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# prDeep: Robust Phase Retrieval with a Flexible Deep Network

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## Abstract

Phase retrieval algorithms have become an important component in many modern computational imaging systems. For instance, in the context of ptychography and speckle correlation imaging, they enable imaging past the diffraction limit and through scattering media, respectively. Unfortunately, traditional phase retrieval algorithms struggle in the presence of noise. Progress has been made recently on developing more robust algorithms using signal priors, but at the expense of limiting the range of supported measurement models (e.g., to Gaussian or coded diffraction patterns). In this work we leverage the regularization-by-denoising framework and a convolutional neural network denoiser to create *prDeep*, a new phase retrieval algorithm that is both robust and broadly applicable. We test and validate prDeep in simulation to demonstrate that it is robust to noise and can handle a variety of system models.

## 1. Introduction

The PR problem manifests when one wants to recover the input from only the amplitude, or intensity, of the output of a linear system. Mathematically, PR refers to the problem of recovering a vectorized signal  $x \in \mathbb{R}^n$  or  $\mathbb{C}^n$  from measurements  $y$  of the form

$$y = |\mathbf{A}x| + w, \quad (1)$$

where the measurement matrix  $\mathbf{A}$  represents the forward operator of the system and  $w$  represents noise.

PR shows up in many imaging applications including microscopy (Zheng & Yang, 2013), crystallography (Millane, 1990), astronomical imaging (Dainty & Fienup, 1987), and inverse scattering (Katz et al., 2014; Metzler et al., 2017b), to name just a few.

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PR algorithms were first developed in the early 1970s and have been continuously studied by the optics community since then (Gerchberg, 1972; Fienup, 1978; 1982; Griffin & Lim, 1984; Pfeifer et al., 2006; Rodriguez et al., 2013). More recently, PR has been taken up by the optimization community. This has produced a number of algorithms with theoretical, if not always practical, benefits (Candes et al., 2013; 2015a; Goldstein & Studer, 2016; Bahmani & Romberg, 2017). For a benchmark study of over a dozen popular PR algorithms, see PhasePack (Chandra et al., 2017).

Following the popularity of compressive sensing, numerous algorithms were developed that use prior information, oftentimes sparsity, to improve PR reconstructions and potentially enable *compressive* PR (Moravec et al., 2007; Schniter & Rangan, 2015). In general, these methods do not improve reconstructions when the signal of interest is dense.

In the last few years, three methods, SPAR (Katkovnik & Astola, 2012; Katkovnik, 2017), BM3D-prGAMP (Metzler et al., 2016b), and Plug-and-Play ADMM (Venkatakrishnan et al., 2013; Heide et al., 2016), have been developed to solve the PR problem using natural-image priors, which make their reconstructions more robust to noise. These methods all apply a natural-image prior via the BM3D image-denoising algorithm (Dabov et al., 2007).

Unfortunately, two of these three methods, SPAR and BM3D-prGAMP, are restricted to i.i.d Gaussian or coded diffraction measurements (Candes et al., 2015b), which prevents their use in many practical applications. All three methods are computationally demanding.

In this work, we make two technical contributions. First, we show how the Regularization by Denoising (RED) (Romano et al., 2017) framework can be adapted to solve the PR problem. We call this adaption prRED. Because it sets up a general optimization problem, rather than using a specific algorithm, prRED is flexible and can handle a wide variety of measurements including, critically, Fourier measurements.

Second, we show how prRED can utilize convolutional neural networks by incorporating the DnCNN neural network (Zhang et al., 2017). We call this combination of RED and DnCNN applied to PR prDeep.

prDeep offers excellent performance with reasonable run

times. In Section 4, we apply prDeep to simulated data and show that it compares favorably to existing algorithms with respect to computation time and robustness to noise.

## 2. Related Work

Our work fits into the recent trend of using advanced signal priors to solve inverse problems in imaging. While priors like sparsity, smoothness, and structured sparsity have been studied for some time, we focus here on plug-and-play priors and deep-learning priors, which together represent the state-of-the-art for a range of imaging recovery tasks.

### 2.1. Plug-and-Play Regularization for Linear Inverse Problems

Image denoising is arguably the most fundamental problem in image processing and, as such, it has been studied extensively. Today there exist hundreds of denoising algorithms that model and exploit the complex structure of natural images in order to remove additive noise.

Earlier this decade, researchers realized they could leverage these highly developed denoising algorithms to act as regularizers in order to solve other linear inverse problems, such as deblurring, superresolution, and compressed sensing (Danielyan et al., 2010). This technique was eventually coined *plug-and-play* regularization (Venkatakrisnan et al., 2013), with the idea being that one could “plug in” a denoiser to impose a specific prior on the inverse problem. A number of techniques have been developed that use this idea to solve various linear inverse problems (Danielyan et al., 2010; Venkatakrisnan et al., 2013; Heide et al., 2014; 2016; Metzler et al., 2016a; Schniter et al., 2016; Romano et al., 2017).

