

Article

## A New Database of Global and Direct Solar Radiation Using the Eastern Meteosat Satellite, Models and Validation

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**Abstract:** We present a new database of solar radiation at ground level for Eastern Europe and Africa, the Middle East and Asia, estimated using satellite images from the Meteosat East geostationary satellites. The method presented calculates global horizontal (G) and direct normal irradiance (DNI) at hourly intervals, using the full Meteosat archive from 1998 to present. Validation of the estimated global horizontal and direct normal irradiance values has been performed by comparison with high-quality ground station measurements. Due to the low number of ground measurements in the viewing area of the Meteosat Eastern satellites, the validation of the calculation method has been extended by a comparison of the estimated values derived from the same class of satellites but positioned at 0°E, where more ground stations are available. Results show a low overall mean bias deviation (MBD) of +1.63 Wm<sup>-2</sup> or +0.73% for global horizontal irradiance. The mean absolute bias of the individual station MBD is 2.36%, while the root mean square deviation of the individual MBD values is 3.18%. For direct normal irradiance the corresponding values are overall MBD of +0.61 Wm<sup>-2</sup> or +0.62%, while the mean absolute bias of the individual station MBD is 5.03% and the root mean square deviation of the individual MBD values is 6.30%. The resulting database of hourly solar radiation values will be made freely available. These data will also be integrated into the PVGIS web application to allow users to estimate the energy output of photovoltaic (PV) systems not only in Europe and Africa, but now also in Asia.

**Keywords:** solar radiation; satellite-based retrieval; global horizontal irradiance; direct normal irradiance

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## Nomenclature

### Acronyms

BSRN	Baseline Surface Radiation Network
CM SAF	Climate Monitoring Satellite Application Facility
DNI	Direct Normal Irradiance
ECMWF	European Centre for Medium-range Weather forecast
GAW	Global Atmosphere Watch
MACC	Monitoring Atmospheric Composition and Climate
MFG	Meteosat First Generation satellites
MSG	Meteosat Second Generation satellites
NWP	Numerical Weather Prediction
PV	Photo-Voltaic
PVGIS	PhotoVoltaic Geographical Information System
RTM	Radiative Transfer Model
SAL	Surface ALbedo
USGS	United States Geological Survey

### Symbols

$AM$	Air Mass (-)
$B, B_{\text{clear}}, B_{\text{all}}$	Direct (beam) radiation, clear-sky & all-sky beam radiation ( $\text{Wm}^{-2}$ )
$CAL$	Effective Cloud Albedo, also called Cloud Index (-)
$D$	Digital count of the satellite instrument
$DNI, DNI_{\text{avg}}$	Direct normal irradiance & average $DNI$ of all stations ( $\text{Wm}^{-2}$ )
$G$	Global horizontal irradiance ( $\text{Wm}^{-2}$ )
$G_{\text{avg}}$	Average measured irradiance of all the stations ( $\text{Wm}^{-2}$ )
$G_{\text{meas}}$	Measured Global horizontal irradiance ( $\text{Wm}^{-2}$ )
$G_{\text{clear}}$	Global horizontal clear-sky irradiance ( $\text{Wm}^{-2}$ )
$H$	Global horizontal irradiation ( $\text{kWh m}^{-2}$ )
$H_{\text{DNI}}$	Direct normal irradiation ( $\text{kWh m}^{-2}$ )
$\Delta H$	Relative difference in $H$ between the two satellites (%)
$k_c$	Clear-sky index (-)
$k_t$	Clearness index (-)
$k'_t$	Modified clearness index (-)
MAB, rMAB	Mean Absolute Bias, relative Mean Absolute Bias

gMAB, rgMAB	global Mean Absolute Bias, relative global Mean Absolute Bias
MBD, rMBD	Mean Bias Deviation, relative Mean Bias Deviation
gMBD, rgMBD	global Mean Bias Deviation, relative global Mean Bias Deviation
$N$	Number of data points used for the validation (-)
$N_s$	Number of ground stations used for the validation (-)
$\rho$	Intensity of reflected light measured by satellite (-)
$\rho_{\text{thres}}$	Threshold value of $\rho$ , for calibration of satellite measurements (-)
$\rho_{cs}, \rho_{\text{max}}$	Clear-sky and monthly maximum values of $\rho$ (-)
$\rho_{\text{max,corr}}$	Value of $\rho_{\text{max}}$ corrected as a function of $\theta$ (-)
RMSD, rRMSD	Root Mean Square Deviation, relative Root Mean Square Deviation
gRMSD, rgRMSD	global Root Mean Square Deviation, relative global Root Mean Square Deviation
$\theta$	Solar zenith angle (-)
Subscript $i$	$i$ 'th time series value
Superscript $m$	measured value
Superscript $r$	retrieved (estimated) value

## 1. Introduction

Retrieval of ground-level solar radiation retrieved from satellite has been performed by many groups going back at least to Möser and Raschke, 1984 [1] and Cano *et al.*, 1986 [2]. Another early work is that of Pinker and Laszlo, 1992 [3]. Since then, both geostationary and polar-orbiting satellites have been used to estimate solar radiation at ground level, and existing data sets cover nearly the whole earth, though with widely varying spatial and temporal resolution [4–8].

A number of solar radiation data sets have been retrieved using data from the *Meteosat* family of European geostationary satellites situated at  $0^\circ$  longitude. Examples include SOLEMI [4,9], the different versions of HelioClim [6,10], CM SAF [9,11], LSA SAF [12,13]. Some of these data are available free of charge, while others license the data on a commercial basis. The commercial/free status of the individual data sets may change with time. However, at the time of writing, to our knowledge there are no freely available high-resolution data for solar radiation retrieved from the eastern *Meteosat* satellites situated over the Indian Ocean (These satellites are officially named *Meteosat* Indian Ocean Data Coverage (IODC)). We find this rather uninformative and will refer to the satellites as “*Meteosat* East”).

Here we describe the results of a procedure to retrieve ground-level solar radiation components using data from the Eastern *Meteosat* satellites. The resulting database consists of hourly values of global horizontal ( $G$ ) and direct normal irradiance ( $DNI$ ) at the earth surface with a spatial resolution similar to the native pixel resolution of the satellite images.

The organization of the paper is as follows: Section 2 will give an overview of the methods used for the calculation of solar irradiance from satellite data and the properties of the resulting solar radiation database. Section 3 presents the results of the validation of hourly global horizontal and direct normal irradiance using ground station data. Section 4 discusses the results and presents some applications of the data. Section 5 contains the conclusions.

## 2. Methods and Input Data for the Retrieval of Solar Radiation From Geostationary Satellite Images

### 2.1. Algorithms for Calculating Surface Solar Radiation Components

The cloudy part of the applied method is based on the well established and widely used Heliosat method (please see [5] for an overview of applications). For the calculation of the all sky solar irradiance the SPECMAGIC algorithm is used [14]. This combination is a standard method for the calculation of solar surface radiation of the CM SAF, the Climate Monitoring Satellite Application Facility. The method is described in more detail in the following paragraphs.

The retrieval of the solar surface irradiance is performed in a two step approach. In a first step the effective cloud albedo ( $CAL$ ) is retrieved from the visible broadband channel of the Meteosat satellite. The observed reflection is corrected for different luminance conditions due to variations in the Sun-Earth distance and the solar zenith angle  $\theta$ . Furthermore, the dark offset of the instrument has to be subtracted from the satellite image counts  $D$ . The observed reflections are therefore normalised by application of Equation (1).

$$\rho = \frac{D - D_0}{\cos(\theta)f(\text{distance})} \quad (1)$$

Here,  $D$  is the observed digital count which includes the dark offset  $D_0$  of the satellite instrument.  $f(\text{distance})$  accounts for the variation in the Earth sun distance. The effective cloud albedo is then derived from the observed normalised reflections by Equation (2):

$$CAL = \frac{\rho - \rho_{cs}}{\rho_{\max} - \rho_{cs}} \quad (2)$$

Here,  $\rho$  is the observed reflection for each pixel and time.

$\rho_{\max}$  is the “maximum” reflection. It is determined by the 95 percentile of all reflection values at local noon in a target region (latitude 45.8°S to 52.7°S and longitude 40°E to 56°E) (This region is for Meteosat East. For the Meteosat satellites at 0° longitude the target region is correspondingly further west). This region is characterised by high frequency of cloud occurrence for each month. In this manner changes in the satellite brightness sensitivity are accounted for.  $\rho_{\max}$  depends slightly on the scattering angle, which is the angle between the sun and satellite.

In order to account for this effect the following empirical relation is applied:

$$\rho_{\max, \text{corr}} = \begin{cases} \rho_{\max} (1 + 0.0017 * (45 - \theta)), & 0 < \theta < 80 \\ \rho_{\max}, & \theta > 80 \end{cases} \quad (3)$$

For the fixed target region the satellite angles are constant and no systematic dependency on the sun azimuth has been found. This explains why the correction depends only on the solar zenith angle. Details on evaluation of the self-calibration approach are given in [15].

