

PLANETARY TRANSMISSION LOAD SHARING: MANUFACTURING ERRORS AND SYSTEM CONFIGURATION STUDY

M. Iglesias, A. Fernandez del Rincon, A. de-Juan,
P. Garcia, A. Diez-Ibarbia, F. Viadero

*Department of Structural and Mechanical Engineering, University of Cantabria. Avda.
de los Castros s/n 39005 Santander, Spain.*

Abstract

This paper addresses the effect of manufacturing errors such as eccentricity and planet pin positioning errors on the quasi-static behavior of a 3 planet planetary transmission, taking into account different configurations regarding the bearing condition of the sun gear shaft. The aim of the paper is to shed light on some untouched aspects of the load sharing behavior of planetary transmissions, such as the effect of radial positioning errors of the planets when different pressure angles are used, and the impact of the different loadings per planet on the actual load per tooth.

A modeling approach is employed, and physical explanations and simplified graphs are provided to help understand the behavior of the transmission when the sun is allowed to float and errors are introduced. The model used, developed by the authors and presented and validated in previous works, hybridizes analytical solutions with finite element models in order to compute the contact forces.

The results obtained show that the teeth loads are much lower than expected compared to the planet uneven loads, both in the non-defected and defected transmission, and that radial positioning errors have non-negligible effect on the load sharing ratio under certain operating conditions.

Keywords: Gears, Planetary transmissions, Load sharing ratio, Transmission error, Run-out, Manufacturing errors

1. Introduction

One of the main advantages of planetary transmissions is its compactness. For high torques, instead of enlarging the wheels and thus its load capacity, planetary transmissions split the load into a number of paths. In this manner, the power is divided among several pinions, so that loadings per unit facewidth remain below nominal values while the torque is multiplied. Besides, planetary transmissions present coaxial input and output and large reduction ratios, being the most compact and lightest possible drives [1]. Under ideal conditions, each path in a planetary transmission carries an equal amount of load. Nevertheless, as in real systems there are inevitable manufacturing deviations due to errors and tolerances, the load is not equally shared amongst the different sun/planet/ring meshes, which can be a problem in terms of both durability (higher loadings per unit facewidth than expected) and dynamic behaviour (vibrations due to changing loads, etc).

The load sharing problem in planetary transmissions has been discussed in a number of publications, assessed by means of experimental tests [2, 3], but mainly based on transmission modeling, from simpler analytical models [4] to more complex models including hybrid semi-analytical and finite element techniques [5].

Due to its spacial configuration, planetary transmissions are complicated to model, but the critical importance of these gear systems in aerospace and energy generation applications makes the effort worth it. The main feature that characterizes the dynamic behavior of gear transmissions is the change in the number of teeth couples simultaneously in mesh. The meshing stiffness is therefore variable, and induces a periodic excitation in the system. Thus, the characterization of this periodic excitation is crucial in order to achieve better simulated results [6]. In a first step to increase modeling realism, the static transmission error has been used as excitation to predict dynamic behavior of planetary transmissions [7, 8, 9]. Nevertheless, more recent studies point that this approach, whilst remaining relatively valid for ordinary transmissions, may not be applicable to multi-mesh transmissions such as planetary ones [10]. With a higher degree of accuracy, at a second step evolution, there are gear models with time-varying stiffness. They give better off-resonance responses, but they are also used to identify regions of large amplitude vibration near resonances, where damping and other non-linear phenomena strongly affect the behavior [11, 12, 13]. The latest and more advanced planetary transmission models are those based on computa-

38 tional approaches, frequently including FEM techniques in combination with
39 different contact models [14]. In some cases, completely flexible bodies are
40 considered in real time simulation [15]. Depending on the particular applica-
41 tion of the model, different emphasize is given to each modeling aspect [16],
42 as is the case of non-stationary operation [17].

43 Studies on load sharing have usually been focus on the behavior of the
44 transmission when defects are present, evaluating the effect of different con-
45 figurations on the resulting load sharing, trying to find methods of improve-
46 ment. As latest works, in [18] the effects of gravity, ring support stiffness and
47 bedplate tilt angle of a wind turbine on the load sharing is studied through
48 modeling approaches. In [19] and again in the wind turbine field the load
49 sharing behaviour of a compound planetary gear transmission in presence of
50 multiple-errors is analysed, adding experimental results to verify the model
51 approach. The effect of floating the sun gear in a planetary gearbox has
52 been studied by [20], in order to absorb the consequences of geometrical
53 imperfections.

54 In this paper, a planetary model is used to study the load sharing in
55 quasi-static conditions, with the aim of shedding light on some untouched
56 aspects of the load sharing behaviour of 3 planet planetary transmissions,
57 such as the effect of radial positioning errors of the planets when different
58 pressure angles are used and the impact of the different loadings per planet
59 on the actual load per tooth. Specially in this last case, the new information
60 can improve the understanding of the tooth load per unit length when uneven
61 LSR occurs, and therefore to produce better gear design processes. Although
62 this new design insight is a direct consequence of the study carried out, the
63 ultimate goal of the planetary gear modeling research presented here is the
64 accurate reproduction of the transmission behaviour in real conditions for
65 on-condition monitoring assessment. The model used hybridizes analytical
66 solutions with finite element models in order to compute the contact forces,
67 making unnecessary the use of mesh stiffness waveform approximations or
68 static transmission error excitation assumptions. The mesh model is based
69 on previous work by the authors [21, 22], extended and improved towards the
70 planetary modeling as it can be found in [23]. Coupling through gear body
71 deformations is also given a special attention, due to the multiple meshes
72 per wheel. With respect to the contact point location and geometric overlap
73 modeling approach used in this work, it has been conceived to allow for the
74 almost direct inclusion of additional modeling features, such as tooth profile
75 modifications (with an approach used in [24]) or the use of shifted gears [25].

76 2. Planet load sharing

77 There are many variations of planetary gear trains. However, whether
78 simple or compound, with straight or helical gears, the vast majority of plan-
79 etary transmissions designs share a fundamental quality: their compactness.
80 This compactness can be understood in two different ways. The first has
81 to do with the kinematic configuration of the planetary transmission (with a
82 rotary planet carrier), which provides much higher ratios than those provided
83 by conventional transmissions. Additionally, this configuration allows coaxial
84 inputs and outputs, which is a plus for many applications, also economizing
85 space. The second reason for which a planetary transmission can be seen
86 as compact is the load capacity. The load capacity of a gear is ultimately
87 determined by the size of their teeth, so that, in general, a large workload nec-
88 essarily implies large gears accordingly. However, as planetary transmissions
89 divide the total load on a variable number of paths (sun-planet and planet-
90 ring pairs), the size of the gears can be reduced in the same proportion as
91 the number of load paths used with respect to an ordinary transmission.

92 Ideally, each of the planetary load paths should transmit the same fraction
93 of the total transmitted load. However, there are a number of reasons for
94 which the load distribution may not be even in the actual operation of the
95 planetary transmissions, the main two are the different path stiffness and
96 the errors in the manufacturing and assembly process. Thus, there will be
97 fluctuations in the working conditions of the various components, running out
98 of the design conditions and causing overloads, in addition to the expected
99 consequences on the dynamic behaviour.

100 Being this uneven load sharing among the different paths highly undesir-
101 able, the first of the two sources mentioned above could be easily avoided.
102 The variable meshing stiffness is an inherent characteristic of gear transmis-
103 sions. An unbalanced load share among the planetary paths can be caused
104 by the different phase of the meshing cycle between paths, and the direct way
105 to avoid this would be the synchronization of the meshing paths. When this
106 design is adopted, another problem arises: the differences between simple
107 and double contact for each mesh of each path would pile up in the complete
108 transmission error or apparent stiffness of the total planetary transmission,
109 magnifying their peak-to-peak fluctuation and becoming a great source of vi-
110 bration and noise. Because of this, the choice of synchronized planet design
111 is not usual and, on the contrary, a softened shape of the global transmission
112 error signal is desired, for which the different meshing paths are phased out

113 $2\pi/n$ (where n is the number of paths or planets).

114 With respect to the second source of uneven load sharing it is not easily
115 avoidable, as it is related to all the errors in the manufacture or assembly
116 of the various components of the planetary transmission. The present work
117 focuses on these and particularly on the errors in the planet positioning.

118 As stated, the uneven load sharing has consequences both in the dy-
119 namic behavior: vibrations due to excitation by varying forces, and durabil-
120 ity: higher workloads than nominal, or worse fatigue behavior of bearings
121 and wheels due to the increase in the stresses fluctuation amplitude. That
122 is why engineers and researchers have dedicated a great deal of attention to
123 the study of planetary gears and the design of techniques to improve the
124 load sharing. The most complicated solutions include flexible supports of
125 the planets, specifically designed to absorb manufacturing or assembly er-
126 rors. Another possible approach is to introduce flexible rings, also capable
127 of suffering deflections which accommodate and absorb the causes of uneven
128 load sharing. However, the simplest solution to achieve an improved load
129 distribution is to allow one of the central members of the transmission to
130 move freely, without any bearing restrictions on its movement around the
131 nominal position. The axisymmetric spatial configuration of the planetary
132 transmission allows to contemplate this solution, causing the central mem-
133 bers to present a theoretical null radial load, and thus allowing the central
134 elements to support themselves by effect of the combination of forces engag-
135 ing on them, without conventional bearing support.

136 In this work, the planet load sharing has been defined as the ratio of the
137 meshing torque in the sun due to each of the 3 planet-sun meshes, to the
138 total input torque T_{ext} calculated as:

$$LSR_i = \frac{T_{(Pi-S)}}{T_{ext}}; \quad i = 1, 2, 3 \quad (1)$$

139 Representing the relationship between the torque transmitted by the sun
140 to each load paths and the total external torque applied. Therefore, in a
141 three planet system as the used for these paper examples, the perfect Load
142 Sharing Ratio (LSR) would be $1/3$. In the following sections, the real value
143 of the LSR for each planet will be discussed, attending to the configuration
144 of the system, the pin point positioning error of the planets in the tangential
145 (e_{tan}) and in the radial (e_{rad}) directions, and also paying attention to the role
146 played by the magnitude of the torque to be transmitted (load level).