Most of these works implicitly assume that the denoiser is a proximal mapping for some cost function  $R(x)$ , i.e.,

$$D(z) = \text{prox}_R(z) \triangleq \arg \min_x \frac{1}{2} \|z - x\|_2^2 + R(x), \quad (2)$$

where  $R(x)$  penalizes image hypotheses that are unnatural. From the maximum a posteriori (MAP) Bayesian perspective,  $R(x)$  is the negative log-prior for the natural-image  $x$ , and  $z$  is a Gaussian-noise corrupted measurement of  $x$ .

With this interpretation, a variety of algorithms, such as (Plug-and-Play) ADMM (Venkatakrisnan et al., 2013; Heide et al., 2014; 2016) or (Denoising-based) AMP (Metzler et al., 2016a; Schniter et al., 2016), can be used to recover  $x$  from the linear measurements  $y$  by solving the optimization problem

$$\arg \min_x \frac{1}{2} \|y - \mathbf{A}x\|_2^2 + R(x), \quad (3)$$

where again  $R(x)$  is an implicit cost function associated with the denoiser.

Because the priors associated with advanced denoisers like BM3D accurately model the distribution of natural images, these methods have offered state-of-the-art recovery accuracy in many of the tasks to which they have been applied.

### 2.2. Plug-and-Play Regularization for PR

Following their success on linear inverse problems, plug-and-play priors were applied to the PR problem as well. We are aware of three prior works that take this approach.

The first, SPAR (Katkovnik & Astola, 2012; Katkovnik, 2017), uses alternating minimization to compute the MAP estimate of  $x$  by using BM3D and a Poisson noise model;  $y^2 = \text{Poisson}(|z|^2)$ , with  $z \triangleq \mathbf{A}x$ . So far, the algorithm has only been successfully applied to coded-diffraction pattern measurements.

The second, BM3D-prGAMP (Metzler et al., 2016b), uses the generalized approximate message passing (AMP) framework (Donoho et al., 2009; Rangan, 2011; Schniter & Rangan, 2015; Metzler et al., 2016a) to compute the minimum mean squared error (MMSE) estimate of  $x$  using BM3D and a Rician channel model:  $y = |z + w|$  with  $w \sim CN(0, \sigma_w^2)$ . It requires the elements of  $\mathbf{A}$  to be nearly i.i.d. Gaussian, which can be approximated using coded-diffraction-pattern measurements.

Most recently, the authors of ProxImaL (Heide et al., 2016) used Plug-and-Play ADMM to estimate  $x$  by solving the optimization problem

$$\arg \min_x \frac{1}{2} \|y - \mathbf{A}x\|_2^2 + R(x), \quad (4)$$

where  $R(x)$  is the cost function implicitly minimized by BM3D. Unlike the aforementioned methods, ADMM supports generic measurement matrices, including Fourier measurements.

### 2.3. Neural Networks for Linear Inverse Problems

Deep learning has recently disrupted computational imaging. Through the use of elaborate learned priors, deep learning methods have competed with and sometimes surpassed the performance of plug-and-play priors, while running significantly faster (when implemented on a GPU). Two prominent examples include SRCNN for superresolution (Dong et al., 2014) and DnCNN for denoising, superresolution, and the removal of JPEG artifacts (Zhang et al., 2017).

In addition, a few works have blended plug-and-play algorithms with neural networks (Chang et al., 2017; Metzler et al., 2017a; Diamond et al., 2017), often using algorithm unfolding/unrolling (Gregor & LeCun, 2010). In doing so, these works are able to incorporate powerful learned priors, while still leveraging the flexibility and interpretability that comes from using a well-defined algorithm (as opposed to a

black-box neural net).

## 2.4. Neural Networks for PR

In the last few years, researchers have raced to apply deep learning to solve the PR problem (Kappeler et al., 2017; Boominathan et al., 2018; Rivenson et al., 2017). So far, each of the proposed methods has been designed for specific PR application, either Ptychography (Kappeler et al., 2017; Boominathan et al., 2018) or Holography (Rivenson et al., 2017). The Ptychography neural networks learn to combine a stack of low-resolution band-pass-filtered images to form a high resolution image. The holographic neural network learns how to remove the twin image component (Goodman, 2005) from a hologram. Because each of these networks learns an application-specific mapping, they do not generalize to new PR problems.

Our work takes a different tack. Rather than setting up a neural network to solve a specific PR problem, we use a neural network as a regularizer within an optimization framework. This technique makes our network applicable to numerous PR problems.