$\rho_{cs}$  is the clear sky reflection, which is a statistical value derived for every pixel and time slot separately. This is essentially done by estimation of the minimal reflection of the satellite images during a certain time span (e.g., a month) for each time slot. In detail this is done by iteration as follows. At the start all reflection values  $\rho$  (within the time span) are used to calculate an average reflection. This average

reflection serves as initial threshold value  $\rho_{\text{thres}}$ . A new average is then calculated with all reflections that are smaller than the initial threshold value plus a small value  $\epsilon$ , given as  $\epsilon = 0.035\rho_{\text{max}}$ . This average is then the threshold for the next iteration, where again all reflection values below  $\rho_{\text{thres}} + \epsilon$  are used to calculate a new value for  $\rho_{\text{thres}}$ . The iteration is proceeded until no reflection value is available that is higher than  $\rho_{\text{thres}} + \epsilon$ , hence until the threshold  $\rho_{\text{thres}}$  does not change any more.  $\rho_{cs}$  is then defined as being equal to  $\rho_{\text{thres}}$ . It is assumed that at this point all cloudy pixels are filtered out.

As a consequence of the law of energy conservation, radiation which is not reflected or absorbed by clouds passes through the clouds. The effective cloud albedo provides therefore a measure for the cloud transmission, *i.e.*, the amount of clear sky (=cloudless) radiation that passes through the clouds. The irradiance can therefore be calculated by Equation (4).

$$G = (1 - CAL)G_{\text{clear}} \quad (4)$$

Here,  $G$  is the solar surface irradiance, often referred as global irradiance and  $G_{\text{clear}}$  is the radiation for clear (=cloudless) skies.

This relation is valid for a  $CAL$  range between 0 to 0.8, and has to be slightly modified outside of this range, please see [16] or [17] for further details.

In the second step of the calculation, the clear-sky irradiance  $G_{\text{clear}}$  is calculated using the SPECMAGIC model for clear-sky solar irradiance [14].  $G_{\text{clear}}$  is then used to calculate the global irradiance using Equation (4).

SPECMAGIC is described in detail in [14]; here, only a brief outline is given. The clear sky model is based on radiative transfer modelling (RTM). It is a hybrid Look-Up-Table approach. The RTM is used to calculate look-up tables. These consider the effect of absorption and scattering of aerosols as well as rayleigh scattering on the irradiance. The effect of the absorbers water vapour and ozone is considered by parameterisation developed and validated by radiative transfer modelling. SPECMAGIC enables the use of enhanced information about the atmosphere: aerosol optical depth, aerosol type, water vapour and ozone.

In addition to the global irradiance, SPECMAGIC performs also the calculation of the direct irradiance. The same method (hybrid Look-Up-Table approach) is used to derive  $B_{\text{clear}}$ . The beam irradiance for all sky conditions,  $B_{\text{all}}$  is then calculated by the use of Equation (5).

$$B_{\text{all}} = B_{\text{clear}} (k_c - 0.38(1 - k_c))^{2.5} \quad (5)$$

where  $k_c$  is the clear-sky index, which equals  $1 - CAL$  in this case. This formula is an adaptation of the Skartveit diffuse model [18].

## 2.2. Sources of Input Data

The accuracy of the estimated  $G_{\text{clear}}$  and  $B_{\text{clear}}$  depends on the accuracy of the information about aerosols and water vapour, the dominant variables for cloudless sky irradiance.

The water vapour information results from the analysis of the global Numerical Weather Prediction model (NWP) of the European Center for Medium Weather Forecast (ECMWF). The weather predictions system is described in [19], and the data are available online [20].

For the aerosol information, 8 years (2003–2010) of reanalysis data from the project “Monitoring Atmospheric Composition and Climate” (MACC) [21–24] have been used to calculate long term monthly means of aerosol optical depth. These data are in turn used to consider the effect of aerosols on the solar surface irradiance. The MACC reanalysis data is generated on a Gaussian T159 grid which corresponds to 120 km resolution. For the use in SPECMAGIC the aerosol data have been re-gridded to a  $0.5 \times 0.5$  degree regular latitude-longitude grid. Further, high AOD values has been smoothed in order to account for the sensitivity of CAL on high aerosol loads.

Though it would be possible to use the monthly aerosol data we have chosen to use the long-term average monthly values. This choice has been made mainly because the time period of the aerosol data is shorter than that of the satellite data. Using the monthly data for the shorter period would introduce an inconsistency in the time series. It would also make these data inconsistent with the longer time series of solar radiation data produced by the CM SAF collaboration, and reduce their usefulness for climate change studies which are the main motivation for the efforts of CM SAF.

For the surface albedo (SAL) values based on the United States Geographical Survey (USGS) land-cover map [25] are used. A formula given in [26] is applied in order to consider the solar zenith angle dependency of SAL. For ozone concentration a climatological value is used.

The METEOSAT images used for the calculation of *CAL* are available from the EUMETSAT web site [27].

### 2.3. Meteosat Satellites Relevant to the Meteosat East Data Retrieval

The solar radiation database presented here is derived from images obtained by a number of different geostationary meteorological satellites operated by EUMETSAT. These satellites belong the the Meteosat First Generation (MFG) class, and the images are obtained using the MVIRI on-board instrument. The relevant satellites are listed in Table 1, with their positions and periods of operation. The Meteosat satellites are generally placed in one of two geostationary locations: the *Meteosat Prime* location at  $0^\circ$  longitude, and the *Meteosat East* locations over the Indian Ocean, at around  $60^\circ\text{E}$ . In this paper we will use these designations (Prime and East) to distinguish the satellite locations.

**Table 1.** List of relevant Meteosat meteorological satellites, with their active periods.

Satellite	Class	Start Date	End Date	Longitude
Meteosat-5	MFG	2 May 1991	13 February 1997	$0^\circ$
Meteosat-5	MFG	1 July 1998	16 April 2007	$63^\circ\text{E}$
Meteosat-7	MFG	6 March 1998	19 July 2006	$0^\circ$
Meteosat-7	MFG	1 November 2006	ongoing	$57^\circ\text{E}$

As can be seen from Table 1, the satellites Meteosat-5 and Meteosat-7 have both been active over the Meteosat Prime location ( $0^\circ$  longitude) and in different positions over the Indian Ocean. During the period July 1998 to July 2006 there were two satellites with similar instruments in different positions. This period allows a direct comparison between the solar radiation estimates from the two satellites.

#### 2.4. Solar Radiation Retrieval

The computer codes described in Section 2.1 have been applied to produce solar radiation data for a number of years for the area covered by Meteosat Prime and Meteosat East. The Meteosat First Generation (MFG) satellites record an image every half hour, but for the present study, only the full-hour slots have been used.

The calculation of *CAL* is performed on the satellite image in the native resolution and projection. However, for the calculation of the solar radiation, the *CAL* data are projected onto a regular latitude/longitude grid with spatial resolution  $0.05^\circ$  (3 arc-minutes). The accuracy of the calculation of *CAL* decreases towards the edge of the image where the satellite views the ground and clouds at a very shallow angle. For this reason the area in which calculations are performed is restricted to a region  $65^\circ$  from the satellite nadir point (for instance, for Meteosat 5, when located at  $63^\circ\text{E}$ , the limits are  $65^\circ\text{N}$ ,  $65^\circ\text{S}$ ,  $2^\circ\text{W}$  and  $128^\circ\text{E}$ ). In the final output the data are given on a latitude-longitude grid (spatial resolution  $0.05^\circ$ ).

### 3. Validation of Global Horizontal and Direct Normal Irradiance Retrieval from MVIRI Data

The validation of the satellite-based retrieval has been performed by comparison with solar radiation measurements performed at ground level. A number of statistical measures have been used in this validation exercise. The definition of these measures can be found in Appendix.

This section will present the main results of the validation exercise. The data used for the validation as well as more detailed analysis is available at a dedicated web page [28].

#### 3.1. Ground Station Measurement Data Used for the Validation

The main source of ground station measurements used in this study is the Baseline Surface Radiation Network, BSRN [29]. The data from these stations are regarded as being of high quality and have been used extensively for validation of satellite-derived solar radiation products.

The geographical distribution of these stations is very uneven. While there are several stations in Western Europe, North America and Japan, Africa and Asia are only covered in a very sparse fashion. For this reason, a few other stations have been added to the validation exercise, especially for Eastern Europe. These station data are available from the Global Atmosphere Watch (GAW) [30] and distributed by the World Radiation Data Centre [31]. Finally, ground station measurements from the Joint Research Centre, Ispra, Italy have been added to the list.

The stations used in the validation are listed in Table 2.

The BSRN stations typically supply global horizontal (*G*) and direct normal irradiance (*DNI*) values at an interval of 1 min. This is also the case for the Ispra station. However, the data for Kishinev and Thessaloniki are available only as hourly values.

### 3.2. Data Selection and Quality Control

There are two datasets per station. On the one hand, the estimated values at the station's location derived from the satellite images and on the other hand, the measured irradiance data at the station. As previously mentioned, only one satellite image per hour will be considered and processed for this study.

Since the satellite image has a finite resolution (about 3–5 km) each pixel effectively is an average over the corresponding area. This means that it is not possible to capture changes in the solar radiation on timescales less than a few minutes. For this reason, the ground station measurements have been averaged over a 10-min window centered around the time the satellite measures the radiance of the image pixel corresponding to the location of the station. This does not apply to the stations for which only hourly data are available. This difference should be kept in mind when inspecting the results. The longer time averaging for the hourly data may lead to smaller values for the mean absolute deviation and the root-mean-square deviation values, while the mean bias deviation should be relatively unaffected by the choice of averaging period.

**Table 2.** List of radiation measurement stations used in the validation of the satellite-based radiation estimates. The “Prime” and “East” columns indicate whether they were used to validate data from the Prime and East satellite positions, respectively. Kishinev and Thessaloniki have only hourly average data available, for the other stations the 1-min data have been averaged over a 10-min interval.