147 **3. Planetary transmission configurations**

148 The study of the behavior of the planetary transmission presented in this
 149 paper is focused on the load sharing. With the aim of analyzing the influence
 150 of the transmission configuration on the load sharing behavior two different
 151 configurations are used in the study: fixed or floating sun, with the rest of
 152 elements constrained to strict rotation. Thus, in the case of the fixed sun
 153 configuration, the static equilibrium is defined for each position marked by
 154 the angular positions of the ring and planet carrier. Hence, the only degrees
 155 of freedom left to determine are the angular positions of planets and sun,
 156 and so four equilibrium equations are needed:

$$\begin{aligned}
 T_{S-Pi} + T_{R-Pi} &= 0 & i = 1, 2, 3 \\
 \sum_{i=1}^3 (T_{S-Pi}) &= T_{ext}
 \end{aligned}
 \tag{2}$$

157 When considering the floating sun configuration, it is also necessary to
 158 determine the position of the sun gear, so the system of equations presented
 159 above must be extended with the balance of forces in the sun as:

$$\sum_{i=1}^3 (\mathbf{F}_{S-Pi}) = \mathbf{0}
 \tag{3}$$

160 As example, a real gear planetary reducer from agricultural machinery is
 161 modeled, whose main parameters are shown in Table 1. The real application
 162 consists of two stages with a common ring, reason for which the ring width is
 163 larger than the rest of the wheels. All the results produced in this paper have
 164 been obtained from the model published in [23], which hybridizes analytical
 165 solutions with finite element models in order to compute the contact forces,
 166 taking into account the coupling through gear body deformations.

167 *3.1. Fixed sun*

168 Besides the load sharing ratio, the Transmission Error (TE) is an impor-
 169 tant factor when studying gear transmission behaviour. Although a more
 170 detailed analysis of the simulated values obtained with the hybrid model
 171 (with analytical and FE features) used in this paper can be shown in [23],
 172 a brief highlight of the non-defected transmission error will be shown in this
 173 section, to facilitate the appreciation of the effects caused by the configura-
 174 tion and manufacturing and mounting errors. Attending to the transmission

Table 1: Modeled transmission parameters (mm)

	Sun	Planet	Ring
Teeth number	16	24	65
Module (md)	4.23	4.23	4.23
Width	25	25	85.9
Pressure angle (tool)	25°	25°	25°
Tooth thickness	6.40	8.30	-
Space width	-	-	8.25
Tip rounding radius	0.05 md	0.05 md	0.05 md
Shaft radius	20	20	156.4
Number of planets	3		
Planets angular phase	120°		
Centre's distance	86.4		
Elastic modulus	207GPa		
Poisson's ratio	0.3		

175 error when the fixed configuration is used, Figure 1 shows how the shape of
176 the resulting error corresponds to the composition of the three independent
177 sun-planet-ring meshes phased out $2\pi/3$.

178 The effect of the transmitted load on the transmission error is threefold:
179 it modifies the average value, the peak-to-peak amplitude and introduces a
180 slight change in the shape of the transmission error curve, as shown in Figure
181 2.

182 The load sharing ratio obtained when a fixed configuration is used rep-
183 resents a direct translation of the relation between the meshing stiffness of
184 each path sun-planet-ring. Thus, in the planetary transmission used in this
185 paper the load sharing among planets undergo strong variations, as shown in
186 Figure 3, reaching differences up to $\pm 9\%$ of the total load. With respect to
187 the theoretical load capacity of the planet, this difference of 9% in the total
188 load transmitted entails a $\pm 27\%$ variation. This means that, although the
189 dynamic behaviour should improve due to the smoother transmission error
190 achieved with the phased-out configuration, the real load transmitted by the
191 gears is about one third higher than the theoretical one.

192 It could be assumed that this considerable fluctuation in the load trans-
193 mitted by each path of the planetary transmission should have serious con-

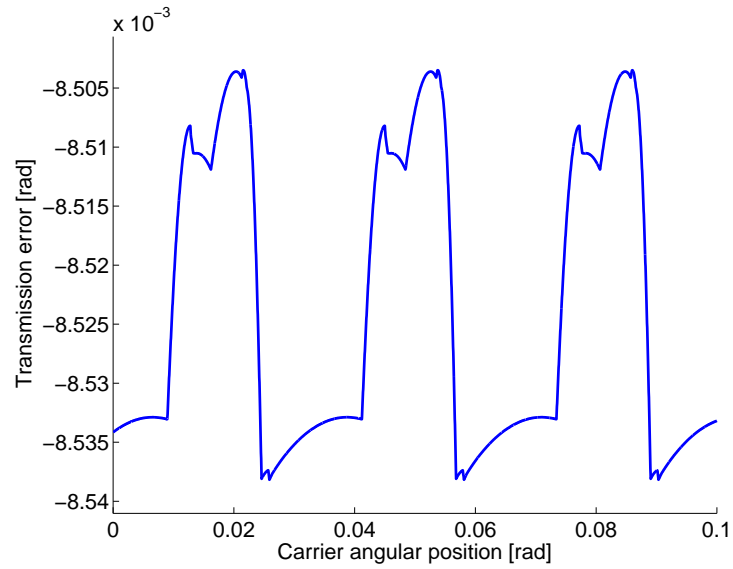


Figure 1: Transmission error with a fixed sun configuration

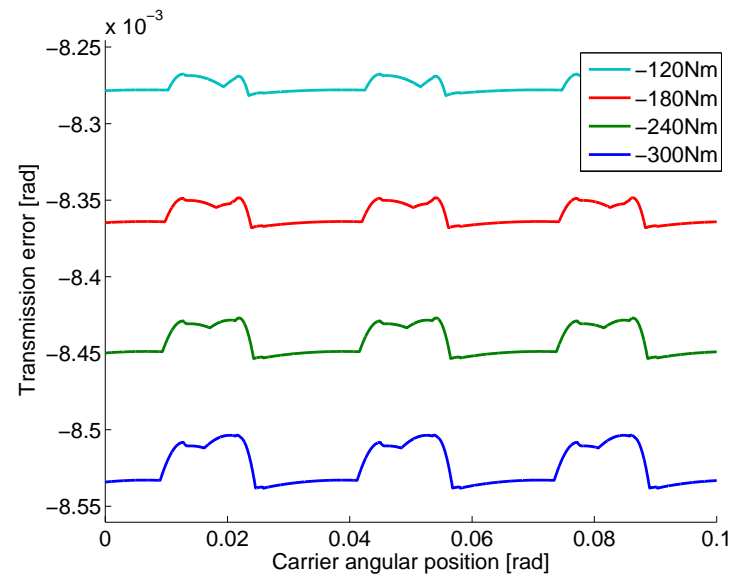


Figure 2: Effect of the load on the transmission error with a fixed sun configuration

194 sequences on the durability of the gear teeth. However, taking into account
 195 the tooth load sharing ratio throughout the meshing cycle shown in Figure
 196 4, it can be appreciated that although the transmitted torque fluctuates up

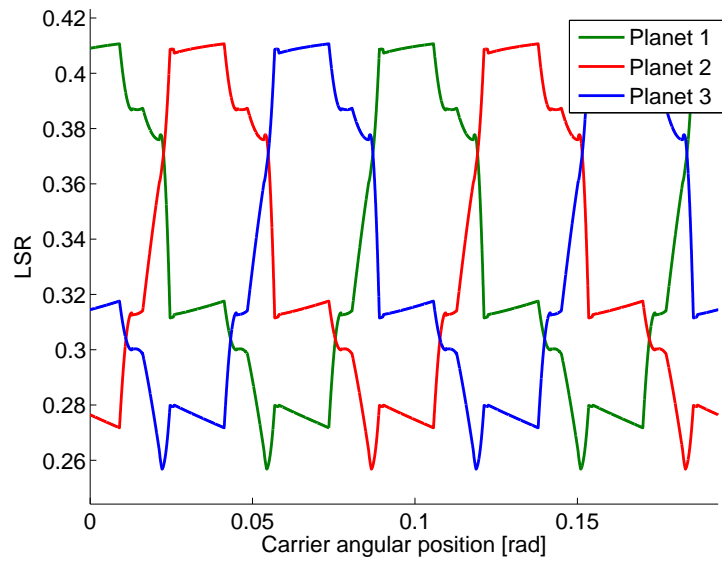


Figure 3: Load sharing ratio with a fixed sun configuration

197 to a $\pm 27\%$ variation, the force does not even reach the 100% of its nominal
 198 value.