## 3. PR via Regularization by Denoising

In this section we first show how the Regularization by Denoising (RED) (Romano et al., 2017) framework can be adapted to solve the PR problem. We call this adaptation prRED. Later, we combine prRED with the state-of-the-art DnCNN neural network (Zhang et al., 2017) to form prDeep.

### 3.1. RED

RED is an algorithmic approach to solving imaging inverse problems that was recently proposed by Romano, Elad, and Milanfar. Like many of the plug-and-play techniques described in Section 2.1, RED can incorporate an arbitrary denoiser to regularize an arbitrary imaging inverse problem. However, whereas the other methods use a denoiser to minimize some *implicit* cost function, RED uses a denoiser to setup and then minimize an *explicit* cost function.

In particular, the RED framework defines the regularizer as

$$R(x) = \frac{\lambda}{2} x^T (x - D(x)), \quad (5)$$

where  $D(x)$  is an arbitrary denoiser. Note that this regularizer serves two roles. First, it penalizes the residual difference between  $x$  and its denoised self; when  $x - D(x)$  is large,  $R(x)$  will tend to be large. Second, it penalizes correlations between  $x$  and the residual. This serves to prevent  $D(x)$  from removing structure from  $x$ ; if  $D(x)$  removes structure from  $x$ , then this structure will show up in the residual, which will be correlated with  $x$ . In effect,

the RED regularizer encourages the residual to look like additive white Gaussian noise (Romano et al., 2017).

When  $D(x)$  satisfies homogeneity and passivity conditions (see Section 3.1 of (Romano et al., 2017)) the proximal mapping of the RED regularization (5) can be implemented recursively as follows<sup>1</sup>

$$\begin{aligned} \text{prox}_R(z) &= \arg \min_x \frac{1}{2} \|x - z\|_2^2 + \frac{\lambda}{2} x^T (x - D(x)) \\ &= v_\infty, \end{aligned}$$

where  $v_j = \frac{1}{1+\lambda}(v_{j-1} + \lambda D(v_{j-1})) \forall j > 0$  and  $v_0 = z$ . In practice, the iterations must be terminated after a finite number  $j$ . However, experiments suggest that  $j = 1$ , which corresponds to calling the denoiser only once per use of the proximal mapping function, leads to good performance.

### 3.2. prRED

To apply RED to PR, we construct a cost function of the form

$$f(x) + R(x), \quad (6)$$

where  $R(x)$  is the RED regularization from (5) and  $f(x)$  is a data-fidelity term that encourages  $\mathbf{A}x$  to match the phaseless measurements  $y$ .

From a Bayesian perspective, the data-fidelity term  $f(x)$  should be proportional to the negative log-likelihood function. For instance, if  $y^2 = |\mathbf{A}x|^2 + w$  with  $w \sim \mathcal{CN}(0, \sigma_w^2 \mathbf{I})$ , then the negative log-likelihood function would be  $-\log p(y|x) \propto \|y^2 - |\mathbf{A}x|^2\|^2$ .

The Bayesian perspective suggests that, when dealing with Poisson noise, which is the focus in this paper, one should use the Poisson log-likelihood function. Interestingly, we experimented with the Poisson log-likelihood function as derived in (Chen & Candes, 2015) and found that it performed slightly worse than the amplitude loss function  $f(x) = \frac{1}{2} \|y - |\mathbf{A}x|\|_2^2$ . This surprising behavior was also noted in (Yeh et al., 2015).

In any case, a variety of data-fidelity terms can be used to solve the PR problem. We adopt the amplitude loss  $\frac{1}{2} \|y - |\mathbf{A}x|\|_2^2$ , which leads us to the non-convex optimization problem

$$\arg \min_x \frac{1}{2} \|y - |\mathbf{A}x|\|_2^2 + \lambda x^T (x - D(x)). \quad (7)$$

Although various solvers could be used to attack (7), we use the FASTA solver (Goldstein et al., 2014). FASTA implements the forward-backward splitting algorithm (a.k.a. the

<sup>1</sup>As these properties do not hold exactly in practice (Reehorst & Schniter, 2018), (6) should be considered only an approximation.

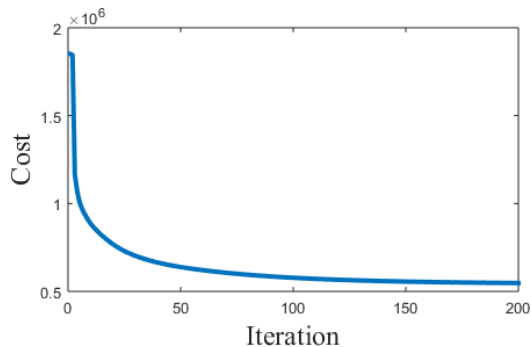


Figure 1. The cost (7) over the number of FASTA iterations. The cost decreases monotonically.