Station	Lat	Lon	Climate type	Prime	East	Years Used
BSRN stations						
Cabauw (NL)	51.97N	4.93E	Temperate maritime	X		2005
Camborne (UK)	50.22N	5.32W	Temperate maritime	X		2005
Carpentras (FR)	44.05N	5.03E	Mediterranean	X		2005
Cocos Island (AU)	12.19S	96.84E	Tropical wet		X	2009
De Aar (ZA)	30.67S	23.99E	Arid	X	X	2002
Florianopolis (BR)	27.53S	48.52W	Humid subtropical	X		2005
Lindenberg (DE)	52.22N	14.12E	Moderate maritime	X		2005
Payerne (CH)	46.81N	6.94E	Semi continental	X		2005
Sde Boqer (IL)	30.87N	34.77E	Dry steppe	X	X	2005, 2009
Solar Village (SA)	24.91N	46.78E	Arid	X	X	2002
Tamanrasset (DZ)	22.79N	5.53E	Hot, desert	X		2005
Toravere (EE)	58.27N	26.47E	Cold humid	X		2005
Xianghe (CN)	39.75N	116.96E	Continental humid		X	2005
Non-BSRN stations						
Ispra (IT)	45.81N	8.63E	Mediterranean moist	X		2005
Kishinev (MD)	47.00N	28.82E	Cool temperature, moist	X	X	2005
Thessaloniki (GR)	40.63N	22.96E	Mediterranean temperate	X	X	2005

As a result, both estimated and measured irradiance datasets present one value per hour. Therefore the first step is the selection of the closest moments in time from both of them. Only if the time difference



between the estimated value and the nearest time-stamped measured value is below 30 min, these values will be selected; otherwise both values are discarded. So mainly the ground data availability restricts the number of points in the comparison.

Once the closest moments in time have been selected from both datasets, a simple quality control procedure is applied to all pairs of data:

- All missing values are removed. BSRN irradiance data usually present “−999” or “−999.9” values to highlight missing data. In these cases, both measured and the corresponding estimated value are eliminated.
- Negative measured or estimated irradiance values when the sun’s elevation angle is below 0° are substituted by zero values and kept in the datasets.
- However, negative measured or estimated irradiance values when the sun’s elevation angle is above 0° are removed from both datasets. If one of the values in the pair is wrong, the whole pair is discarded.

For the analysis of the validation results, apart from the general comparison between estimated and measured irradiance values of both datasets as a whole, the validation has been performed, also distributing the data according to different criteria, like months, solar coordinates, angular distance between the sun and the satellite or sky type. In order to distribute data into different sky types, the sky classification defined by Ineichen *et al.* [32], has been used. This one is based on a modified clearness index,  $k'_t$ , (Equation (6)), defined by Perez *et al.* [33], which has the advantage of being relatively less dependent on the solar elevation angle than the normal clearness index,  $k_t$ .

$$k'_t = \frac{k_t}{1.031 \cdot \exp \frac{-1.4}{0.9 + \frac{9.4}{AM}} + 0.1} \quad (6)$$

Here  $AM$  is the optical air mass as defined by Kasten [34]. Sky conditions are delineated into three intervals according to the  $k'_t$  value as follows:

- Clear sky conditions:  $0.65 < k'_t \leq 1$ ;
- Intermediate conditions:  $0.3 < k'_t \leq 0.65$ ;
- Overcast conditions:  $0 < k'_t \leq 0.3$ .

For all stations, the validation period is one year. Some of the stations have been compared against both the Meteosat Prime and Meteosat East estimates, if the station lies within the field of view of both satellites.

Using the measurements from the stations listed in Table 2 the validation metrics in Equations (A2)–(A5) can be calculated from the hourly measured and estimated values. In the following sections, the MBD values are calculated considering both day and night values, as is common in the climatological community. This should be borne in mind when comparing these results with those reported by others. The relative statistics (rMBD, rMAB and rRMSD) will not be affected by this choice.

## 3.3. Validation Results, Global Horizontal Irradiance

The results of the validation of the global irradiance estimates are shown in Table 3.

**Table 3.** Validation of the satellite-based retrieval of global horizontal irradiance,  $G$ . The validation metrics have been calculated according to Equations (A2)–(A5). The values for average irradiance include also night hours.

Station	Year	$(G_{\text{meas}})$ $\text{Wm}^{-2}$	MBD $(\text{Wm}^{-2})$	rMBD $(\%)$	MAB $(\text{Wm}^{-2})$	rMAB $(\%)$	RMSD $(\text{Wm}^{-2})$	rRMSD $(\%)$
Meteosat PRIME								
De Aar	2002	257	1.54	0.60	27.42	10.67	66.13	25.73
Solar Village	2002	280	−7.16	−2.55	22.22	7.93	48.11	17.17
Cabauw	2005	137	−0.50	−0.37	28.31	20.68	62.19	45.44
Camborne	2005	130	−2.40	−1.85	26.85	20.71	58.26	44.93
Carpentras	2005	185	10.10	5.47	22.85	12.38	52.75	28.57
Florianopolis	2005	188	0.54	0.29	39.63	21.13	89.54	47.73
Ispra	2005	154	13.85	9.02	27.41	17.85	58.65	38.20
Kishinev	2005	149	2.05	1.37	30.78	20.59	61.89	41.40
Lindenberg	2005	133	−4.29	−3.22	27.26	20.49	62.59	47.04
Payerne	2005	151	0.89	0.58	29.69	19.61	64.49	42.60
Sde Boqer	2005	250	8.53	3.41	23.76	9.94	53.00	21.18
Tamanrasset	2005	266	6.98	2.63	26.56	10.00	64.96	24.45
Thessaloniki	2005	194	6.90	3.55	31.55	16.22	61.23	31.49
Toravere	2005	119	−4.94	−4.14	27.43	23.00	60.90	51.07
Meteosat EAST								
De Aar	2002	250	−3.02	−1.21	28.91	11.55	71.32	28.50
Solar Village	2002	273	−0.62	−0.23	20.50	7.51	48.92	17.91
Kishinev	2005	151	0.75	0.50	32.62	21.64	64.76	42.97
Sde Boqer	2005	251	9.89	3.94	27.25	10.84	60.07	23.90
Thessaloniki	2005	194	−7.05	−3.64	40.14	20.70	79.40	40.95
Xianghe	2005	174	1.46	0.84	44.35	25.48	96.90	55.68
Cocos Island	2009	242	1.38	0.57	39.25	16.25	82.25	34.05
Sde Boqer	2009	252	8.41	3.34	24.34	9.66	53.62	21.27

The MBD remains within  $\pm 10 \text{ Wm}^{-2}$  for all stations except two (Ispra and Carpentras). These two stations are also the only ones where rMBD is higher than 5%. rRMSD is much higher, ranging from 17% to more than 50% in Toravere and Xianghe. This result is common for satellite-based solar radiation estimates and reflects the difficulty of correctly estimating the precise position and timing of clouds. RMSD is generally higher in cloudy climates and at locations near the edge of the satellite images, such as Toravere and Xianghe.

From the values in Table 3 we can calculate the global statistical quantities in Equations (A7)–(A9). The average measured irradiance of all the stations is  $G_{\text{avg}} = 196.59 \text{ Wm}^{-2}$ , and the global bias deviation is  $\text{gMBD} = 1.63 \text{ Wm}^{-2}$ . The relative global bias is  $\text{rgMBD} = 0.73\%$ , while  $\text{rgMAB} = 2.36\%$  and  $\text{rgRMSD} = 3.18\%$ . For the calculation of these values, a station used to validate data from both satellites

is included with the results for both satellites, as if it were two different stations. This is the case of De Aar, Solar Village, Kishinev, Sde Boqer and Thessaloniki. However, the results for Sde Boqer for 2005 and 2009 considering Meteosat East have been averaged before calculating the aggregate statistics for all stations.

The mean bias deviation has also been calculated according to the type of sky, as described in Section 3.2. The results are shown in Table 4. For clear-sky conditions the method tends to underestimate the irradiance: at most of the stations, there is a negative bias. This bias tends to be larger near the edge of the satellite images (Toravere and Xianghe have the largest negative bias values), while the bias tends to be relatively low in desert areas (Sde Boqer, Solar Village and Tamanrasset). At the same time the irradiance under intermediate sky conditions tends to be overestimated, sometimes by as much as 25%. In overcast situations the overestimation of the irradiance is even more evident, with the extreme example of Xianghe where the estimate is 100% higher than the average measured value. This may be due to the fact that the station is close to the edge of the satellite field of view.

**Table 4.** Relative mean bias deviation of the global horizontal irradiance retrieval, according to sky type. In this table the average irradiance values (given in  $Wm^{-2}$ ) are calculated considering only daytime values.