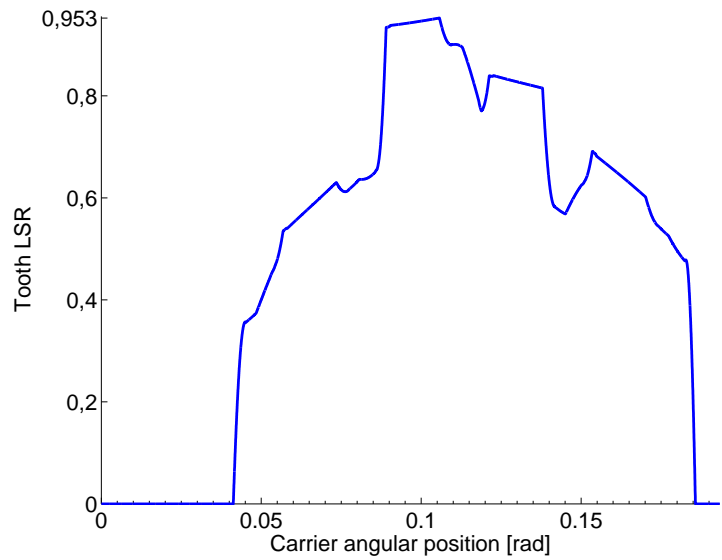


Figure 4: Tooth load sharing ratio with a fixed sun configuration

199 To understand this fact, it is necessary to explain how it is calculated the
200 tooth load sharing ratio, which is the ratio force-in-the-teeth to maximum-
201 theoretical-force in the teeth. The numerator is the value of the contact
202 force simulated in the teeth, and the denominator is the theoretical force
203 that equilibrates the transmitted torque when applied at a base-radius dis-
204 tance, divided by the number of paths. In other words, the tooth load sharing
205 ratio represent the difference between the tooth force in the real transmission
206 and the tooth force in a theoretical transmission with only one teeth in con-
207 tact for each path, transmitting in even load sharing conditions. It is then
208 easy to understand that the fluctuation of $\pm 27\%$ in the transmitted load by
209 the planet does not necessarily imply the same fluctuation in its teeth. The
210 maximum peak of transmitted load corresponds to the maximum value of the
211 path stiffness, which takes place precisely during the double contact period.
212 Thus, the $+27\%$ is shared by a teeth pair, and the corresponding tooth load
213 sharing ratio falls below the unit value. This fact can be appreciated com-
214 paring Figure 3 and Figure 4, where the tooth load sharing ratio corresponds
215 to the planet 3 planet load sharing ratio both in blue.

216 3.2. *Floating sun*

217 When the configuration used includes a floating member, as the sun in
218 this case, the load sharing situation changes drastically. The sun has to meet
219 the static equilibrium of forces in addition to the torque equilibrium, and
220 this causes the sun making an orbit around its central position, as shown in
221 Figure 5. This motion corresponds with the sun keeping away from the most
222 stiff meshes in the direction of the meshing lines, until it reaches a point
223 where the sum of products of the individual overlapping distances and their
224 corresponding stiffness values are equal for all paths.

225 In Figure 6, three graphic constructions are shown to help visualize the
226 motion of the sun when it is allowed to move freely. At the top sub-figure,
227 it is represented the forces and torque equilibrium situation for the fixed sun
228 configuration. The system of gear forces results in a balance of torques with
229 a nonzero force component caused by the greater stiffness of the planet-sun
230 mesh 1 (in red).

231 When the sun is released from all constraints, it will translate and rotate
232 to find the new position of equilibrium. For the shake of clarity, the movement
233 is decomposed in two independent parts: translation and rotation.

234 Starting from the initial graphic construction, the sun must move away
235 from the most rigid contact (mesh 1), until the forces triangle becomes closed

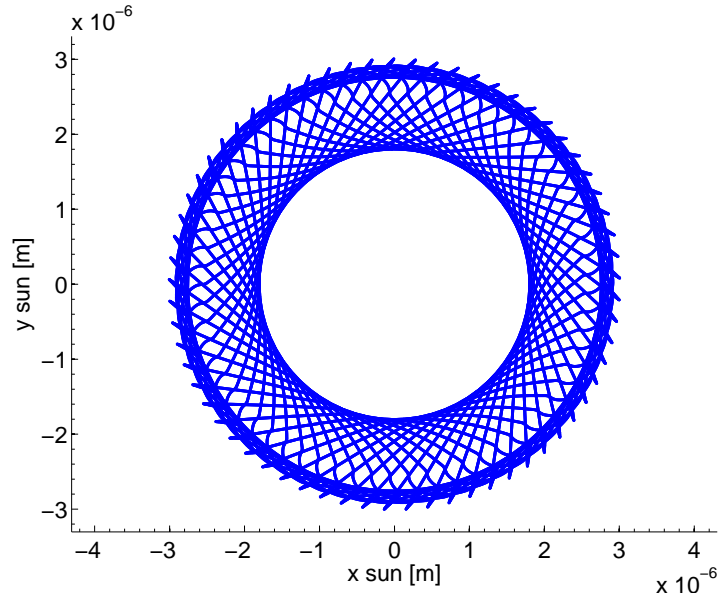


Figure 5: Sun orbit with a floating sun configuration

236 (equilateral due to the symmetry of the problem). The gear contact over-
 237 lap of mesh 1 is reduced by a length of $\Delta\delta_1$. Consequently, the other two
 238 gear contact overlaps will increase by certain lengths $\Delta\delta_2$ and $\Delta\delta_3$. These
 239 variations in the overlap distances will have proportional consequences in the
 240 values of the contact forces. There will be a position of the gear center that
 241 will comply with the equilibrium of forces, closing the triangle as shown in
 242 the second sub-figure.

243 Since the base radius remains unchanged during these operations, the
 244 torque values due to the contact forces will be consequently proportional
 245 to said contact forces. In the position of the gear center found in the first
 246 movement of the sun, the equilibrium of forces has been satisfied, but now the
 247 balance of torques has varied, and the torques produced by the gear meshes
 248 are not balanced with the external applied torque T_{ext} . As the position of the
 249 gear center guarantees the force equilibrium, to balance the external torque
 250 will suffice with the addition of a certain rotation to the gear, increasing the
 251 overlaps and thus the sum of the individual torques. These two steps should
 252 be repeated until both equilibrium conditions are satisfied: translations to
 253 maintain the forces triangle closed (equilibrium of forces) and rotation to
 254 maintain the torque equilibrium.

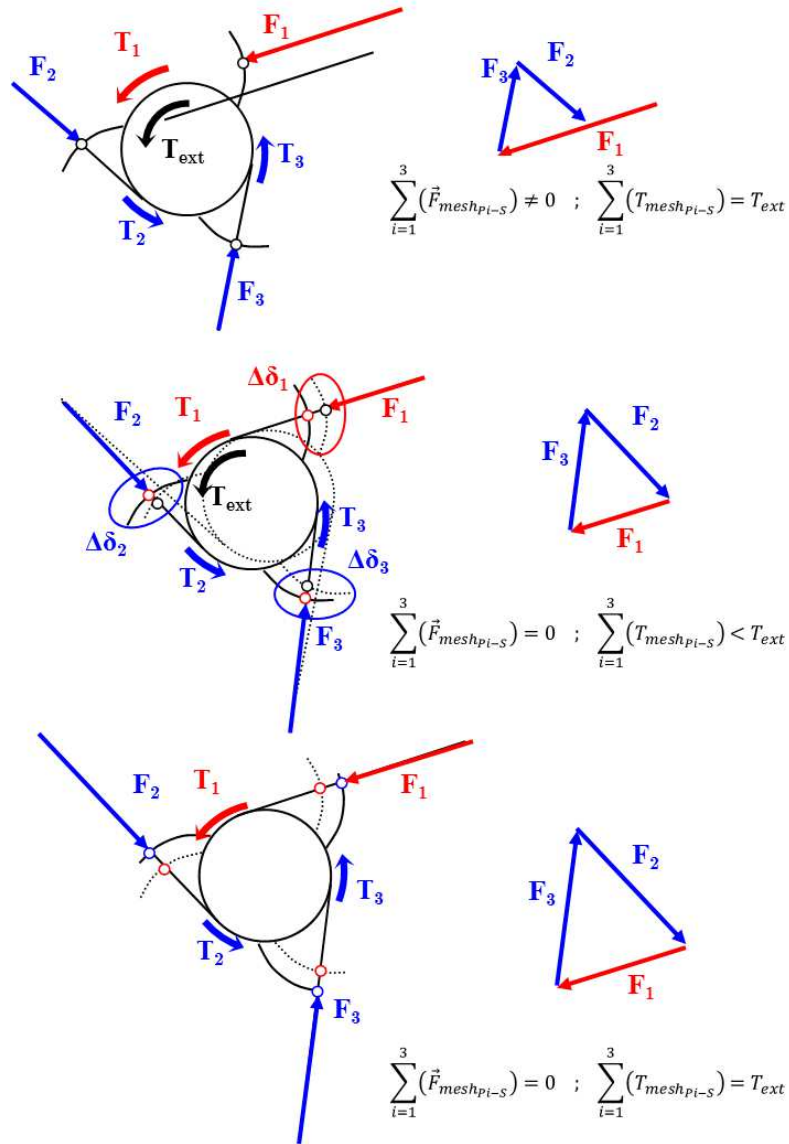


Figure 6: Translation and rotation of the sun due to variable meshing stiffness

255 Understanding this necessary composition of movements for the sun when
 256 it is released from its bearing constraints can also help to understand two
 257 important consequences in the transmission behaviour. First of all, due to
 258 the necessity of satisfy the forces equilibrium and the symmetry of the prob-
 259 lem, the load sharing must be even among planets. Each planet will carry

260 the same force vector (one side of the equilateral triangle shown in the Fig-
 261 ure 6). The second consequence is related to the transmission error signal
 262 obtained from the planetary transmission, shown in Figure 7. The gear has
 263 to perform an extra rotation with respect to the configuration with fixed
 264 center, to compensate the loss of torque due to the translation. Another
 265 interpretation of this fact, with maybe more profound implications, is that
 266 the implementation of floating central members in planetary transmissions
 267 brings, as a consequence, the reduction of the overall rotational stiffness of
 268 the system.