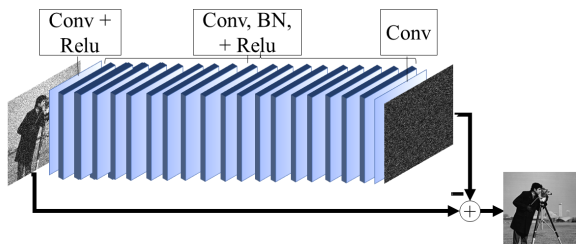


Figure 2. Network architecture of the DnCNN denoiser. Note that it takes advantage of residual learning.

proximal gradient method) and incorporates adaptive step sizes for acceleration. FASTA is very simple to use; after defining the loss function, one need only provide the solver with a proximal mapping for the regularization term (6) and a (sub)gradient for the data-fidelity term with respect to  $z$  (for  $z \triangleq \mathbf{A}x$ ). For our adopted data-fidelity term, a useful subgradient is

$$z - y \circ \frac{z}{|z|} \in \partial_z \frac{1}{2} \|y - |z|\|_2^2, \quad (8)$$

where  $\circ$  denotes the Hadamard (i.e., elementwise) product and  $\partial_z f(z)$  denotes the subdifferential of  $f$  with respect to  $z$ .

In practice, the FASTA solver converges quickly. As an example, Figure 1 shows a typical cost (7) per iteration trajectory for the FASTA solver. There it can be seen that the cost drops monotonically and converges after about 200 iterations.

### 3.3. prDeep

The prRED framework can incorporate nearly any denoising algorithm. We call the special case of prRED with the DnCNN denoiser “prDeep”.

DnCNN (Zhang et al., 2017) is a state-of-the-art denoiser for removing additive white Gaussian noise from natural images. DnCNN consists of 16 to 20 convolutional layers

(we used 20) of size  $3 \times 3 \times 64$ . Sandwiched between these layers are ReLU (Krizhevsky et al., 2012) and batch-normalization (Ioffe & Szegedy, 2015) operations. DnCNN is trained using residual learning (He et al., 2016).

In practice, DnCNN noticeably outperforms the popular BM3D algorithm. Moreover, thanks to parallelization and GPU computing, it runs hundreds of times faster than BM3D.

We trained four DnCNN networks at different noise levels. To train, we loosely followed the procedure outlined in (Zhang et al., 2017). In particular, we trained with 300 000 overlapping patches drawn from 400 images in the Berkeley Segmentation Dataset (Martin et al., 2001). For each image patch, we added additive white Gaussian noise with a standard deviation of either 60, 40, 20, or 10, where our images had a dynamic range of  $[0, 255]$ . We then setup DnCNN to recover the noise-free image. We used the mean-squared error between the noise-free ground truth image and our denoised reconstructions as the cost function. We trained the network with stochastic gradient descent and the ADAM optimizer (Kingma & Ba, 2014) with a batch size of 256. Our training rate was 0.001, which we dropped to 0.0001 and then 0.00001 when the validation error stopped improving. Training took just over 3 hours per noise level on an Nvidia Pascal Titan X.

## 4. Experimental Results

In this section we compare prDeep to several other PR algorithms on simulated data with varying amounts of Poisson noise. We test the algorithms with both coded diffraction pattern (CDP) and Fourier measurements. In both sets of tests, we sample and reconstruct 6 “natural” and 6 “unnatural” (real and nonnegative) test images, which are presented in Figures 3 and 4.

### 4.1. Experimental Setup

**Competing Algorithms.** We compare prDeep against Hybrid Input-Output (HIO) (Fienup, 1982), Oversampling Smoothness (OSS) (Rodriguez et al., 2013), Wirtinger Flow (WF) (Candes et al., 2015a), DOLPHIn (Tillmann et al., 2016), SPAR (Katkovnik & Astola, 2012; Katkovnik, 2017), and BM3D-prGAMP (Metzler et al., 2016b). We also compare with Plug-and-Play ADMM (Venkatakrishnan et al., 2013; Heide et al., 2016) using both the BM3D and DnCNN denoisers. HIO and WF are baseline algorithms designed for Fourier and CDP measurements, respectively. OSS is an alternating projection algorithm designed for noisy Fourier measurements. It imposes a smoothness constraint to regions outside of the target’s support. DOLPHIn is an iterative PR algorithm designed to reconstruct images from noisy CDP measurements. It imposes a sparsity constraint with

