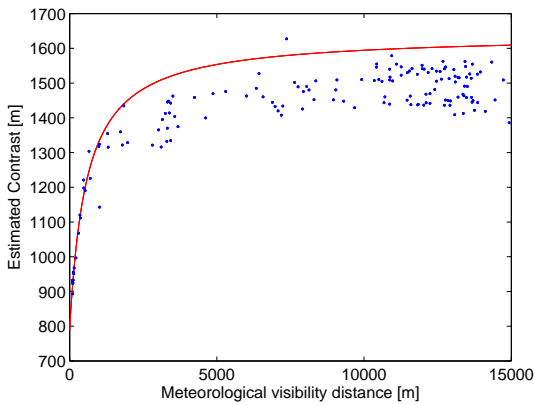


Figure 7: Enhanced data fitting process with the exponential distribution model on short visibility data and extrapolated on higher visibility ranges. The data are plotted in blue. The fitted model is plotted in red.

and can be guessed using only two images grabbed in different foggy weather conditions. Visibility data are thus no longer required, which is an important progress towards generic methods without any learning phase.

However, a continuous distribution of object with respect to the distance in the scene is assumed. This assumption may be false in urban areas, where depth discontinuities exist because of the presence of vertical buildings. Using the actual depth distribution of the scene could improve the results. In this aim, spatial information systems could be used to estimate a more accurate depth distribution at a camera location, so as to get rid of the process proposed in section 3.4. However, it may be difficult to register accurately the 3-D GIS on the image.

Another limitation observed in the test scene is due to the fact that the range of distribution of Lambertian targets is a few hundred meters. We are thus not able to use all the visual cues which are present in the landscape. This can be due to non-uniform illumination when selecting Lambertian targets and could be reduced using the image processing filter proposed by Liaw et al. (2010). A second solution consists in changing the location of the camera and for example increasing its mounting height, so as to get a better perspective. A complementary solution consists in using a camera of better quality, so as to get less noisy images. We aim at exploring these different strategies.

Nevertheless, thanks to results shown in this article, we believe that an ordinary camera is able to monitor the atmospheric visibility whatever the envisaged application: road safety, aeronautic and air quality. This allows envisioning the development of multipurpose environmental monitoring cameras.

6. Conclusion

Camera-based methods are being developed to estimate the atmospheric visibility. However, the methods are either dedicated to road safety (low visibility ranges) or to air quality monitoring (high visibility ranges).

In this article, a generic model-driven approach is presented, which estimates the atmospheric visibility distance through use of ordinary outdoor cameras based on the contrast expectation in the scene. Unlike previous data-driven approaches, a physics-based model is proposed which expresses the mapping between the contrast and the atmospheric visibility distance. Contrary to previous approaches, the model is non-linear which explains why it is able to encompass a larger spectrum of applications. Due to its intrinsic physical constraints, the calibration of the system is also less sensitive to the input data. In particular, the model takes into account the actual distribution of visual targets in the scene, which is estimated by a novel dedicated process which only needs two different fog images. Visibility data are thus not mandatory anymore to calibrate the system. It is also invariant to illumination variations in the scene by selecting the Lambertian surfaces in the scene.

To evaluate the relevance of our approach, the publicly available MALTILDA database is used. Using these experimental data, promising results are obtained, which improve the previous results obtained with this database. When models are fitted to all data, data-driven approaches seem to be more effective for high visibility ranges. When the non-linear models are fitted to the reliable data only, the data-driven approach and the model-driven approach give more or less the same results.

In future work, an ambitious objective is to estimate the contrast expectation function without any additional meteorological sensor, based only on the characteristics of the camera and the properties of the scene (geometry, texture) collected by remote sensing techniques. Such a generic model-driven approach would pave the road to methods without any constraining learning phase.

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