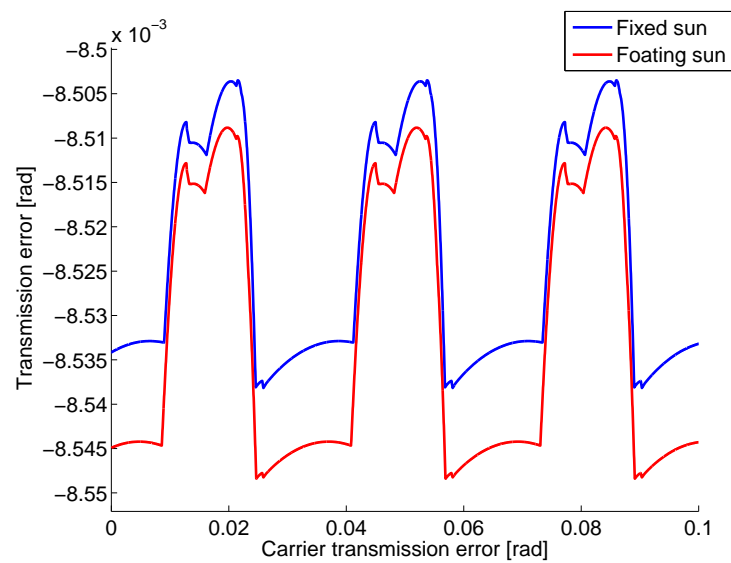


Figure 7: Transmission error with a fixed sun vs. floating sun

269 As stated, the LSR must be even for all the load paths, as shown in Figure
 270 8. There are some negligible differences in the distribution, due to the loss
 271 of symmetry when contacts out of the line of action take place (the example
 272 transmission does not present profile modifications, and consequently there
 273 are corner contacts in the teeth tips).

274 4. Errors in planetary transmissions

275 It is easy to find numerous references in the literature dealing with the
 276 behavior of planetary train in the presence of defects, especially after the
 277 recent boom in transmissions computational modeling.

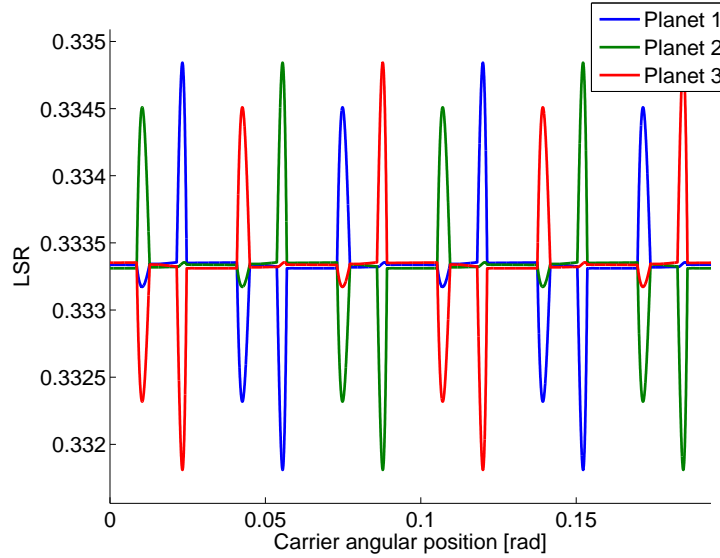


Figure 8: Load sharing ratio with a floating sun configuration

278 A large percentage of the published works are based on the study of the
 279 Load Sharing Ratio, analyzing the causes and defects for which an even dis-
 280 tribution does not occur, and proposing solutions or tools for predicting the
 281 amount of imbalance [26, 27]. It is generally accepted that transmissions
 282 with 3 planets are those with a better load sharing capability. The option
 283 of configuring the transmission with a floating central member is the solu-
 284 tion that provides the best results, absorbing deviations, manufacturing and
 285 assembly errors, and improving the LSR.

286 There are two sources related to manufacturing and assembly errors which
 287 are known to have a strong impact on the load sharing distribution in plan-
 288 etary transmissions: the errors in the positioning of the planets [26] and the
 289 eccentricity of the wheels [28]. The following describes each of the errors and
 290 their implementation on the model developed.

291 4.1. Planet pinhole positioning errors

292 The planet positioning error occurs when due to manufacturing tolerances
 293 of the carrier, the planet shaft positions differ from their theoretical location.
 294 As the impact of this error is determined by its direction, in this study the
 295 error will be divided in two components: radial or tangential component, as
 296 shown in Figure 9. For simplicity and clarity, the positioning error effect will

297 be studied in one single planet, remaining the other two planets located in
 298 their nominal position. As the considered errors are sufficiently small, the
 variation of the pressure angle will be neglected.

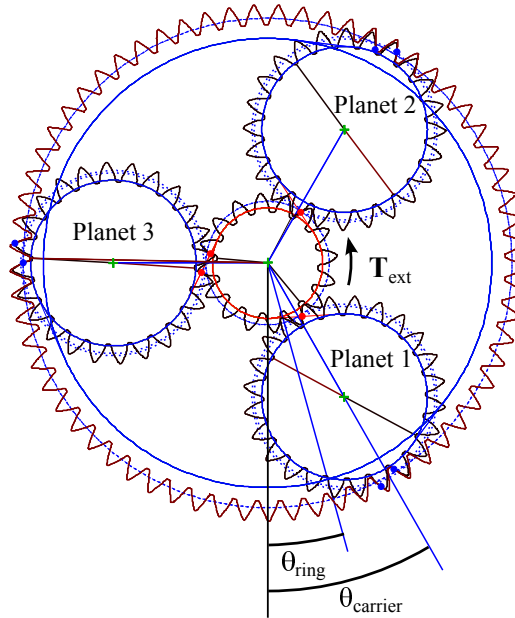


Figure 9: Planet position errors

299
 300 The positioning error of the planet is decomposed into the e_{tan} , parallel
 301 to the line tangent to the circumference at the planet center points, and
 302 the e_{rad} , which corresponds to deviations in the radial direction towards the
 303 center of the system. The sign of the value is positive when the deviations
 304 coincides with the direction shown in Figure 9.

305 The tangential component of the positioning error has the effect of ad-
 306 vancing or delaying the beginning of the contact teeth of the affected planet,
 307 depending on the direction of the error. In Figure 10 it is shown a planet
 308 with positive error e_{tan} . It can be appreciated that the movement of the
 309 planet centre with respect to its theoretical position causes the apparition
 310 of variation in the geometric overlap for the sun-planet ($\Delta\delta_{S-P}$) and ring-
 311 planet ($\Delta\delta_{P-R}$) meshes respectively. These variations in the geometric over-
 312 laps bring the consequent variation of the associated contact forces in both
 313 meshes, and thus in the torque transmitted by the corresponding sun-planet-
 314 ring path affected with the error. In this case, the positive error combined

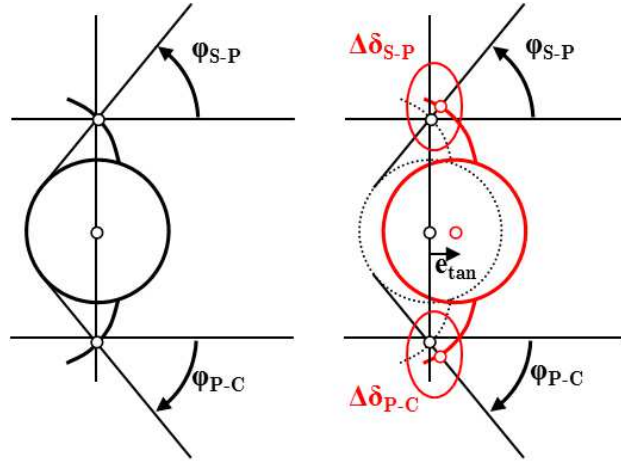


Figure 10: Overlapping variation due to e_{tan}

315 with the direction of the torque, results in a preload of the affected planet.
 316 This excess of transmitted torque ΔT is given by:

$$\frac{(k_{P-R}\Delta\delta_{P-R} + k_{S-P}\Delta\delta_{S-P})}{2} = \Delta T \quad (4)$$

$$\text{where } \begin{cases} \Delta\delta_{S-P} = e_{tan} \cos(\varphi_{S-P}) \\ \Delta\delta_{P-R} = e_{tan} \cos(\varphi_{P-R}) \end{cases}$$

317 Where it can be appreciated the influence of the contact stiffness k_{S-P}
 318 and k_{P-R} and the pressure angles φ_{S-P} and φ_{P-R} .

319 The amount of excess transmitted torque is much lesser for radial position-
 320 ing errors depicted in Figure 11.

321 In this situation, the displacement of the planet produces a variation in
 322 the geometric overlapping $\Delta\delta$ in the opposite direction for each mesh. Thus,
 323 while the overlap variation is positive in the case of the sun-planet mesh, it
 324 is negative for the planet-ring mesh. As the torque balance in the planet
 325 must be zero, it seems apparent that the planet will suffer a rotation moving
 326 away from the sun-planet contact and closing the gap opened in the planet-
 327 ring contact. If the geometric overlaps are the same in absolute value, they
 328 cancel each other and the only effect of the planet radial displacement is an
 329 additional rotation of the gear with consequences in the phase of the system,
 330 but no effect on the transmitted load. Nevertheless, the fact that in some
 331 planetary drives the pressure angle of the sun-planet pair is different than
 332 the planet-ring has been ignored by some authors. It is common to find in

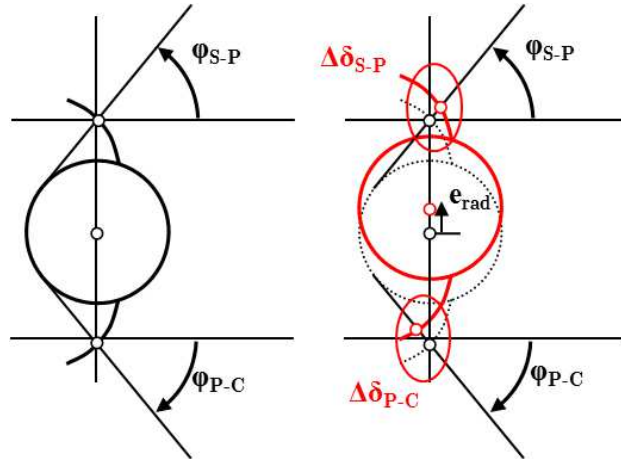


Figure 11: Overlapping variation due to e_{rad}

333 the literature assertions about the negligible effect of the radial component
 334 of the positioning error on the load distribution through the different paths
 335 in planetary transmissions [26].