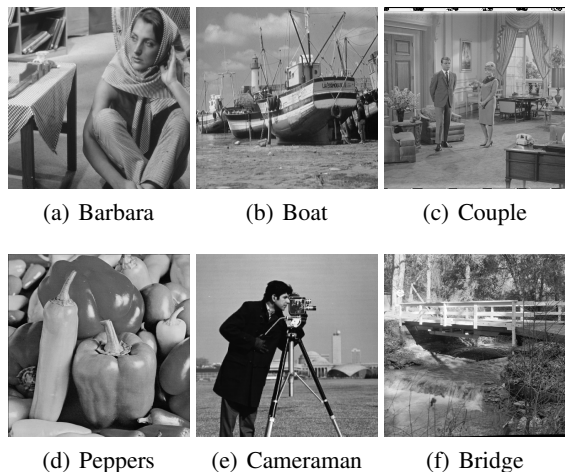


Figure 3. The six “natural” test images. They follow a distribution very similar to the training data.

respect to a learned dictionary. SPAR, BM3D-prGAMP, and Plug-and-Play ADMM were described in Section 2.2.

**Implementation.** All algorithms were tested using Matlab 2017a on a desktop PC with an Intel 6800K CPU and an Nvidia Pascal Titan X GPU. Dolphin, SPAR, and BM3D-prGAMP used their respective authors’ implementations. We created our own version of Plug and Play ADMM based off of code original developed in (Chan et al., 2017). We use FASTA to solve the nonlinear least squares problem at each iteration of the algorithm. prDeep and DnCNN-ADMM used a MatConvNet (Vedaldi & Lenc, 2015) implementation of DnCNN. A public implementations of prDeep is available at <https://github.com/ricedsp/prDeep>.

**Measurement and Noise Model.** Model mismatch and Poisson shot noise are the dominant sources of noise in many PR applications (Yeh et al., 2015). In this paper, we focus on shot noise, which we approximate as

$$y^2 = |z|^2 + w \text{ with } w \sim N(0, \alpha^2 \text{Diag}(|z|^2)),$$

where  $z = \mathbf{A}x$  with known  $\mathbf{A}$ , and where  $\text{Diag}(|z|^2)$  is a diagonal matrix with diagonal elements  $|z|^2$ .

Some algebra and the central limit theorem can be used to show that  $y^2/\alpha^2 \sim \text{Poisson}((|z|/\alpha)^2)$ . In effect,  $y^2$  is a rescaled Poisson random variable. The term  $\alpha$  controls the variance of the random variable and thus the effective signal-to-noise ratio in our problem.

**Parameter Tuning.** HIO was run for 1000 iterations. WF was run for 2000 iterations. BM3D-prGAMP was run for 50 iterations. prDeep was run for 200 iterations four times; once for each of the denoisers networks (trained at stan-

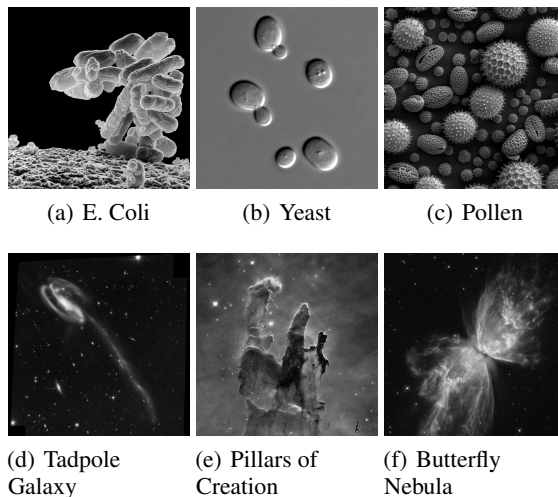


Figure 4. The six “unnatural” test images. They follow a distribution distinct from the training data. Images are from Wikipedia and follow public domain licenses.

dard deviations 60, 40, 20, and 10). The result from reconstructing with the first denoiser was used to warm-start reconstructing with the second, the second warm-started the third, etc. Plug and Play ADMM was similarly run for 50 iterations four times. ADMM converged faster than prDeep and did not benefit from additional iterations.

SPAR reconstructed  $x$  using  $y^2/\alpha^2$  rather than  $y$ , as its loss function expects Poisson distributed random variables. prDeep’s parameter  $\lambda$ , which determines the amount of regularization, was set to  $\bar{\sigma}_w$  when dealing with Fourier measurements and  $0.1\bar{\sigma}_w$  when dealing with CDP measurements, where  $\bar{\sigma}_w^2$  denotes the sample variance of the noise. The parameter in ADMM analogous to  $\lambda$  was set to  $.2\bar{\sigma}_w$  when dealing with Fourier measurements and  $0.02\bar{\sigma}_w$  when dealing with CDP measurements. The algorithms otherwise used their default parameters.