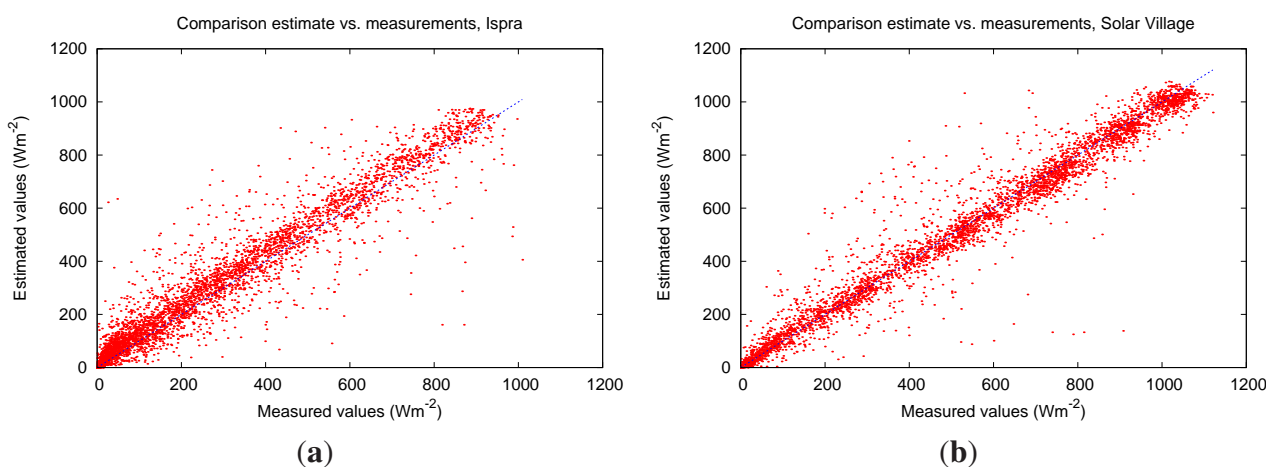
Station	Year	Clear		Intermediate		Overcast	
		$G_{meas}$	rMBD (%)	$G_{meas}$	rMBD (%)	$G_{meas}$	rMBD (%)
Meteosat PRIME							
De Aar	2002	604	1.82	293	−3.04	123	−41.27
Solar Village	2002	657	−0.64	240	−11.75	96	−57.38
Cabauw	2005	423	−6.76	243	2.33	86	34.28
Camborne	2005	448	−9.45	235	1.97	85	30.01
Carpentras	2005	476	1.28	231	20.38	83	56.40
Florianopolis	2005	611	−8.38	316	10.77	109	37.80
Ispra	2005	510	2.26	233	17.23	72	57.37
Kishinev	2005	535	−3.64	287	8.00	68	31.41
Lindenberg	2005	428	−7.90	225	0.33	80	28.47
Payerne	2005	492	−5.84	226	8.29	86	40.64
Sde Boqer	2005	637	1.17	212	18.70	81	32.24
Tamanrasset	2005	654	−1.03	273	23.50	95	63.07
Thessaloniki	2005	608	−1.43	310	17.34	72	34.43
Toravere	2005	382	−10.80	199	1.57	74	33.27
Meteosat EAST							
De Aar	2002	580	−3.94	322	15.98	101	61.84
Solar Village	2002	645	−2.18	193	21.87	90	56.92
Kishinev	2005	537	−4.60	288	8.61	68	27.54
Sde Boqer	2005	637	0.55	212	25.03	82	65.63
Thessaloniki	2005	609	−8.42	312	9.47	73	29.07
Xianghe	2005	503	−9.12	305	3.92	94	100.11
Cocos Island	2009	651	−7.16	353	14.14	130	55.53
Sde Boqer	2009	631	−0.17	216	25.80	89	73.82

These results indicate that SPECMAGIC, while representing well the average irradiance, tends to make estimates that give intermediate values rather than the extreme values that are measured under clear-sky or overcast situations.

A graphical comparison of the measured and estimated values is shown in Figure 1, for two stations: Solar Village in a desert climate and Ispra, Italy which is a humid temperate climate. The overall MBD for Solar Village is very low, while the Ispra location has the highest MBD of all stations. The pattern of results is rather different for the two stations. For Solar Village, the majority of points are below the diagonal (estimated is lower than measured), with a number of scattered points far above the diagonal. This is consistent with the finding that the estimates tend to be too low for clear-sky conditions and too high for more cloudy conditions. For Ispra, the estimated values are generally higher than measurements at all irradiance levels. This could be due to an underestimate of the effects of aerosols. It might also indicate a problem in the measurements.

As mentioned in the beginning of this section, all the validation data are available online. The interested reader may use these data to produce similar plots for any of the stations used in the validation exercise.

**Figure 1.** Two examples of scatterplots comparing measured and estimated values of  $G$ . (a) Station in Ispra, Italy; (b) Station at Solar Village, Saudi Arabia. The solid line represents the diagonal  $y = x$ .



### 3.4. Validation Results, Direct Normal Irradiance

The algorithm described in Section 2.1 can also estimate the direct component of the solar radiation, either as direct horizontal or direct normal irradiance.  $DNI$  is also measured at nearly all BSRN stations as well as at some of the non-BSRN stations used for the validation in the previous section.

Table 5 shows the results of the validation of  $DNI$  estimates from the SPECMAGIC model, using the stations that have  $DNI$  data available.

The relative global mean absolute bias (rgMAB) and relative global root mean square deviation (rgRMSD) have values  $\text{rgMAB} = 5.03\%$  and  $\text{rgRMSD} = 6.30\%$ . These values are significantly higher than the corresponding values for the global horizontal irradiance.

For most stations, the mean bias deviation value is higher than the one observed in the validation of the global irradiance. Also, the rMAB and rRMSD values are approximately twice as high as those of the global irradiance. However, when global statistics are analyzed we can see that the global MBD for all the stations and the relative global MBD are lower in the validation of the beam component than for the global irradiance. The higher deviations of the estimated beam irradiance, in some cases as overestimation and in other as underestimation, tend to compensate better than in the case of the global irradiance. The very low overall MBD for  $DNI$  should therefore be regarded as fortuitous.

As for the global irradiance estimates, we can calculate the global statistical quantities in Equations (A7)–(A9). The average measured DNI of all the stations is  $DNI_{avg} = 214.62 \text{ Wm}^{-2}$ , and the global mean bias deviation is  $gMBD = 0.61 \text{ Wm}^{-2}$ . The relative global bias is  $rgMBD = 0.62\%$ .

**Table 5.** Validation of the satellite-based retrieval of direct normal irradiance,  $DNI$ . The validation metrics have been calculated according to Equations (A2)–(A5). The values for average irradiance include also night hours.

Station	Year	$(DNI_{meas})$ ( $\text{Wm}^{-2}$ )	MBD ( $\text{Wm}^{-2}$ )	rMBD (%)	MAB ( $\text{Wm}^{-2}$ )	rMAB (%)	RMSD ( $\text{Wm}^{-2}$ )	rRMSD (%)
Meteosat PRIME								
De Aar	2002	330	−13.32	−4.03	72.43	21.93	146.64	44.40
Solar Village	2002	280	−10.18	−3.64	64.15	22.93	123.17	44.02
Cabauw	2005	116	10.32	8.89	52.51	45.24	120.28	103.63
Camborne	2005	112	0.92	0.82	49.75	44.56	117.75	105.48
Carpentras	2005	234	10.93	4.67	56.96	24.35	121.13	51.78
Florianopolis	2005	153	5.94	3.89	59.50	38.96	141.33	92.54
Kishinev	2005	152	−4.29	−2.83	53.90	35.01	124.75	82.27
Lindenberg	2005	143	0.02	0.01	52.06	36.30	120.52	84.02
Payerne	2005	165	3.03	1.87	63.90	39.50	147.06	90.90
Sde Boqer	2005	288	−11.23	−3.90	68.06	23.64	134.53	46.73
Tamanrasset	2005	269	30.58	11.35	75.32	27.96	144.47	53.62
Toravere	2005	133	−17.27	−12.94	60.16	45.09	142.49	106.80
Meteosat EAST								
De Aar	2002	321	−25.41	−7.91	71.97	22.40	143.74	44.74
Solar Village	2002	272	8.14	2.99	63.99	23.51	125.37	46.07
Kishinev	2005	152	−4.45	−2.94	56.37	37.16	128.01	84.40
Sde Boqer	2005	288	−8.67	−3.01	74.70	25.91	150.98	52.36
Xianghe	2005	112	4.36	3.87	65.96	58.64	141.14	125.46
Cocos Island	2009	191	22.94	12.02	71.60	37.51	148.28	77.68
Sde Boqer	2009	294	3.09	1.05	71.21	24.21	140.29	47.70