336 Whilst it is true that the compensatory nature of the overlaps produced
 337 by the radial movement of the planet minimizes its effect on the load sharing
 338 ratio (especially when compared to the case of the tangential displacement),
 339 the LSR is far from perfect when there are differences in the pressure angles,
 340 as in this case:

$$\begin{aligned}\Delta\delta_{S-P} &= e_{tan} \sin(\varphi_{S-P}) \\ \Delta\delta_{P-R} &= e_{tan} \sin(\varphi_{P-R})\end{aligned}\tag{5}$$

341 4.2. Run-out errors

342 Eccentricity (run out) occurs when there is a difference between the geo-
 343 metric centre of the gear and the position of its rotation centre. Analyzing
 344 its nature, from a kinematic point of view, the run-out error is a kind of
 345 positioning error in which the radial and tangential components vary sinu-
 346 soidally with the angular position of the planet gear (or carrier angular po-
 347 sition). Thus, the implementation of eccentricity is based on the positioning
 348 error implementation, with the added complexity of the sinusoidal variation
 349 function of the planet gear centre angular position.

350 5. Modeled behavior of the planetary transmission with errors

351 As stated before, the fewer degrees of freedom, the greater impact of er-
352 rors in a transmission has. As compliances are introduced, the load sharing
353 behavior improves, as it has been demonstrated by means of theoretical, ge-
354 ometric and modelling approaches in the previous sections. In the following,
355 positioning errors will be analyzed by components in order to ratify the con-
356 clusions drawn from geometric analysis presented before. To illustrate the
357 effect of the configuration on the LSR, the results obtained for a fixed sun
358 will be shown first, and afterwards the improvement of the behavior when
359 the sun is allowed to orbit will be presented.

360 5.1. Tangential component of the positioning error in fixed sun configuration

361 In Figure 12, it is shown the LSR for a planetary transmission with fixed
362 sun and an e_{tan} of varying magnitude in planet 1. In a dashed line, at the
363 center of the figure, it is reproduced the LSR for a transmission without
364 defects (same results as those shown in Figure 4) to be used as reference. In
365 a dotted line, the results of LSR for a tangential positioning error of $30 \mu m$
366 are shown. At this level of error the planet 1 assumes at certain positions
367 of the meshing cycle up to 80% of the total transmitted load (900Nm at the
368 sun).

369 The loss of symmetry in the results is generalized. The non-faulty plan-
370 ets present of course a very different level of load, but also the shape and
371 amplitude of the LSR curve is affected, now having three different forms of
372 LSR for each planet. When the positioning error is increased to $50 \mu m$, the
373 overlap variation of the planet 1 is such to cause the complete unloading of
374 the non-faulty path loads, and the defected planet carries the entire amount
375 of transmitted torque.

376 The effect of a tangential error in the opposite direction (negative error)
377 is contrary to the previous results shown. Instead of increasing the overlap
378 distance between profiles in the planet, this error spread the profiles apart,
379 avoiding contact and thus the transmission of load. This fact can be con-
380 firmed by the LSR results shown in Figure 13, where there can be appreciated
381 a drop of the load carried by planet 1 when a tangential negative error of 20
382 μm is introduced.

383 Another conclusion can be drawn from the different effect that positive
384 or negative tangential error have on the LSR behaviour. When a planet is
385 affected by a tangential error, in an extreme case of a negative error, the

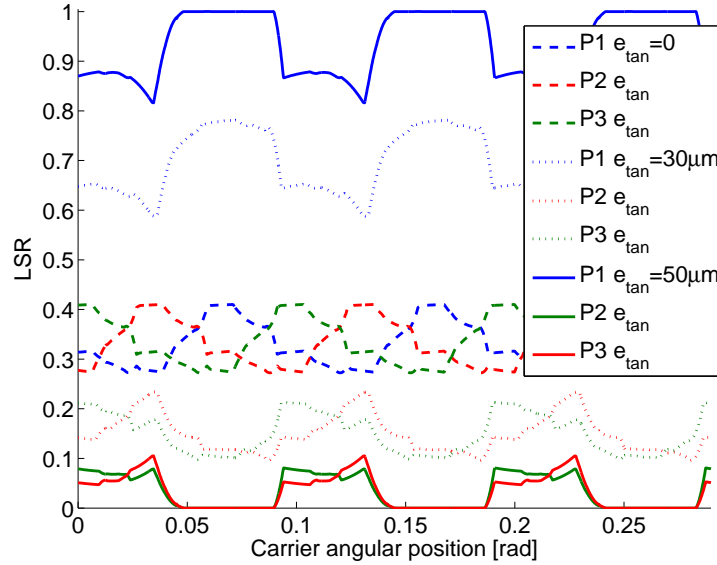


Figure 12: Load sharing ratio with fixed sun and positive e_{tan}

386 faulty planet becomes unloaded and the rest of the paths shares the excess of
 387 torque. Nevertheless, for the extreme case of positive error, the faulty planet
 388 assumes the whole amount of load. It is apparent that it is far more critical
 389 a situation in which a single planet is loaded with all the transmitted torque
 390 (positive error) than the situation in which a set of planets must share the
 391 load corresponding to the faulty unloaded planet (negative error).

392 To summarize the effect of the tangential error in the LSR, in Figure 14
 393 the LSR of planet 1 is shown for a range of e_{tan} , from negative to positive
 394 values. While the shape of the curves along the cycle varies strongly with
 395 the error magnitude, the average level of the LSR varies proportionally with
 396 the positioning error.

397 As it was first postulated in section 3.1, it could be assumed that these
 398 considerable fluctuations in the load transmitted by each planet of the trans-
 399 mission should have proportional impact on the teeth forces, with conse-
 400 quences on the gears durability. Looking again at the tooth load sharing
 401 ratio of planet 1 throughout the meshing cycle shown in Figure 15 for an e_{tan}
 402 of $30 \mu m$. Now, attending to Figure 12, planet 1 carries roughly 60% to 80%
 403 of the total transmitted load. Since the nominal load assigned to a planet
 404 in a 3 planet transmission is $1/3$, this means that planet 1 is transmitting
 405 between 1.8 and 2.4 times the assigned value of the load. Nevertheless, in

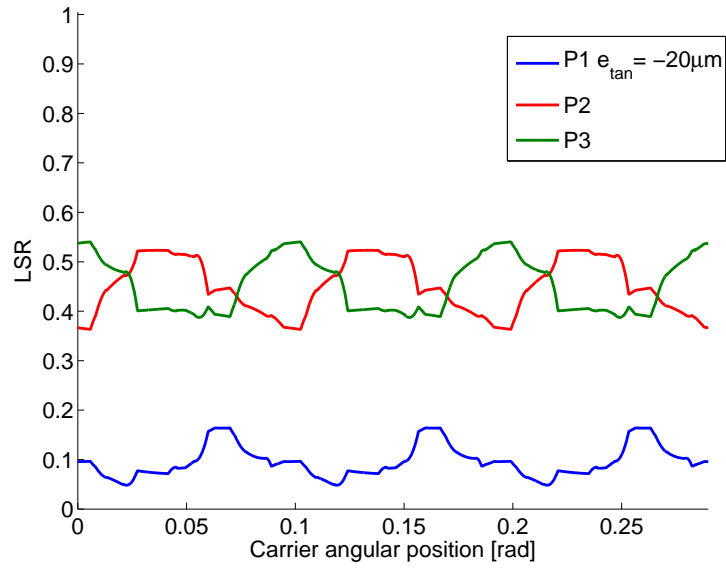


Figure 13: Load sharing ratio with fixed sun and negative e_{tan}

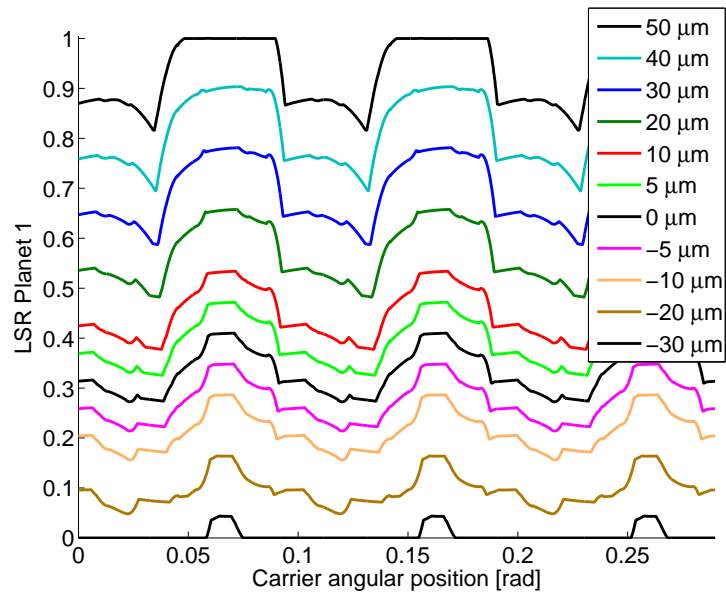


Figure 14: Load sharing ratio with fixed sun and positive and negative e_{tan}

406 the figure, the tooth load barely reaches 2 times the nominal assigned force
 407 value. The explanation resides again in the lower stiffness that a one tooth

408 contact presents, allowing the rest of the paths with more teeth pairs in mesh
 409 to carry corresponding greater forces.