**Initialization.** With oversampled CDP measurements of real-valued signals, none of the algorithms were particularly sensitive to initialization; initializing with a vector of ones worked sufficiently well. In contrast, with Fourier measurements, the algorithms were very sensitive to initialization. We experimented with various spectral initializers, but found they were ineffective with (noisy and only  $4\times$  oversampled) Fourier measurements. Instead, we first ran the HIO algorithm (for 50 iterations) 50 times, from random initializations, to form 50 estimates of the signal:  $\hat{x}_1, \hat{x}_2, \dots, \hat{x}_{50}$ . We then used the reconstruction  $\hat{x}_i$  with the lowest residual ( $\|y - |\mathbf{A}\hat{x}_i|\|_2$ ) as an initialization for HIO. HIO was then run for 1000 iterations, and the result was used to initialize the other algorithms. This process was repeated three times and the reconstruction with the smallest residual

Table 1. PSNRs and run times (sec) of  $128 \times 128$  reconstructions with four intensity-only CDP measurements and varying amounts of Poisson noise.

	$\alpha = 9$			$\alpha = 27$			$\alpha = 81$		
	PSNR	PSNR	Time	PSNR	PSNR	Time	PSNR	PSNR	Time
	Natural	Unnatural		Natural	Unnatural		Natural	Unnatural	
HIO	36.1	36.0	4.2	26.1	26.0	5.9	14.8	16.9	5.9
WF	34.3	34.2	0.8	24.7	24.0	0.9	13.2	13.0	6.7
DOLPHIn	31.0	28.8	1.1	27.9	26.9	1.1	17.9	20.8	1.1
SPAR	34.4	36.0	15.7	29.4	31.1	15.7	<b>24.5</b>	26.4	15.7
BM3D-prGAMP	34.1	35.4	17.2	31.4	32.2	17.7	22.5	22.3	23.1
BM3D-ADMM	38.5	39.0	51.6	31.9	33.4	40.0	24.4	<b>26.9</b>	37.7
DnCNN-ADMM	<b>39.9</b>	<b>40.6</b>	17.1	32.6	33.4	9.8	21.6	22.7	7.9
prDeep	39.1	40.1	25.4	<b>32.8</b>	<b>34.0</b>	23.2	20.9	24.5	23.1

Table 2. PSNRs and run times (sec) of  $128 \times 128$  reconstructions with  $4 \times$  over-sampled intensity-only Fourier measurements and varying amounts of Poisson noise.

	$\alpha = 2$			$\alpha = 3$			$\alpha = 4$		
	PSNR	PSNR	Time	PSNR	PSNR	Time	PSNR	PSNR	Time
	Natural	Unnatural		Natural	Unnatural		Natural	Unnatural	
HIO	22.2	20.8	10.6	19.9	17.9	10.4	17.9	15.5	10.7
WF	15.2	18.8	6.7	15.1	18.5	6.6	15.2	18.5	6.7
OSS	18.6	23.8	18.3	18.4	23.1	18.6	18.6	22.7	18.1
SPAR	22.0	23.9	79.1	19.6	22.6	77.0	19.4	20.5	83.0
BM3D-prGAMP	24.5	26.0	84.3	23.2	24.4	82.8	21.0	22.7	84.4
BM3D-ADMM	27.5	28.3	155.7	24.5	25.0	156.8	21.8	23.1	156.8
DnCNN-ADMM	<b>29.3</b>	<b>31.3</b>	65.0	26.0	25.8	62.4	22.0	23.4	64.0
prDeep	28.4	30.6	105.1	<b>28.5</b>	<b>26.4</b>	105.0	<b>26.4</b>	<b>25.1</b>	107.1

was used as the final estimate. The reported computation times (for Fourier measurements) include the time required to initialize the algorithms and run them three times.

#### 4.2. Simulated Coded Diffraction Measurements

We first test the algorithms with CDP intensity-only measurements. CDP is a measurement model proposed in (Candes et al., 2015b) that uses a spatial light modulator (SLM) to spread a target’s frequency information and make it easier to reconstruct. Under a CDP measurement model, the target is illuminated by a coherent source and then has its phase immediately modulated by a known random pattern using an SLM. The complex field then undergoes far-field Fraunhofer diffraction, which can be modeled by a 2D Fourier transform, before its intensity is recorded by a standard camera. Multiple measurements, with different random SLM patterns, are recorded. In this work, we model the capture of four measurements using a phase-only SLM. Mathemati-

cally, our measurement operator is as follows

$$\mathbf{A} = \begin{bmatrix} \mathbf{F}\mathbf{D}_1 \\ \mathbf{F}\mathbf{D}_2 \\ \mathbf{F}\mathbf{D}_3 \\ \mathbf{F}\mathbf{D}_4 \end{bmatrix}, \quad (9)$$

where  $\mathbf{F}$  represents the 2D Fourier transform and  $\mathbf{D}_1, \mathbf{D}_2, \dots$  are diagonal matrices with nonzero elements drawn uniformly from the unit circle in the complex plane.