### 3.5. Comparison of Irradiance Probability Density

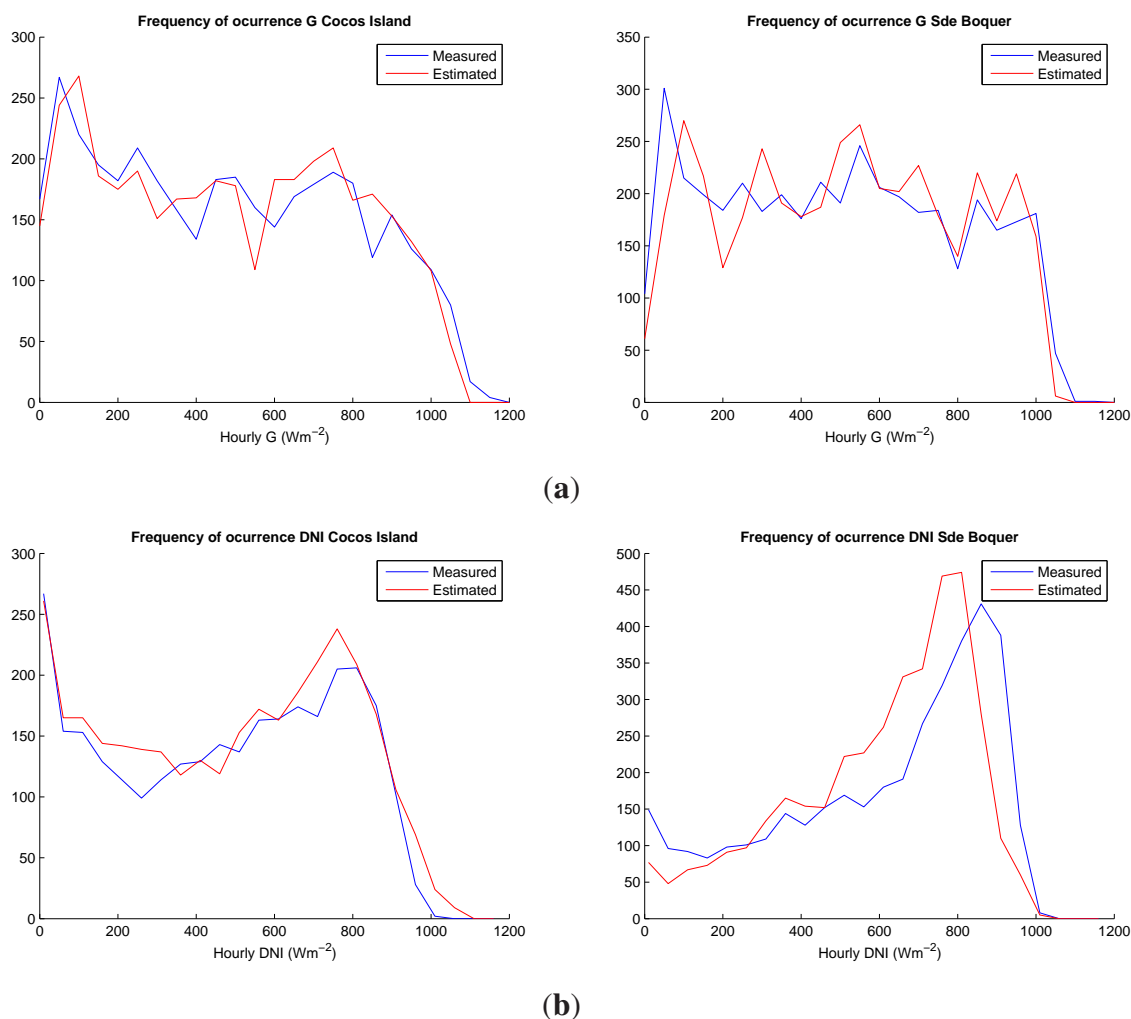
Besides the statistical measures (MBD, MAB, RMSD, *etc.*) used for the validation of the estimation model, it is interesting to analyze the probability distribution of the irradiance values, to see if the occurrence of the different irradiance levels are properly represented by the model.

For this analysis, the data when the sun’s elevation angle is below 5 degrees are not considered. The remaining data are distributed into bins of width  $50 \text{ Wm}^{-2}$  from 0 to  $1200 \text{ Wm}^{-2}$ . For the direct irradiance the  $0 \text{ Wm}^{-2}$  value will not be displayed, in order to avoid the high peak in the probability distribution at that level of irradiance.

As an example of the probability distribution graphs, those of the global and direct normal irradiances for Cocos Island and Sde Boquer are shown in Figure 2. These two stations have been chosen as they represent quite different climates.

For the global irradiance, in general, the model does not provide the very low and high irradiance values measured at the stations. For intermediate irradiance levels, there is no clear pattern. For the two stations shown in Figure 2, the overestimation derived from the model for Sde Boquer ( $+8.41 \text{ Wm}^{-2}$ ) is reflected in the graph where it is seen that the model has too many occurrences at high irradiance ( $800\text{--}900 \text{ Wm}^{-2}$ ) and too few at low irradiance ( $<100 \text{ Wm}^{-2}$ ). For Cocos Island there are no consistent differences in the intermediate irradiance range.

**Figure 2.** Comparison of irradiance probability density functions for two stations (Cocos Island and Sde Boquer). (a) Global horizontal irradiance; (b) Direct normal irradiance.



Regarding the frequency distribution of the beam normal irradiance values, the general trend for the various stations is the one observed in Sde Boqer. There is a higher number of observations of very high irradiance values that are not reached by the estimation model, while for the intermediate irradiance levels, the estimated values show a higher occurrence. However, it is the opposite case in Cocos Island, especially in the range of high beam irradiance values. This is reflected in the overall high positive bias (+22.94 Wm<sup>-2</sup>). This is likely to be due to an underestimation of the aerosol content in the area, leading to an overestimate of DNI under clear-sky conditions.

The lack of very high DNI values in the satellite estimates may be due to the fact that we are using long-term averages of AOD data. Periods with exceptionally low aerosol load will tend to have very high DNI values, which will be seen in the measurements but not captured by the satellite-based estimates.

The hourly estimated and measured data are available in the online supplementary material. This allows interested readers to construct these graphs for any station in the validation data set.

## 4. Results and Discussion

The methods described in Section 2 have been used to calculate solar radiation parameters for every full-hour time slot in the period 1999–2013, with a hiatus in 2006 due to uncertainty about the quality of data from Meteosat-5 towards the end of the operational life of that satellite. Thus, a total of 14 years are available, which makes it possible to calculate long-term averages with low uncertainty from interannual variation.

### 4.1. Sample Results

An example of a map of solar irradiance (global horizontal and direct normal) is shown in Figure 3. The time slot is 09:00GMT on 14 March 2005, showing the maximum irradiance, corresponding to cloudless conditions, just west of the center of the image (which is at 63°E), and illustrating the overall extent of valid data.

### 4.2. Comparison of Meteosat Prime and East Estimates

From 1998 to 2006 two very similar satellites were both in orbit at the same time at widely different positions (Meteosat 7 at 0° and Meteosat 5 at 63°E). This makes it possible to do a direct comparison of the solar radiation retrieval from the two satellites in the area where they overlap.

Figure 5 shows a map of the relative difference between  $H$  for the year 2005 calculated from the two satellites, defined as:

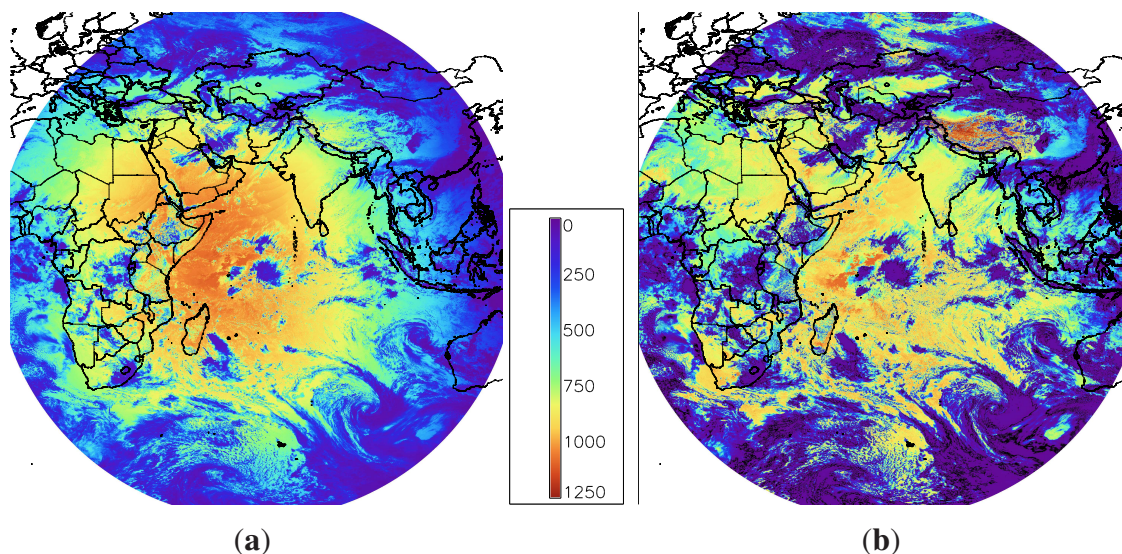
$$\Delta H = \frac{H_{\text{prime}} - H_{\text{east}}}{H_{\text{prime}}} \quad (7)$$

The result shows a significant gradient in the east-west direction, with estimates from Meteosat Prime higher than those of Meteosat East in the western part and vice versa in the eastern part. This means that the retrieval algorithm tends to give lower values near the edge of the satellite image.