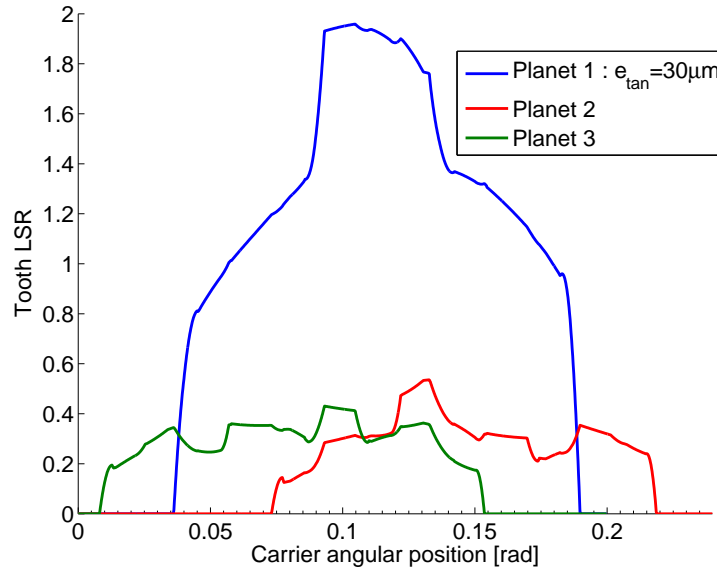


Figure 15: Tooth load sharing ratio with a fixed sun configuration and $e_{tan} = 30\mu m$

410 The effect of the torque level on the LSR is very relevant, as shown in
 411 Figure 16. In general, geometry related errors in gear transmissions have a
 412 higher visibility for lower load levels, as higher loads imply greater elastic
 413 deformations and therefore the dilution of the error deflections.

414 Although the focus of this study is centered on the LSR consequences of
 415 manufacturing errors, it is evident that there are other measurable signals
 416 that show disturbances due to these errors. The transmission error presented
 417 in Figure 2 has three lobes per planetary meshing period. In Figure 17 the
 418 same planetary transmission error is shown (for an input torque of 900Nm),
 419 when different and progressively greater values of e_{tan} are introduced in planet
 420 1. As the faulty planet gradually assumes the torque corresponding to the
 421 other two loading paths (planets 2 and 3), the transmission error lobes caused
 422 by these unloading planets become reduced and absorbed by the shape of a
 423 transmission error characteristic of a sun-planet-ring transmission.

424 5.2. Radial component of the positioning error in fixed sun configuration

425 The radial component positioning error of the planets, as introduced in
 426 Section 4, has a much lesser effect on the LSR than the tangential compo-

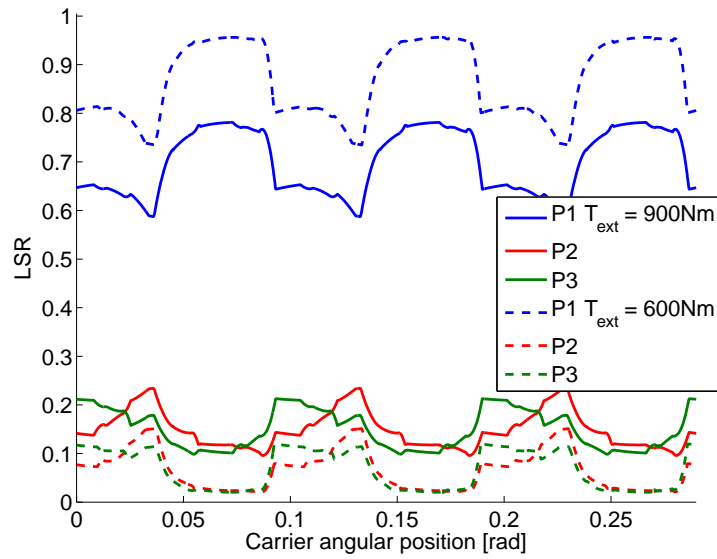


Figure 16: Load sharing ratio with a fixed sun configuration and $e_{tan} = 30\mu m$ and different loads levels

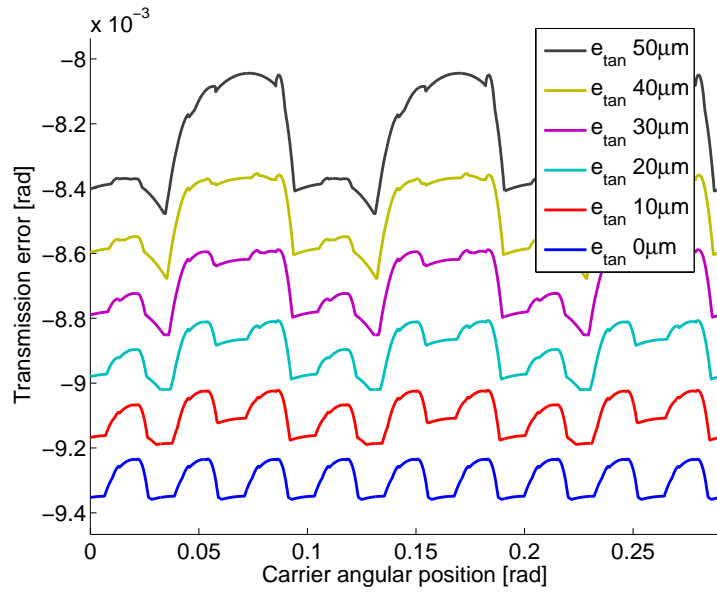


Figure 17: Transmission error with a fixed sun configuration under different e_{tan} values

427 nent. It mainly depends on the difference between the operating pressure

428 angles of the sun-planet and planet-ring meshes, and its effect is practically
 429 nonexistent when these angles are equal (which is usual). In the example
 430 transmission, the operating pressure angles for each gear pair are calculated
 431 for the planetary transmission as $\varphi_{P-R} = 24.45^\circ$, $\varphi_{S-P} = 27.37^\circ$.

432 Figure 18 shows the LSR of the transmission when a radial positioning
 433 error of $40 \mu m$ is introduced in planet 1. The difference between the average
 434 LSR value of the faulty path track relative to the other two is about 2 per-
 435 centage points, which shows the much lesser impact of the radial component
 436 compared to the results shown in Figure 12.

437 In order to illustrate the comparison between both component effects of
 438 the positioning error, it has been found that an error of $40 \mu m$ in the radial
 439 direction is equivalent to a tangential error of $1 \mu m$ in terms of LSR impact.
 440 Both cases are shown in an overlapping manner in the Figure to assess its
 441 coincidence. The LSR values corresponding to the tangential error are shown
 442 in black and different types of dashed line.

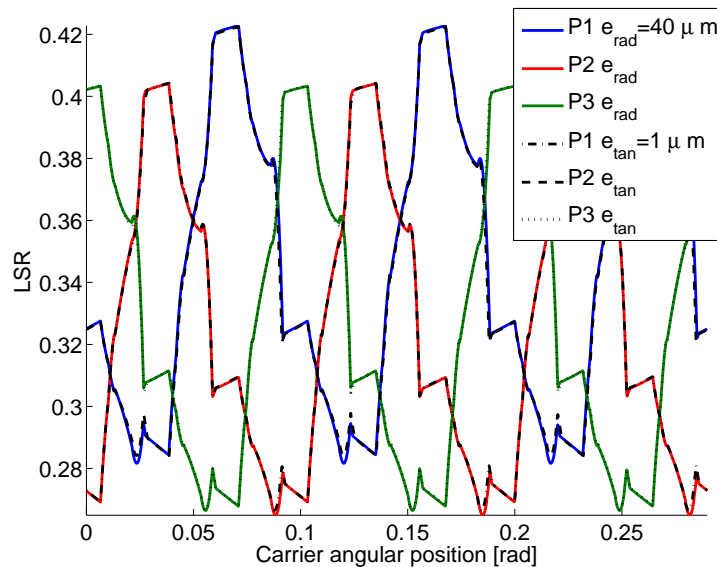


Figure 18: Load sharing ratio comparison with fixed sun between e_{rad} and e_{tan}

443 5.3. Positioning errors in floating sun configuration

444 When the configuration of a planetary transmission with positioning errors
 445 is modified from fixed to floating sun, the LSR becomes homogenized
 446 approaching a perfect $1/3$ for each load path (or planet). The sun describes

447 an orbit around its central position, compensating for the positioning error
 448 of the planet by adjusting its separation distance (and thus overlap distance)
 449 in the direction of the line of action, as shown in Figure 19.

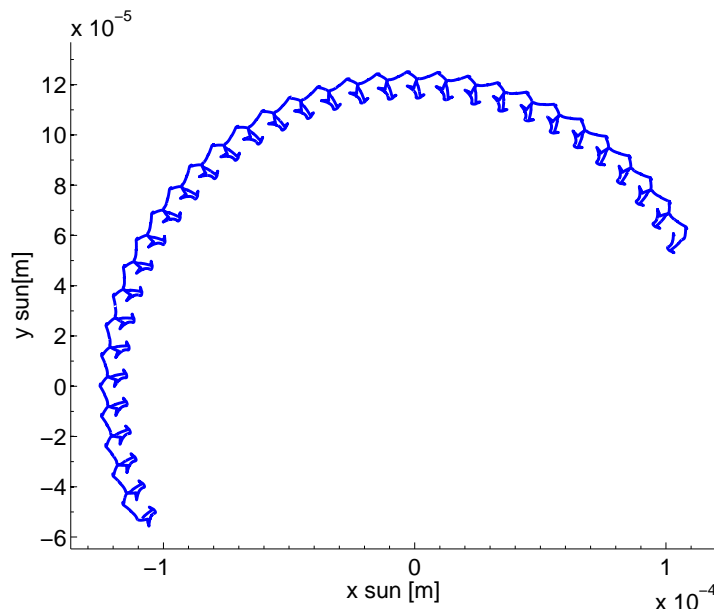


Figure 19: Sun orbit with $e_{tan} = 30\mu m$

450 In this figure, the sun orbit is shown for positions of the carrier from 0 to
 451 π radians, and considering a positive tangential positioning error of planet 1
 452 of $100\mu m$. For the initial position shown in the figure, it can be seen how the
 453 sun moves away from the center in the same direction as the pressure angle,
 454 therefore moving away from the contact with planet 1. It can be appreciated
 455 from the orbit graph that the radius of the orbit described approximates the
 456 magnitude of the planet positioning error.