In Table 1 we compare the performance of the various PR algorithms. We do not include a comparison with OSS as it is setup specifically for Fourier measurements. We report recovery accuracy in terms of mean peak-signal-to-noise ratio (PSNR) across two sets of test images. We report run time in seconds.

Table 1 demonstrates that, when dealing with CDP measurements at low SNRs (large  $\alpha$ ), all five plug-and-play methods produce similar reconstructions.

<sup>1</sup>PSNR =  $10 \log_{10}(\frac{255^2}{\text{mean}((\hat{x} - x_o)^2)})$  when the pixel range is 0 to 255.

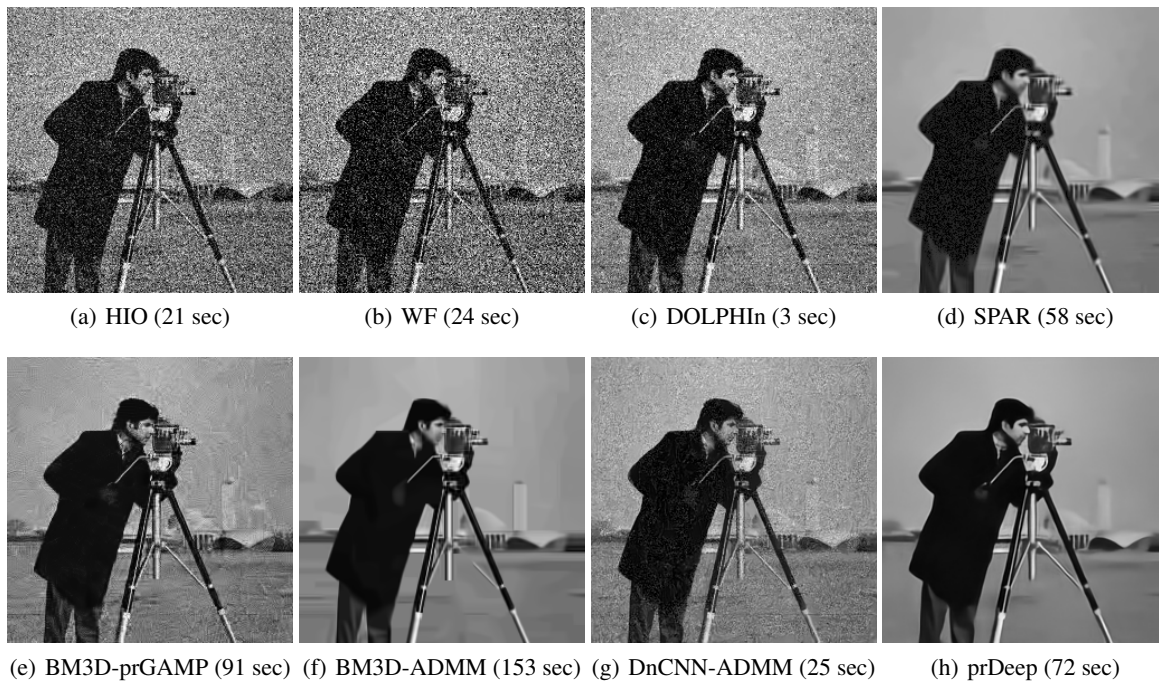


Figure 5. Reconstructions of  $256 \times 256$  image from four noisy intensity-only CDP measurements ( $\alpha = 81$ ) with several PR algorithms. All five plug-and-play methods provide similar reconstructions.

In Figure 5, we visually compare the reconstructions of a  $256 \times 256$  image using the algorithms under test. With CDP measurements, all of the plug-and-play algorithms do a good job reconstructing the signal from noisy measurements.

### 4.3. Simulated Fourier Measurements

Fourier measurements of real-valued signals are prevalent in many real-world applications. Many PR applications exploit the Fourier-transform property  $\mathbf{F}(x \star x) = |\mathbf{F}x|^2$ , where  $\star$  denotes correlation. This implies that in applications where one can measure or estimate the autocorrelation function of an object one can also measure the modulus squared of its Fourier transform. This allows PR algorithms to reconstruct the object.

This relationship has been used in multiple contexts. In astronomical imaging, this relationship has been used to image through turbulent atmosphere (Dainty & Fienup, 1987). In laser-illuminated imaging, this relationship has been used to reconstruct diffuse objects without speckle noise (Fienup & Idell, 1988). More recently, this relationship has been used to image through random scattering media, such as biological tissue (Katz et al., 2014). In all of these applications, one reconstructs the real-valued intensity distribution of the object.

In these tests we oversampled the spectrum by  $4\times$ . That is, we first placed  $128 \times 128$  images at the center of a  $256 \times 256$

square and then took the 2D Fourier transform. We assumed that the support, i.e., the location of the image within the  $256 \times 256$  grid, was known a priori.