The effects of aerosols and water vapour are calculated by the clear-sky radiation model and are therefore independent of the viewing angle of the satellite, so the differences shown in Figure 5 must in some way be due to the effect of clouds, either in the calculation of  $CAL$  or in the treatment of clouds by

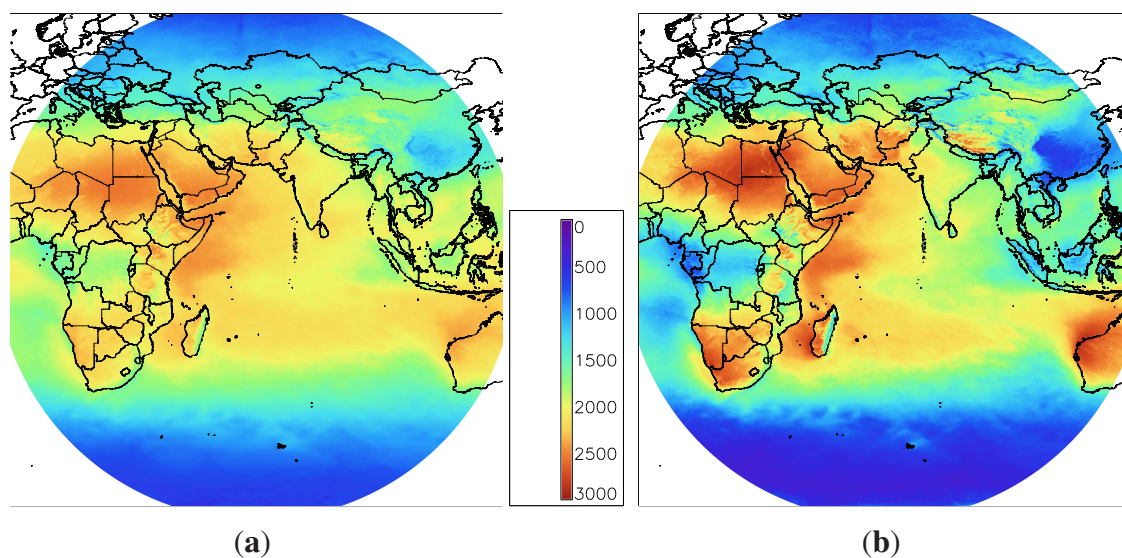
the SPECMAGIC algorithm. A possible explanation is that due to the finite thickness of clouds, broken cloud fields can be identified when viewed directly from above (close to the satellite nadir) but will appear as continuous cloud cover when viewed from a shallow angle (as is the case near the edge of the image). This would lead to an overestimation of cloudiness and hence an underestimation of radiation.

**Figure 3.** Example of solar irradiance from the calculations. The slot shown is 14 March 2005 at 09:00GMT. (a) Global horizontal irradiance  $G$ ; (b) Direct normal irradiance  $DNI$ . The values are in  $Wm^{-2}$ .



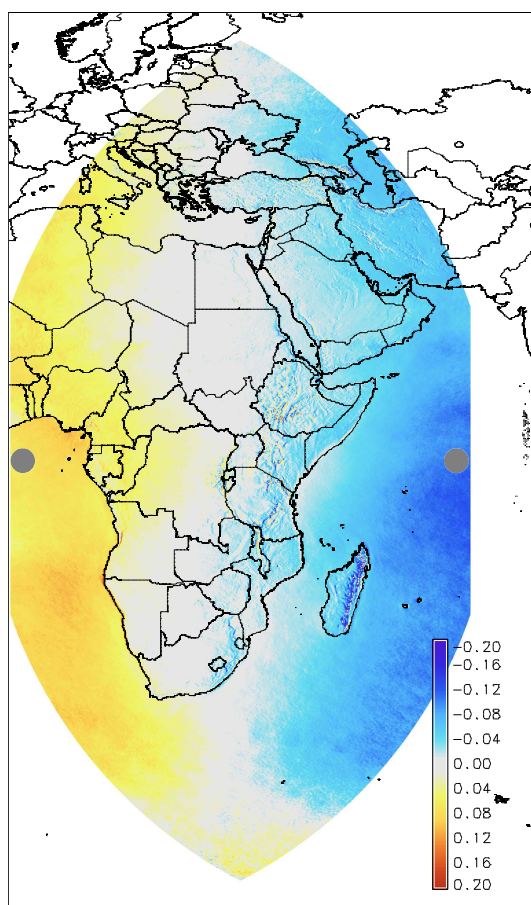
The individual hourly slots can be summed to produce maps of the global horizontal and direct normal irradiation for longer time periods. Figure 4 shows  $H$  and  $H_{DNI}$  for the year 2002.

**Figure 4.** Yearly solar irradiation for the year 2002. (a) Global horizontal irradiation  $H$ ; (b) Direct normal irradiation,  $H_{DNI}$ . The values are in  $kWh m^{-2}$ .





**Figure 5.** Relative difference  $\Delta H$  in yearly global horizontal irradiation, calculated from the two satellites for the year 2005 using Equation (7).



In the validation results shown in Table 3 some of the stations were visible from both satellites at the same time. For these stations the same trend is seen. De Aar and Thessaloniki are both closer in longitude to the Meteosat Prime nadir, and their MBD is more positive for Meteosat Prime (at De Aar, rMBD is 0.60% for Meteosat Prime vs.  $-1.21\%$  for Meteosat East, while for Thessaloniki the values are  $+3.55\%$  vs.  $-3.64\%$ ). Solar Village is closer to Meteosat East and has rMBD of  $-2.55\%$  for Meteosat Prime and  $-0.23\%$  for Meteosat East. Sde Boqer and Kishinev are close being at equal distance to the two satellites, and here we see rather little difference in rMBD ( $3.41\%$  vs.  $3.94\%$  for Sde Boqer,  $1.37\%$  vs.  $0.50\%$  for Kishinev).

An analogous case is evident in mountain areas which often have very localized gradients in the available solar radiation. In some cases the local patterns are shifted relative to each other in the data sets obtained from the two different satellites. This may be due to the effect of parallax when the two satellites are observing persistent cloud systems over mountains from different angles. This effect will lead to an increased uncertainty in mountainous regions.

#### 4.3. Comparing the Results with Existing Databases

In the scientific bibliography there are several publications in which solar radiation estimates retrieved from satellite images are validated by comparison with ground measurements. The validation is

commonly performed on daily or monthly average irradiance values, although in some cases hourly irradiance values are analyzed, like in this paper.

Two examples of hourly validation data were performed by Ineichen (2011 and 2013) [35,36]. The author compared the irradiance values derived from various satellite products with the measurements registered in several stations. It should be noted that these data sets are derived from the Meteosat MSG class of satellites, so the methods employed are somewhat different. Nevertheless it is possible to get an idea of how the accuracy of the present data set compares with other recent efforts.

A detailed description of the method we used to compare these results is found in the supplementary material [28].

Data of global horizontal and beam normal irradiance are validated in hourly values. The magnitude of the errors is similar of those reported in this paper.

In the first report [35], five different satellite products are validated: SolarGis, Helioclim 3 data bank, 3Tier algorithm, EnMetSol considering two different clear sky models (denoted Solis and Dumortier) and IrSolAv estimates. The estimates are validated with irradiance measurements registered at 20 stations in the case of the global horizontal irradiance and 12 stations for the beam component. The validation results for  $G$  and  $DNI$  are shown in Table 6.

**Table 6.** General statistics for the validation of global horizontal and direct normal irradiance estimation, Ineichen, 2011 [35]. The results for the present work are given as the last column for comparison.  $N_s$  is the number of stations used for the comparison for each product.

	SolarGis	Heliosat 3	3Tier	EnMetSol (solis)	EnMetSol (dumor)	IrSolAv	SPECMAGIC
Global horizontal irradiance							
$N_s$	20	20	20	20	20	15	16
rgMBD (%)	0.75	1.90	1.20	0.95	0.95	0.33	0.73
rgMAB (%)	2.15	4.30	3.00	3.05	3.25	2.87	2.36
rgRMSD (%)	2.69	5.54	3.59	3.57	3.75	3.75	3.18
Direct normal irradiance							
$N_s$	12	12	12	12	12	7	14
rgMBD (%)	-2.50	10.83	8.08	-1.75	6.25	-0.29	0.62
rgMAB (%)	5.33	13.00	10.92	5.58	9.75	2.86	5.03
rgRMSD (%)	6.18	18.67	14.13	6.84	10.97	3.96	6.30

The magnitude of the general statistics are similar to those obtained in the present paper. The validation of the SPECMAGIC estimates reports rgMBD, rgMAB and rgRMSD values as low as the best satellite product analyzed in [35], for both  $G$  and  $DNI$ .

If the validation of the global irradiance on the horizontal plane is performed distributing the data according to sky type, we obtain the results shown in Table 7. Again, the results are similar to those derived from the SPECMAGIC method.

For all satellite products considered, including SPECMAGIC estimates, the global horizontal irradiance is underestimated in clear sky situations, while for intermediate and more significantly overcast situations, the solar resource is overestimated.

It should be noted that although the number of stations used in [35] is similar to that used in the present study, there is not a great overlap in the actual list of stations. In fact, only 8 stations used in [35] are the same as those we have used. This may skew the results because some station locations may be more “challenging” than others for the algorithms. We have therefore repeated the comparison using these 8 stations: Cabauw, Camborne, Carpentras, Payerne, Sde Boqer, Thessaloniki, Toravere (Thessaloniki is used for the DNI comparison). IrSolAv was not validated for all these stations and therefore has been excluded from this comparison. These results are given in Table 8.

**Table 7.** General statistics for the validation of global horizontal irradiance estimation for clear, intermediate and overcast sky situations’ [35]. The results for the present work are given as the last column for comparison.