457 As it was stated before, by describing this orbit, the planetary forms a
 458 zero forces system, for which the resulting contact forces must be equal due to
 459 the symmetry, arriving as consequence to the final 1/3 LSR shown in Figure
 460 20. The only exceptions for this perfect 1/3 can be found at the contacts
 461 out of the line of action at the teeth tip (the peaks in the figure), where the
 462 symmetry is broken because of the change in the effective pressure angles.
 463 The net differences between the LSR of each planet, which can be appreciated
 464 in Figure 20, are also due to a slight and negligible loss of symmetry caused
 465 by the $100\mu m$ of variation in the vertex of the triangle formed by the three

466 planet pin points.

467 Complementary to the homogenization that occurs in the LSR when the
 468 sun is allowed to move, the individual tooth load sharing ratio also regains
 469 symmetry for the three planets, as seen in Figure 21. In this case, the
 470 maximum value is the unit corresponding to the force that balances a third
 471 outer torque applied during simple contact area, as the floating sun now can
 472 also compensate for stiffness differences, which did not happened in the fixed
 473 sun case shown in Figure 4.

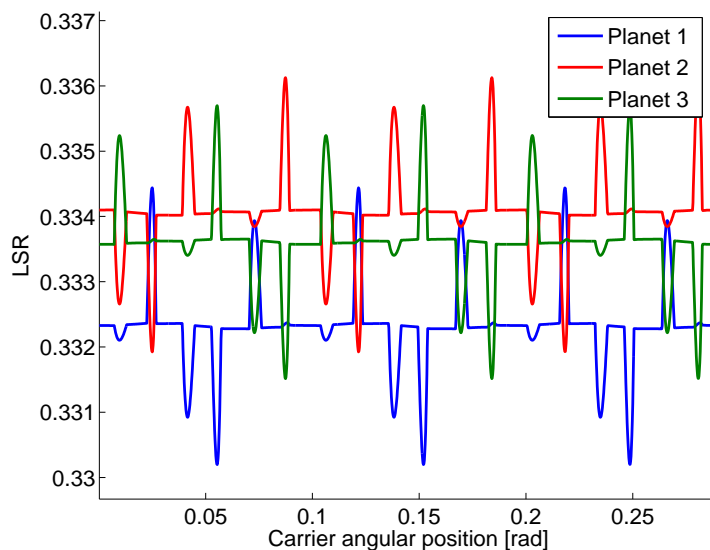


Figure 20: Load sharing ratio with floating sun and $e_{tan} = 100\mu m$

474 Regarding the transmission error, the differences between both configu-
 475 rations can be appreciated in Figure 22. As happened in the results without
 476 positioning errors shown in Figure 7, the sun must perform an extra rota-
 477 tion to compensate the loss of torque due to orbit translation, for which
 478 the transmission error increases. For the floating configuration, the shape
 479 of the transmission error exhibits symmetrical behaviour for each path, with
 480 three lobes per meshing period, whereas for the fixed sun configuration the
 481 predominant faulty planet dominates the transmission error shape.

482 5.4. Run-out errors in fixed sun configuration

483 All eccentricity simulations presented below are carried out with an error
 484 of $20\mu m$ in the positioning of the axis of rotation of planet 1 with respect

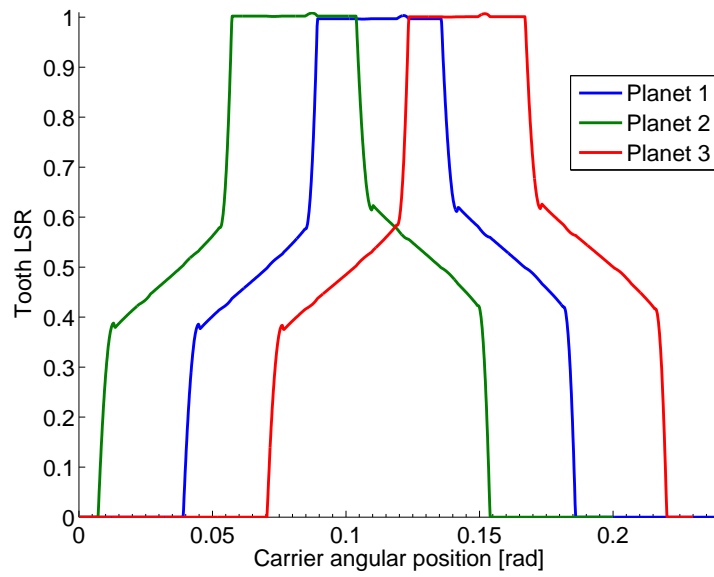


Figure 21: Tooth load sharing ratio with floating sun and $e_{tan} = 100\mu m$

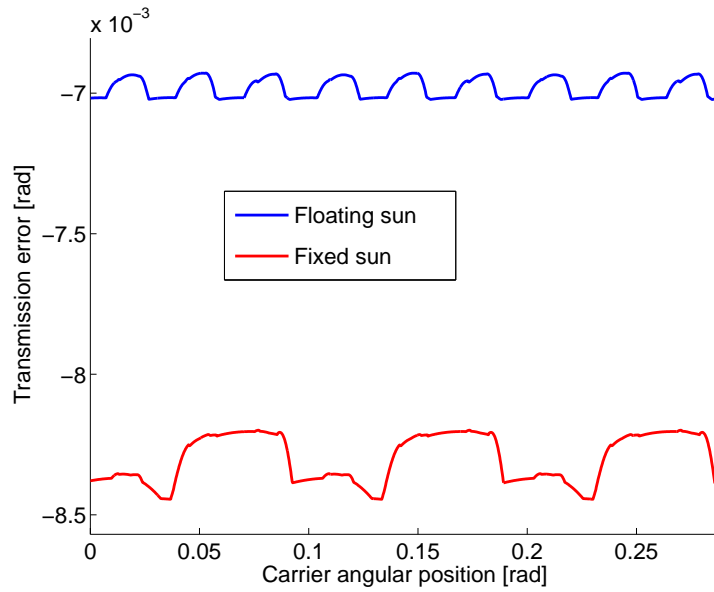


Figure 22: Transmission error with a fixed sun vs. floating sun and $e_{tan} = 30\mu m$

485 to its geometrical center. For the initial position (zero angular position of
 486 the planet carrier), the geometric center of the wheel is displaced $20\mu m$ in

487 the negative radial direction and thus, in this initial position, the planetary
488 configuration coincides to that of a negative e_{rad} of $20 \mu m$. Therefore, the
489 LSR behaviour with eccentricity has to be identical as that presented for the
490 transmission with radial positioning error of the same magnitude. Indeed,
491 in Figure 23 it is shown that the LSR curves present very little deviation
492 from the 1/3 load sharing that would be expected with an error of radial
493 positioning of such relative low magnitude. As the planet carrier rotates
494 counterclockwise, the planet does clockwise, and the projection of the ec-
495 centricity of each of its components begins acquiring a negative tangential
496 component, which progressively causes the unload of the planet as seen in
497 the previous sections. As the system rotates, the tangential component in-
498 creases and the radial decreases, up to the point where the planet performs
499 a relative rotation of $\frac{\pi}{2}$ with respect to the carrier. At this point, the eccen-
500 tricity will place the geometric center of the planet at $20 \mu m$ tangentially
501 negative, reaching the maximum amount of unload for the faulty planet. In
502 the following equation, it is given the absolute rotation value of the carrier
503 between positions of pure radial and tangential components ($\frac{\pi}{2}$ or relative
504 rotation of the planet with respect to the carrier).

$$\theta_{P/carrier} = \theta_P - \theta_{carrier} = -\frac{Z_R - Z_P}{Z_P} \theta_{carrier} - \theta_{carrier} = \frac{\pi}{2} \rightarrow \theta_{carrier} \quad (6)$$

$$\theta_{carrier} = 0.58rad$$

505 In Figure 23, it can be seen that for the pure negative tangential com-
506 ponent and the level of torque used in this simulation (600Nm), planet 1
507 becomes load-free and it can also be observed that the LSR curve shape
508 for each of the planets corresponds to that shown in previous sections for
509 negative e_{tan} . From this position, and as the system continues to rotate,
510 the eccentricity error progressively loses tangential component, reaching the
511 maximum positive value of $20 \mu m e_{rad}$ at carrier position of 1.16 rad. In
512 this position, the LSR also presents a value close to 1/3 due to the lesser
513 effect of the radial positioning error. When the system further rotates, the
514 eccentricity acquires again positive values of e_{tan} component, showing one
515 more time a LSR behaviour similar to that of a positive tangential pin point
516 positioning error of the planet at 1.74 rad. The sinusoidal behaviour of the
517 LSR for the run-out error will repeat itself periodically as the carrier rotates
518 2.32 rad, alternating LSR curve shapes characteristic of e_{tan} and e_{rad} .