Table 2 compares the reconstruction accuracies and recovery times of several PR algorithms.<sup>2</sup> We do not include results for DOLPHIn, as it completely failed with Fourier measurements. Note that, at a given noise level, the reconstructions from Fourier measurements are far less accurate than their CDP counterpart. Table 2 demonstrates that with Fourier measurements and large amounts of noise prDeep is superior to existing PR algorithms.

In Figure 6, we again visually compare the reconstructions from the algorithms under test, this time with Fourier measurements. In this regime prDeep produces fewer artifacts than competing methods.

## 5. Conclusions and Future Work

In this paper, we have extended and applied the Regularization by Denoising (RED) framework to the problem of PR. Our new algorithm, prDeep, is exceptionally robust to noise thanks to the use of the DnCNN image denoising neural network. As we demonstrated in our experiments, prDeep is also able to handle a wide range of measurement

<sup>2</sup>Our results account for the translation and reflection ambiguities associated with Fourier measurements

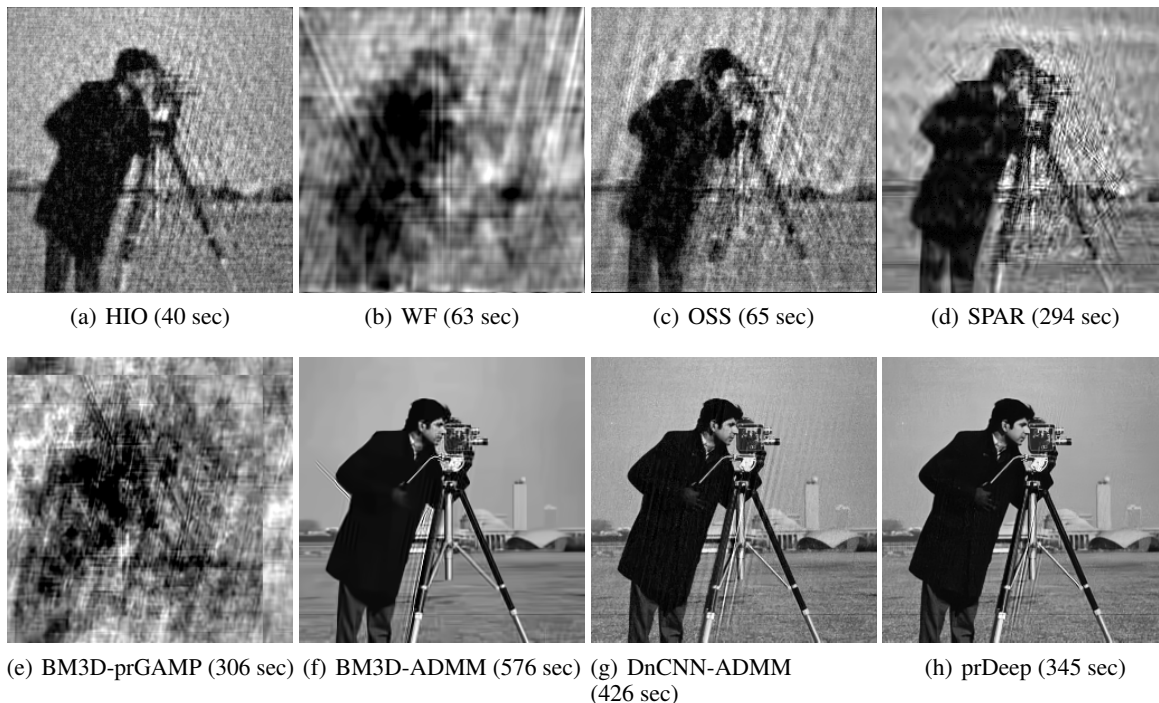


Figure 6. Reconstructions of  $256 \times 256$  image from noisy  $4 \times$  oversampled intensity-only Fourier measurements ( $\alpha = 3$ ) with several PR algorithms. Plug and Play ADMM and prDeep provide the best reconstructions.

matrices, from intensity-only coded diffraction patterns to Fourier measurements.

By integrating a neural network into a traditional optimization algorithm, prDeep inherits the strengths of both optimization and deep-learning. Like other optimization-based algorithms, prDeep is flexible and can be applied to PR problems with different measurement models, noise levels, etc., without having to undergo costly retraining. Like other deep-learning-based techniques, for any given problem prDeep can take advantage of powerful learned priors and outperform traditional, hand-designed methods.

prDeep is not a purely academic PR algorithm; it can handle Fourier measurements of real-valued signals, a measurement model that plays a key role in many imaging applications. However, prDeep does have a major limitation; it is presently restricted to amplitude-only targets. Extending prDeep to handle complex-valued targets is a promising and important direction for future research.

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