Clear Sky Conditions							
	SolarGis	Heliosat 3	3Tier	EnMetSol (solis)	EnMetSol (dumor)	IrSolAv	SPECMAGIC
rgMBD (%)	−2.75	−5.25	−4.35	−3.70	−3.70	−5.40	−4.03
rgMAB (%)	2.95	5.35	4.55	4.10	4.00	5.67	4.68
rgRMSD (%)	3.88	6.11	5.37	4.80	4.86	6.16	5.75
Intermediate Sky Conditions							
rgMBD (%)	5.05	10.85	4.65	8.00	7.90	7.87	10.24
rgMAB (%)	6.05	11.15	4.85	8.40	8.40	7.87	11.65
rgRMSD (%)	7.37	15.12	6.91	10.44	10.40	10.39	13.94
Overcast Sky Conditions							
rgMBD (%)	24.80	62.05	55.85	35.55	34.55	49.33	37.21
rgMAB (%)	24.80	62.05	55.85	35.55	34.55	49.33	46.61
rgRMSD (%)	28.19	77.47	59.36	39.96	39.49	55.21	49.94

**Table 8.** General statistics for the validation of global horizontal and direct normal irradiance estimation, Ineichen, 2011 [35], using only the 8 stations that all validations have in common.

	SolarGis	Heliosat 3	3Tier	EnMetSol (solis)	EnMetSol (dumor)	SPECMAGIC
Global horizontal irradiance						
rgMBD (%)	−0.38	−0.75	−0.25	−0.13	−0.13	0.93
rgMAB (%)	1.13	3.50	1.75	2.13	2.63	2.93
rgRMSD (%)	1.46	3.97	2.12	2.26	2.98	3.30
Direct normal irradiance						
rgMBD (%)	−3.43	8.57	5.29	−2.00	6.57	1.22
rgMAB (%)	6.00	12.29	8.14	6.86	10.00	5.81
rgRMSD (%)	6.46	15.02	9.85	8.26	11.56	7.26

Given the low number of stations, it is difficult to clearly distinguish between the values for the different products. It is possible to do a test for statistical significance of the differences in the global variances (the square of rgRMSD). A two-tailed f-test statistic ( $7 \times 7$  degrees of freedom)

on the values for  $G$  in Table 8 shows that only SolarGIS is significantly better than the results from SPECMAGIC ( $p < 0.04$ ). For the  $DNI$  comparison (7 stations), none of the rgRMSD results from [35] are significantly different from the SPECMAGIC result.

Similarly, for the results in Table 6, few of the differences in rgRMSD between SPECMAGIC and the other products are statistically significant: only the differences with Heliosat 3 ( $G$  and  $DNI$ ) and 3Tier ( $DNI$ ) are statistically significant at  $p < 0.05$ . As noted previously, this may be in part due to the difference in choice of stations.

In the second report [36], six different satellite products are validated: SolarGis, Helioclim 3, Solemi, Heliomont (Meteoswiss), EnMetSol, IrSolAv. Estimates are compared with global horizontal and beam normal irradiance experimental data registered at 17 stations. Results are shown in Table 9. Again, the range of the statistics is similar to that observed in the present work. In this case, the number of stations in common is even smaller, and a comparison for these few stations does not make statistical sense.

**Table 9.** General statistics for the validation of global horizontal and direct normal irradiance estimation, Ineichen, 2013 [36]. The results for the present work are given as the last column for comparison.

	SolarGis	HelioClim 3	Solemi	Heliomont	EnMetSol	IrSolAv	SPECMAGIC
Global horizontal irradiance							
$N_s$	17	17	17	17	17	15	16
rgMBD (%)	0.00	1.06	2.29	0.71	−0.59	0.47	0.73
rgMAB (%)	1.65	3.41	3.47	2.94	2.71	3.27	2.36
rgRMSD (%)	2.14	4.38	3.84	3.45	3.22	5.39	3.18
Direct normal irradiance							
$N_s$	17	17	17	17	17	15	16
rgMBD (%)	−2.00	7.94	−12.94	0.18	1.35	4.47	0.62
rgMAB (%)	5.53	10.53	13.53	8.76	7.71	8.60	5.03
rgRMSD (%)	7.43	16.30	16.19	10.40	8.69	12.02	6.30

## 5. Conclusions

We have presented the results of applying the Heliosat and SPECMAGIC algorithms to images from the Meteosat Eastern geostationary satellites in order to produce hourly maps of surface solar irradiance (global and direct). For validation purposes, estimates from the Meteosat Prime satellites were also produced. Validation from 16 ground stations show a low overall yearly bias (gMBD) in the estimation of the global horizontal irradiance of  $+1.63 \text{ Wm}^{-2}$  or  $+0.73\%$ . The mean absolute bias of the individual station bias values (rgMAB) is  $2.36\%$  while the root-mean-square of the station bias values (rgRMSD) is  $3.18\%$ . For the annual average DNI estimation, the corresponding values are:  $\text{gMBD} = +0.61 \text{ Wm}^{-2}$  or  $+0.62\%$ ,  $\text{rgMAB} = 5.03\%$ , and  $\text{rgRMSD} = 6.30\%$ . The very low gMBD value is a fortuitous cancelling of individual deviations. The two latter values give a better idea of the uncertainty in estimates at individual locations. A comparison with validation results given by Ineichen [35,36] shows that the accuracy of the present results is similar to that of other recent works on satellite-based solar radiation databases.

The overall deviations from measured data are quite low. However, there are still areas where improvements could be made. A direct comparison between solar radiation estimates from the Meteosat Prime and East over the same time period shows that the radiation estimates tend to be higher near the center of the satellite image than near the edge. This is probably due to the fact that the finite thickness of clouds will make the cloud cover seem more extensive when viewed at a shallow angle as is the case near the edge of the satellite image. To resolve this issue would require more information about cloud type than is currently available in the method.

For the time being, there is no particular treatment of snow cover in the method, so areas affected by snow may have larger uncertainties. Since none of the validation stations are in high mountain or arctic areas, this problem is not seen in the validation results but should be acknowledged anyway. Methods to account for snow cover are being studied at present and will be applied to future versions of the solar radiation database.

For the present study, long-term monthly averages of aerosol data have been employed, as described in Section 2.2. This choice may have an influence on the year-to-year variability of the total solar radiation, which is of interest to investors and operators of solar energy power plants. We plan to investigate the influence of using monthly aerosol data on the resulting solar radiation time series.

Once processing has been completed, the hourly maps of solar radiation will be made freely available to the public. To our knowledge, this new data set is the first freely available high-resolution solar radiation data set covering Asia (there exist several for Europe and Africa, and free data are to some extent also available for the Americas). The new data set should be useful for researchers, planners and installers of solar energy systems, as well as for studies of climate and biological processes.

The SPECMAGIC method can also calculate spectrally resolved solar irradiance. We plan to use this feature to produce maps of spectrally resolved irradiance to study the effects of spectral changes on PV system productivity for the area covered by the data. Maps of photosynthetically active radiation may also be produced.

Solar radiation data are essential for the estimation of solar energy system performance. An important application of the data presented in this study will be to perform estimates of solar energy system performance. In particular, it is planned that these data will be incorporated into the web-based PV estimation tool PVGIS [37,38].

## Acknowledgments

We would like to thank EUMETSAT for the access to satellite image data and the CM SAF Project for the possibility to use the algorithms developed within that project.

## Author Contributions

The contributions of each author to the work described in this paper are as follows:

**Ana Gracia Amillo** did most of the work on validation of satellite data against ground station data and the comparison with other recent work and wrote the section on validation and parts of the results & discussion.

**Thomas Huld** performed the data processing for Meteosat East and some of the Meteosat Prime data, as well as postprocessing to produce annual means and differences between satellites (Figures 3–5). He wrote most of the introduction and conclusions and part of the discussion.

**Richard Müller** produced an updated version of the SPECMAGIC code and processed the climatological data needed as input. He wrote most of the section on methods.

## Appendix

### Statistical Measures Used for the Validation

A number of different statistical measures have been used to estimate the uncertainty in the irradiance retrieved from satellite data. These are defined briefly here.

For a given set of data (station or satellite-based estimate), the annual average is calculated from the  $N$  values:

$$\langle G \rangle = \frac{1}{N} \sum_{i=1}^N (G_i) \quad (\text{A1})$$

The *Mean Bias Deviation* (MBD) is calculated as shown in Equation (A2):

$$\text{MBD} = \frac{1}{N} \sum_{i=1}^N (G_i^r - G_i^m) \quad (\text{A2})$$

Here  $G_i^m$  is the measured irradiance at the  $i^{\text{th}}$  time point while  $G_i^r$  is the estimated irradiance at the same time point. In this calculation both day and night time slots are considered.  $N$  is the total number of time points for which both measured and estimated data are available.

The *relative MBD* (rMBD) is defined in Equation (A3) as:

$$\text{rMBD} = \frac{\sum_{i=1}^N (G_i^r - G_i^m)}{\sum_{i=1}^N (G_i^m)} = \frac{\text{MBD}}{\langle G \rangle} \quad (\text{A3})$$

The *Mean Absolute Bias* (MAB) is calculated as (Equation (A4)):

$$\text{MAB} = \frac{1}{N} \sum_{i=1}^N |G_i^r - G_i^m| \quad (\text{A4})$$

The *Root Mean Square Deviation* (RMSD) is given as:

$$\text{RMSD} = \sqrt{\frac{\sum_{i=1}^N (G_i^r - G_i^m)^2}{N}} \quad (\text{A5})$$

Relative values of MAB and RMSD are calculated like Equation (A3):  $\text{rMAB} = \text{MAB} / \langle G \rangle$  and  $\text{rRMSD} = \text{RMSD} / \langle G \rangle$

To get an overall estimate of the bias in yearly average solar irradiation, we can calculate the average of the MBD values at all the station locations. This is termed the *global* MBD (and respectively, relative global MBD):

$$g\text{MBD} = \frac{1}{N_s} \sum_{s=1}^{N_s} \text{MBD}_s \quad (\text{A6})$$

$$rg\text{MBD} = \frac{1}{N_s} \sum_{s=1}^{N_s} r\text{MBD}_s \quad (\text{A7})$$

Here,  $\text{MBD}_s$  is the MBD of station number  $s$ .  $N_s$  is the number of stations.