519 Regarding the transmission error for the planetary transmission with run-

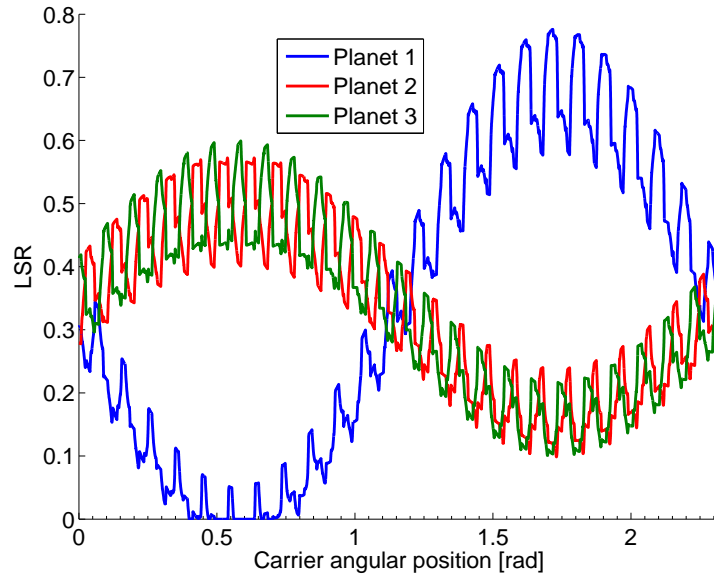


Figure 23: Load sharing ratio with $20\mu m$ eccentricity error and 600 Nm with a fixed sun configuration

520 out planet errors, a parallel analysis to that reported for the LSR can be
 521 performed. In Figure 24 the same four areas can be identified as the ec-
 522 centricity changes its dominant component from negative radial to negative
 523 tangential to positive radial and positive tangential. Attending to this rela-
 524 tively easily measurable signal, an eccentricity error can thus be identified in
 525 a quasi-static analysis.

526 5.5. Run-out errors in floating sun configuration

527 Similarly to what happened for transmission with positioning error, al-
 528 lowing the sun to move freely causes the absorption of the eccentricity effect
 529 on the LSR, as shown in Figure 26. Again the variations from the perfect
 530 distribution is caused by the loss of symmetry explained before. The orbit
 531 described by the sun is presented in Figure 25. The distance to the cen-
 532 ter depends mainly on the tangential component of the eccentricity for each
 533 position, causing the appearance of lobes in the orbit shape.

534 6. Conclusions

535 Two different planetary configurations have been studied, considering the
 536 possibility of both fixed and free translation of the sun and taking into ac-

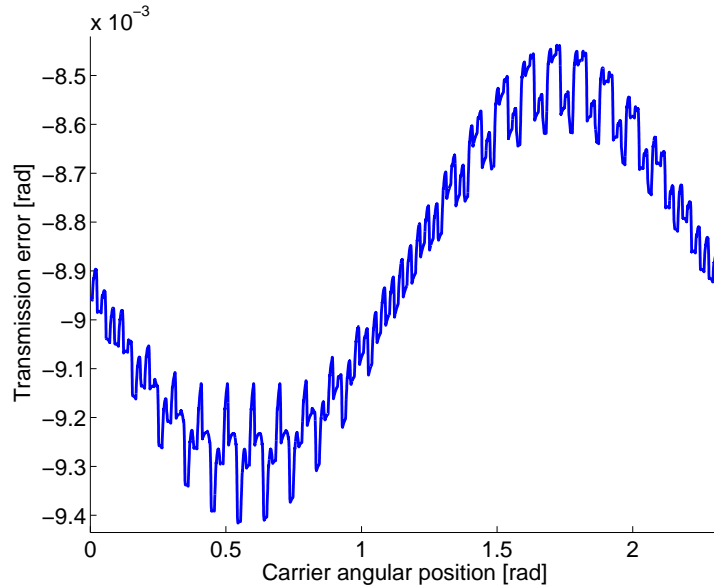


Figure 24: Transmission error with $20\mu\text{m}$ eccentricity error with a fixed sun configuration

537 count the resulting values of load sharing ratio (LSR) between paths. The
 538 results obtained in this paper are valid for any three planet transmission
 539 with a $2\pi/3$ distribution of the meshing phase, and similar values of operat-
 540 ing pressure angles as the ones shown here. The tendency of the load sharing
 541 ratio when errors are included can be extrapolated for n-planet transmissions
 542 when the configuration is fixed (of course taking into account the proportion-
 543 ality between the number of planets and the ideal LSR and its deviations),
 544 but there are critical differences in the behaviour of 3-planet versus n-planet
 545 systems when there is a floating central member, as explained in [4]. As
 546 expected, the LSR varies widely in the fixed transmission scenario, due to
 547 the different path stiffness. These variations on the values of the LSR have
 548 been found to have no impact on the maximum level of contact forces, which
 549 will not exceed its maximum nominal value in transmissions without defects.
 550 In the floating sun scenario, the LSR becomes almost perfect, except for the
 551 areas where contacts out of the line of action take place. Nevertheless, even
 552 in those zones, the variations of the LSR are insignificant, never above a
 553 tenth of percentage point.

554 Planet positioning errors have been analysed: tangential and radial com-
 555 ponents in a separately way. The tangential error has a great impact on the

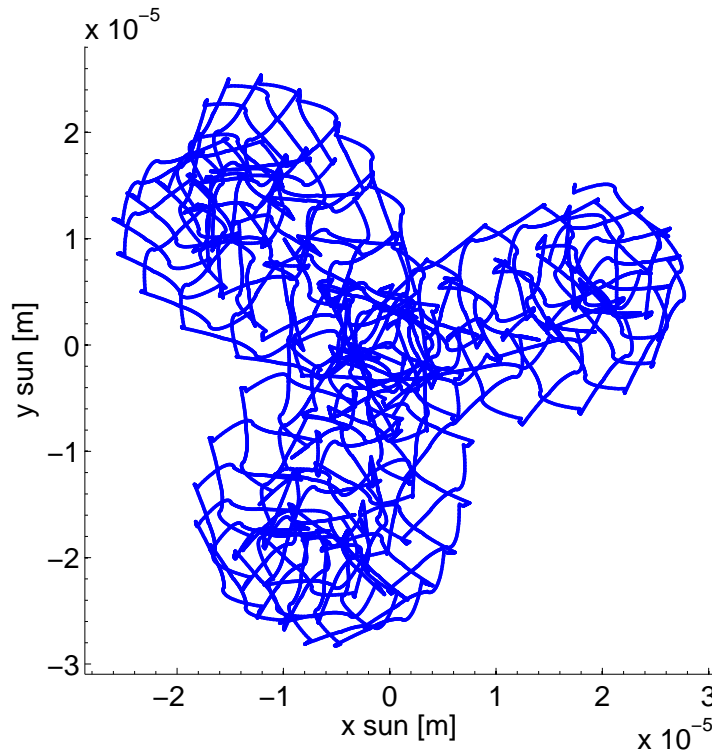


Figure 25: Sun orbit with a floating configuration and $20\mu\text{m}$ eccentricity error

556 LSR in non floating configurations. It can even cause the complete discharge
 557 of one or more paths under certain conditions of deviation and load. Whereas
 558 a negative tangential deviation (with respect to the line of action) can cause
 559 the unload of the defected planet, a positive deviation has the potential to
 560 unload all the non-defected paths. Thus, the consequences of the positive
 561 deviations are much more serious than the ones posed by the negative ones.

562 The effect of the tangential component on the LSR has an impact in the
 563 maximum contact forces level. Nevertheless, the relationship between the
 564 increase on the contact forces and the variation of the LSR has been found
 565 to be nonlinear.

566 The radial component of the positioning error has a significantly lower
 567 impact on the LSR (found to be approximately 40 times lower for the example
 568 transmission) than the tangential component. Nonetheless, this impact is
 569 not null, in spite of what can be found in some bibliographic references. It
 570 strongly depends on the design and mounting conditions of the planetary

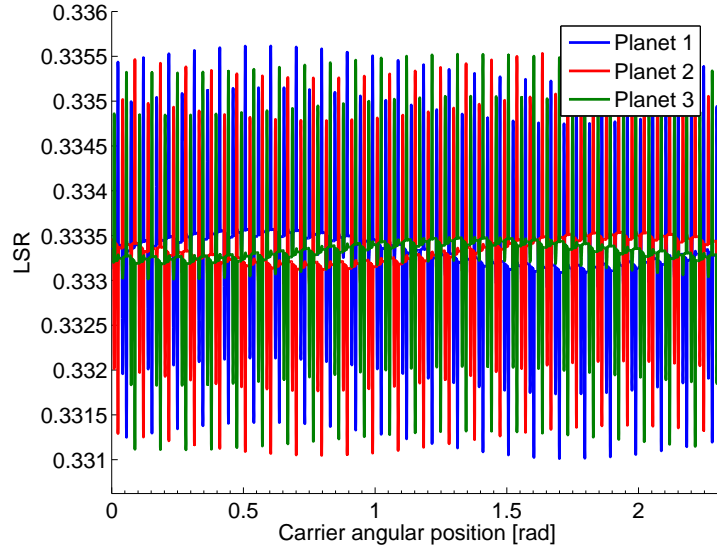


Figure 26: Load sharing ratio with $20\mu\text{m}$ eccentricity error and 600 Nm with a floating sun configuration

571 transmission; more specifically on the difference between the pressure angles
 572 of sun-planet and planet-ring.

573 When errors are present, the variations on the values of the LSR have
 574 been found to have a much lesser impact on the maximum level of teeth
 575 contact forces. Thus, the initial concern that uneven LSR could have great
 576 importance on the teeth durability due to higher loads than nominal is par-
 577 tially unfounded.

578 The increment of the transmitted load results in the dilution of the posi-
 579 tioning error effects, due to the presence of greater levels of global deflections.
 580 For the floating configuration, the sun follows an orbit of radius similar to
 581 the positioning error magnitude (for tangential components). In this way,
 582 the sun displacement absorbs the error and virtually evens the LSR.

583 The effects of the eccentricity error have also been studied. It has been
 584 concluded that they are equivalent with a sinusoidal evolution of the posi-
 585 tioning error, through its different components and directions.

586 **Acknowledgments** The authors would like to acknowledge Project DPI2013-
 587 44860 funded by the Spanish Ministry of Science and Technology for support-
 588 ing this research.

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