A measure of the uncertainty of the yearly average irradiation in a given point can be given as the *relative global* MAB and RMSD:

$$rg\text{MAB} = \frac{1}{N_s} \sum_{s=1}^{N_s} |r\text{MBD}_s| \quad (\text{A8})$$

$$rg\text{RMSD} = \sqrt{\frac{\sum_{s=1}^{N_s} (r\text{MBD}_s)^2}{N_s}} \quad (\text{A9})$$

## Conflicts of Interest

The authors declare no conflict of interest.

## References

1. Möser, W.; Raschke, E. Incident solar radiation over Europe estimated from METEOSAT data. *J. Clim. Appl. Meteor* **1984**, *23*, 166–170.
2. Cano, D.; Monget, J.; Albuissou, M.; Guillard, H.; Regas, N.; Wald, L. A method for the determination of the global solar radiation from meteorological satellite data. *Solar Energy* **1986**, *37*, 31–39.
3. Pinker, R.; Laszlo, I. Modelling surface solar irradiance for satellite applications on a global scale. *J. Appl. Meteorol.* **1992**, *31*, 166–170.
4. Meyer, R.; Hoyer, C.; Schillings, C.; Trieb, F.; Diedrich, E.; Schroedter, M. SOLEMI: A new satellite-based service for high-resolution and precision solar radiation data for Europe, Africa and Asia. In Proceedings of the 2003 ISES Solar World Congress, Gäddede, Sweden, 14–19 June 2003.
5. Müller, R.; Matsoukas, C.; Gratzki, A.; Behr, H.; Hollmann, R. The CM-SAF operational scheme for the satellite based retrieval of solar surface irradiance—A LUT based eigenvector hybrid approach. *Remote Sens. Environ.* **2009**, *113*, 1012–1024.
6. Rigollier, C.; Lefevre, M.; Wald, L. The method Heliosat-2 for deriving shortwave solar radiation from satellite images. *Solar Energy* **2004**, *77*, 159–169.
7. Laszlo, I.; Ciren, P.; Liu, H.; Kondragunta, S.; Tarpley, J.D.; Goldberg, M.D. Remote sensing of aerosol and radiation from geostationary satellites. *Adv. Space Res.* **2008**, *41*, 1882–1893.

8. Karlsson, K.-G.; Riihelä, A.; Müller, R.; Meirink, J.F.; Sedlar, J.; Stengel, M.; Lockhoff, M.; Trentmann, J.; Kaspar, F.; Hollmann, R.; *et al.* CLARA-A1: The CM SAF cloud, albedo and radiation dataset from 28 yr of global AVHRR data. *Atmos. Chem. Phys. Discuss.* **2013**, *13*, 935–982.
9. SOLEMI. Available Online: <http://www.dlr.de/tt/solemi> (accessed on 26 August 2014).
10. Blanc, P.; Gschwind, B.; Lefèvre, M.; Wald, L. The HelioClim project: Surface solar irradiance data for climate applications. *Remote Sens.* **2011**, *3*, 343–361, doi:10.3390/rs3020343.
11. CM SAF. Available Online: <http://www.cmsaf.eu> (accessed on 26 August 2014).
12. Geiger, B.; C. Meurey, D.; Lajas, L.; Franchistéguy, L.; Carrer, D.; Roujean, J.-L. Near real-time provision of downwelling shortwave radiation estimates derived from satellite observations. *Meteorol. Appl.* **2008**, *15*, 411–420.
13. LSA SAF. Available Online: <http://landsaf.meteo.pt> (accessed on 26 August 2014).
14. Müller, R.; Behrendt, T.; Hammer, A.; Kemper, A. A new algorithm for the satellite-based retrieval of solar surface irradiance in spectral bands. *Remote Sens.* **2012**, *4*, 622–647.
15. Müller, R.; Trentmann, J.; Träger-Chatterjee, C.; Posselt, R.; Stöckli, R. The Role of the Effective Cloud Albedo for Climate Monitoring and Analysis. *Remote Sens.* **2011**, *3*, 2305–2320.
16. Hammer, A.; Heinemann, D.; Hoyer, C.; Kuhlemann, R.; Lorenz, E.; Müller, R.; Beyer, H. Solar energy assessment using remote sensing technologies. *Remote Sens. Environ.* **2003**, *86*, 423–432.
17. Posselt, R.; Müller, R.; Stöckli, R.; Trentmann, J. Spatial and temporal homogeneity of solar surface irradiance across satellite generations. *Remote Sens.* **2011**, *3*, 1029–1046.
18. Skartveit, A.; Olseth, J.; Tuft, M. An hourly diffuse fraction model with correction for variability and surface albedo. *Solar Energy* **1998**, *63*, 173–183.
19. Dee, D.P.; Uppala, S.M.; Simmons, A.J.; Berrisford, P.; Poli, P.; Kobayashi, S.; Andrae, U.; Balmaseda, M.A.; Balsamo, G.; Bauer, P.; *et al.* The ERA-Interim reanalysis: Configuration and performance of the data assimilation system. *Q. J. R. Meteorol. Soc.* **2011**, *137*, 553–597.
20. ECMWF. Available Online: <http://www.ecmwf.int> (accessed on 26 August 2014).
21. Morcrette, J.-J.; Boucher, O.; Jones, L.; Salmond, D.; Bechtold, P.; Beljaars, A.; Benedetti, A.; Bonet, A.; Kaiser, J.W.; Razinger, M.; *et al.* Aerosol analysis and forecast in the European Centre for Medium-Range Weather Forecasts Integrated Forecast System: Forward modeling. *J. Geophys. Res.* **2009**, *114*, doi:10.1029/2008JD011235.
22. Benedetti, A.; Morcrette, J.-J.; Boucher, O.; Dethof, A.; Engelen, R.J.; Fisher, M.; Flentje, H.; Huneeus, N.; Jones, L.; Kaiser, J.W.; *et al.* Aerosol analysis and forecast in the European Centre for Medium-Range Weather Forecasts Integrated Forecast System: 2. Data assimilation. *J. Geophys. Res.* **2009**, *114*, doi:10.1029/2008JD011115.
23. Inness, A.; Baier, F.; Benedetti, A.; Bouarar, I.; Chabrillat, S.; Clark, H.; Clerbaux, C.; Coheur, P.; Engelen, R.J.; Errera, Q.; *et al.* The MACC reanalysis: An 8 yr data set of atmospheric composition. *Atmos. Chem. Phys.* **2013**, *13*, 4073–4109.
24. MACC. Available Online: <http://www.gmes-atmosphere.eu> (accessed on 26 August 2014).
25. Brown, J.F.; Loveland, T.; Merchant, J.W.; Reed, B.C.; Ohlen, D.O. Using multi-source data in global land-cover characterization: Concepts, requirements, and methods. *Photogramm. Eng. Remote Sens.* **1983**, *59*, 977–987.



26. Dickinson, R.E. Land surface processes and climate—Surface albedos and energy balance. *Adv. Geophys.* **1983**, *25*, 305–353.
27. EUMETSAT. Available Online: <http://www.eumetsat.int> (accessed on 26 August 2014).
28. A New Database of Global and Direct Solar Radiation Using the Eastern Meteosat Satellite, Models and Validation. Available Online: <http://re.jrc.ec.europa.eu/esti/meteast/index.html> (accessed on 26 August 2014).
29. Baseline Surface Radiation Network. Available Online: <http://bsrn.awi.de> (accessed on 26 August 2014).
30. Global Atmosphere Watch. Available Online: <http://www.wmo.int/gaw/> (accessed on 26 August 2014).
31. World Radiation Data Centre. Available Online: <http://wrdc.mgo.rssi.ru> (accessed on 26 August 2014).
32. Ineichen, P.; Barroso, C.S.; Geiger, B.; Hollmann, R.; Marsouin, A.; Müller, R. Satellite Application Facilities irradiance products: Hourly time step comparison and validation over Europe. *Int. J. Remote Sens.* **2009**, *30*, 5549–5571.
33. Perez, R.; Ineichen, P.; Seals, R.; Zelenka, A. Making full use of the clearness index for parametrizing hourly insolation conditions. *Solar Energy* **1990**, *45*, 111–114.
34. Kasten, F. A simple parameterization of two pyr heliometric formulae for determining the Linke turbidity factor. *Meteorol. Rdsch.* **1980**, *33*, 124–127.
35. Ineichen, P. Five Satellite Products Deriving Beam and Global Irradiance Validation on Data from 23 Ground Stations. Available online: <http://archive-ouverte.unige.ch/unige:23669> (accessed on 26 August 2014).
36. Ineichen P. Long Term Satellite Hourly, Daily and Monthly Global, Beam and Diffuse Irradiance Validation. Interannual Variability Analysis. Available online: <http://archive-ouverte.unige.ch/unige:29606> (accessed on 26 August 2014).
37. Huld, T.; Müller, R.; Gambardella, A. A new solar radiation database for estimating PV performance in Europe and Africa. *Solar Energy* **2012**, *86*, 1803–1815.
38. PVGIS. Available Online: <http://re.jrc.ec.europa.eu/pvgis/> (accessed on 26 August 2014).

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